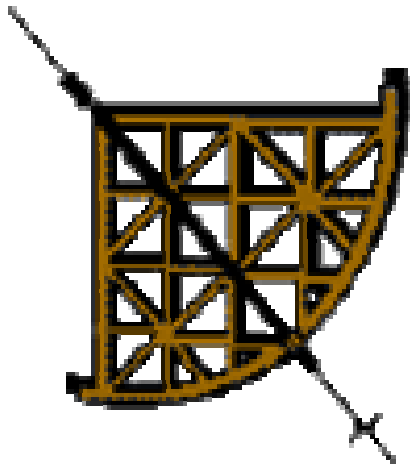


Dal *Sidereus Nuncius* alla cosmologia fisica: la fecondità del genio galileiano



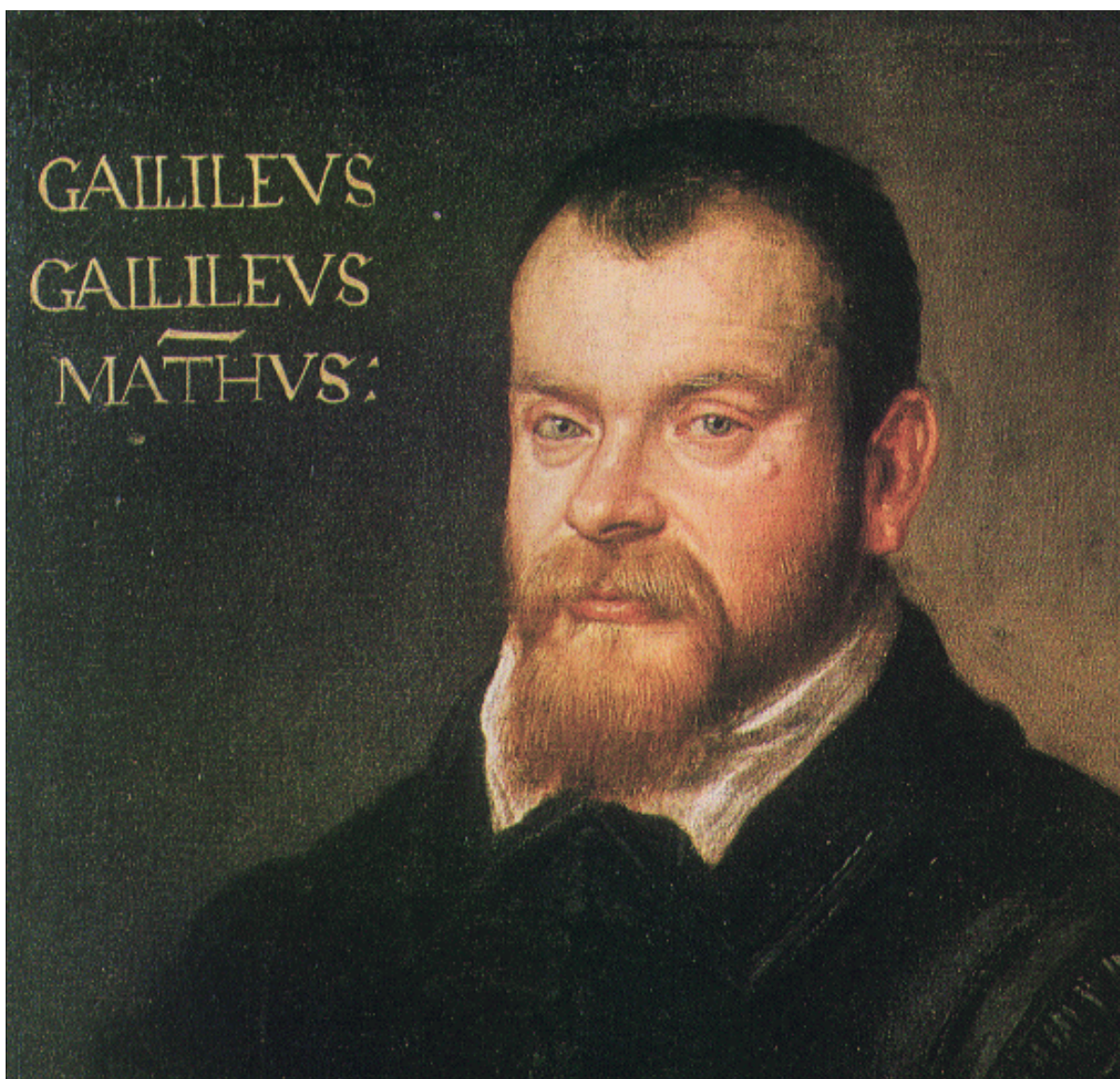
Giorgio G.C. Palumbo
Università degli Studi di Bologna
Dipartimento di Astronomia



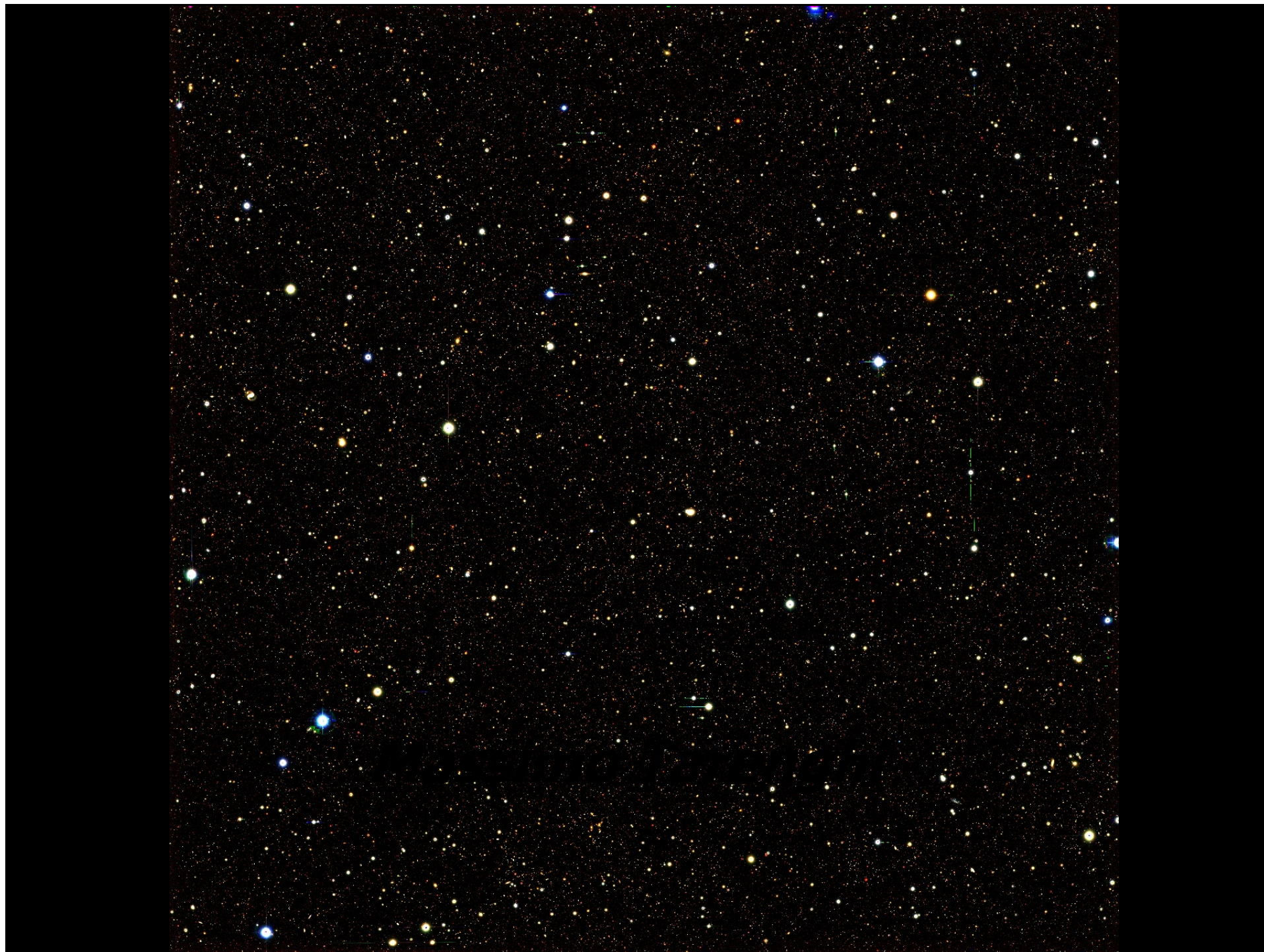


In un documento datato 25 Settembre 1608 del governo di Zeeland ai propri delegati al governo generale delle Netherlands li esorta ad aiutare il consegnatario " che dice di avere un certo aggeggio per mezzo del quale tutte le cose a grandi distanze appaiono come se fossero vicine guardando attraverso vetri che egli afferma essere una nuova invenzione"

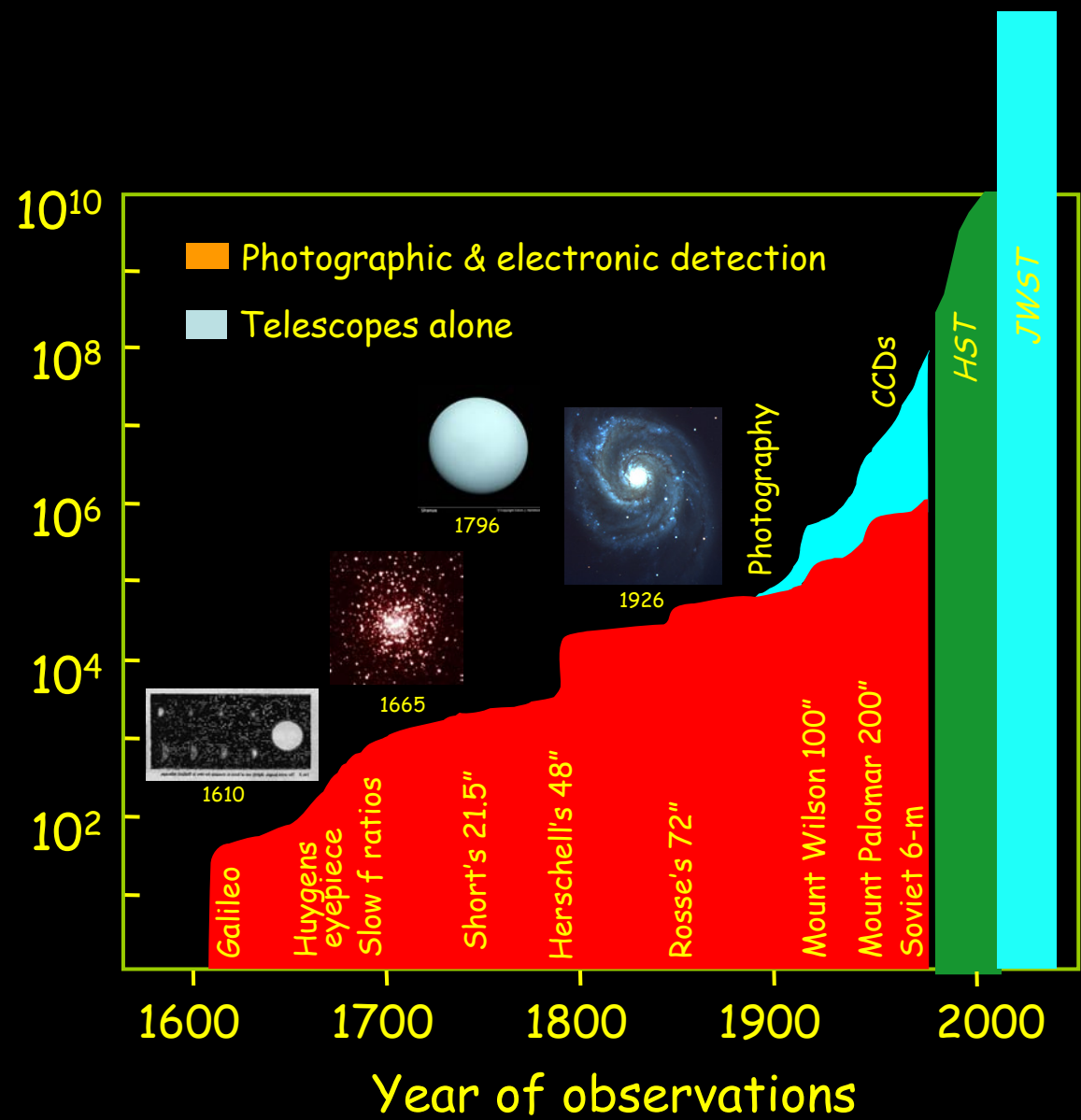
Hans Lippersheim (Lippershey) 1570-1619, Wesel, Middelburg



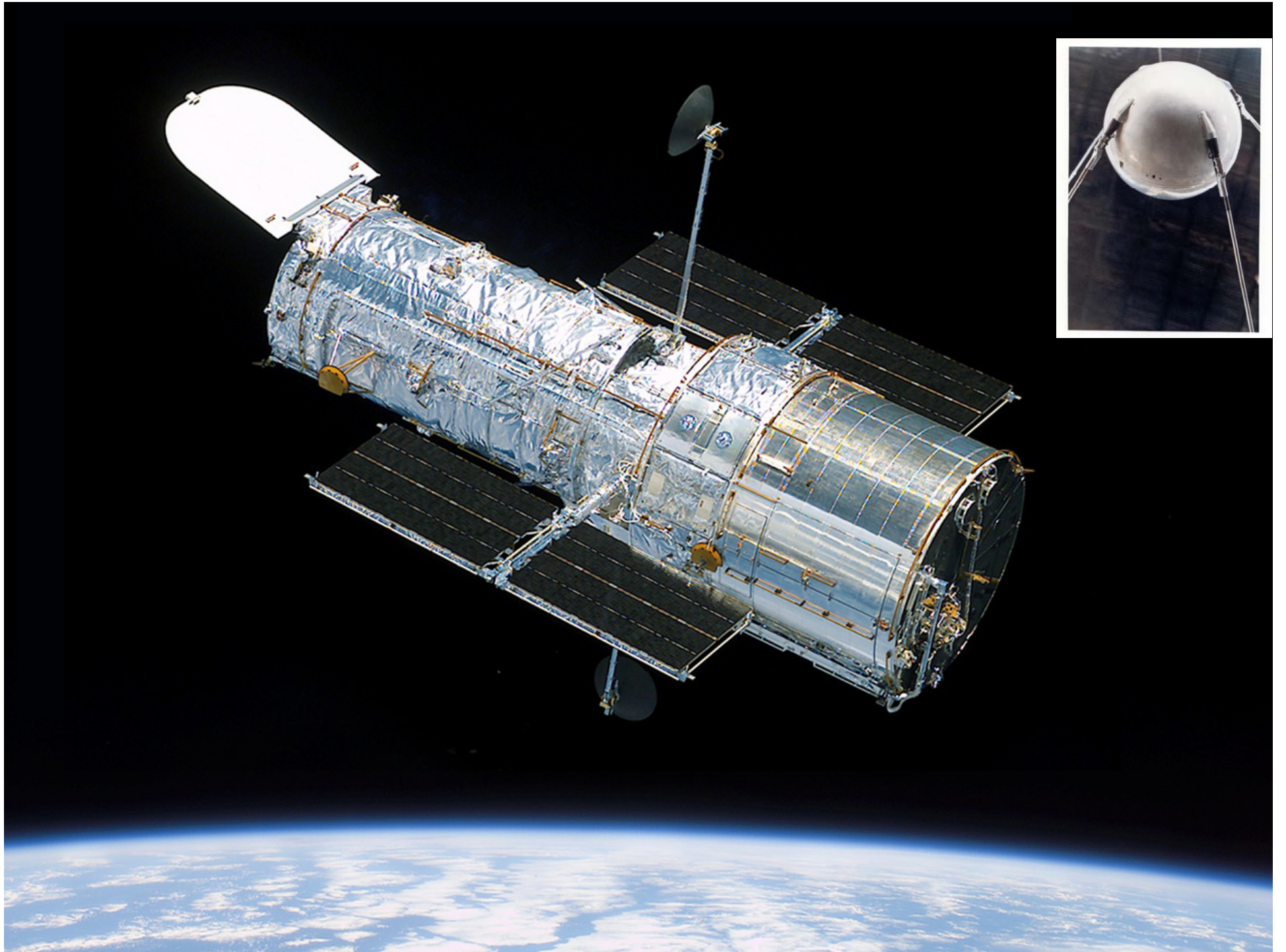




Miglioramento
rispetto
all'occhio umano



Adapted from *Cosmic
Discovery*, M. Harwit



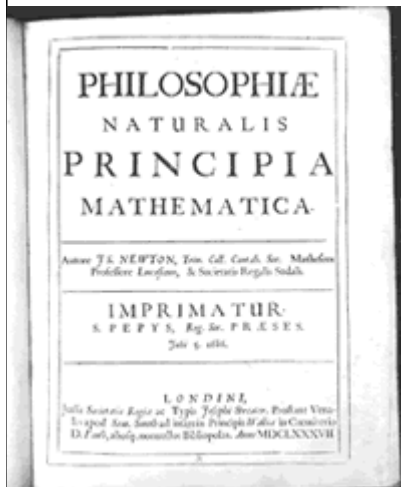




La legge di gravitazione universale

Isaac Newton (1643-1727)

1687

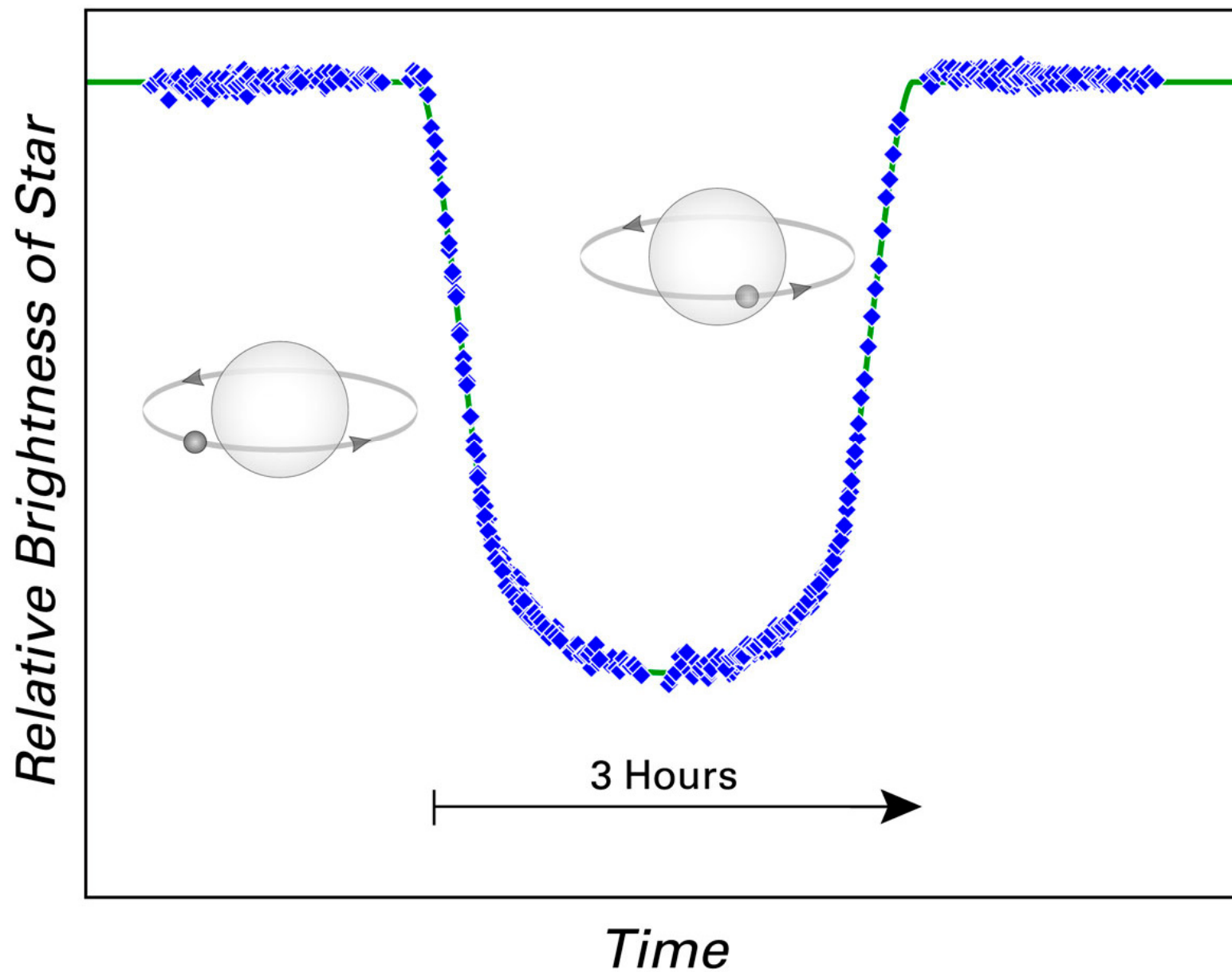


$$F = -G \frac{M_1 M_2}{R^2}$$

Da Newton a Kepler

$$\frac{R^3}{P^2} = C$$

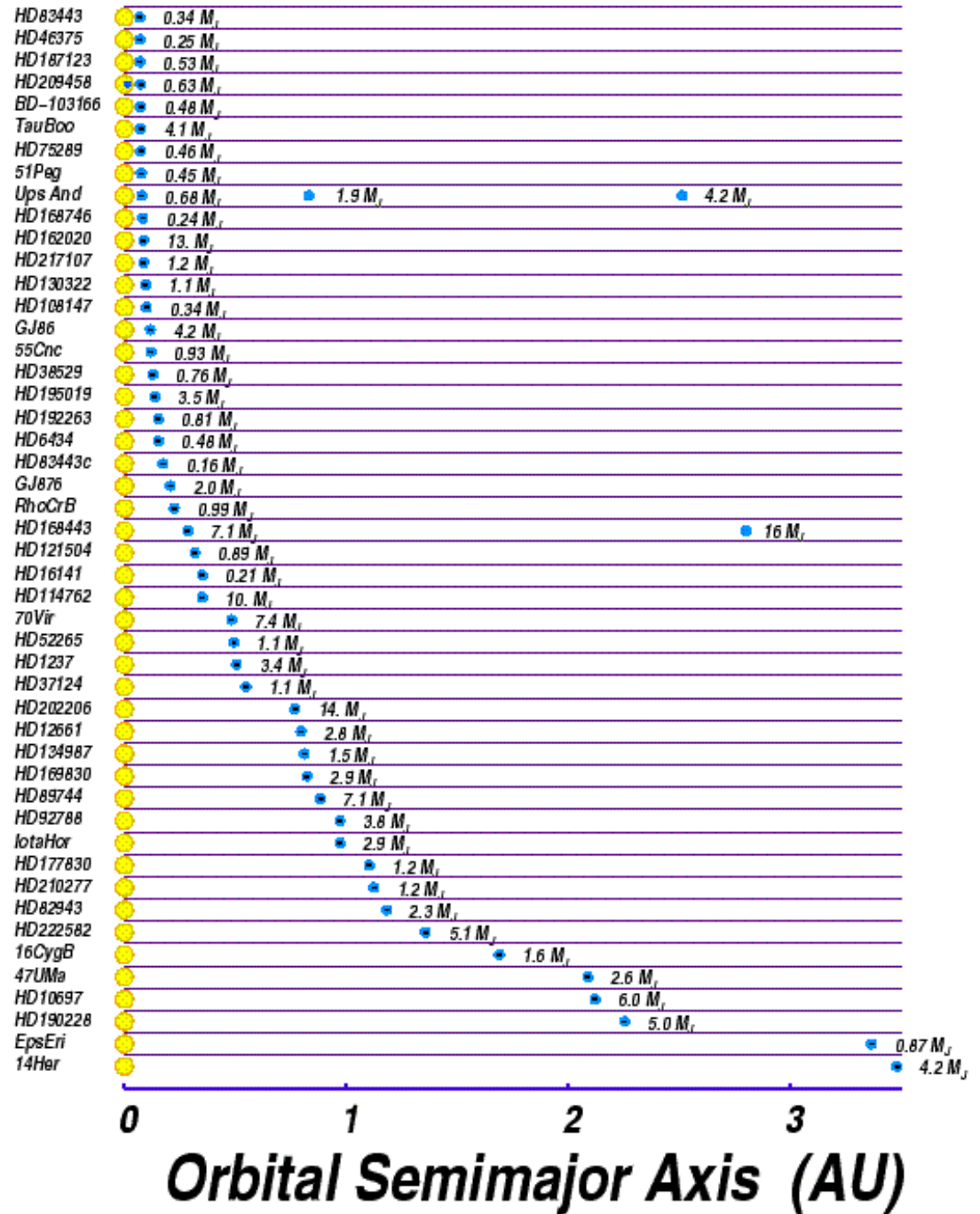
Planet Eclipsing Star HD 209458



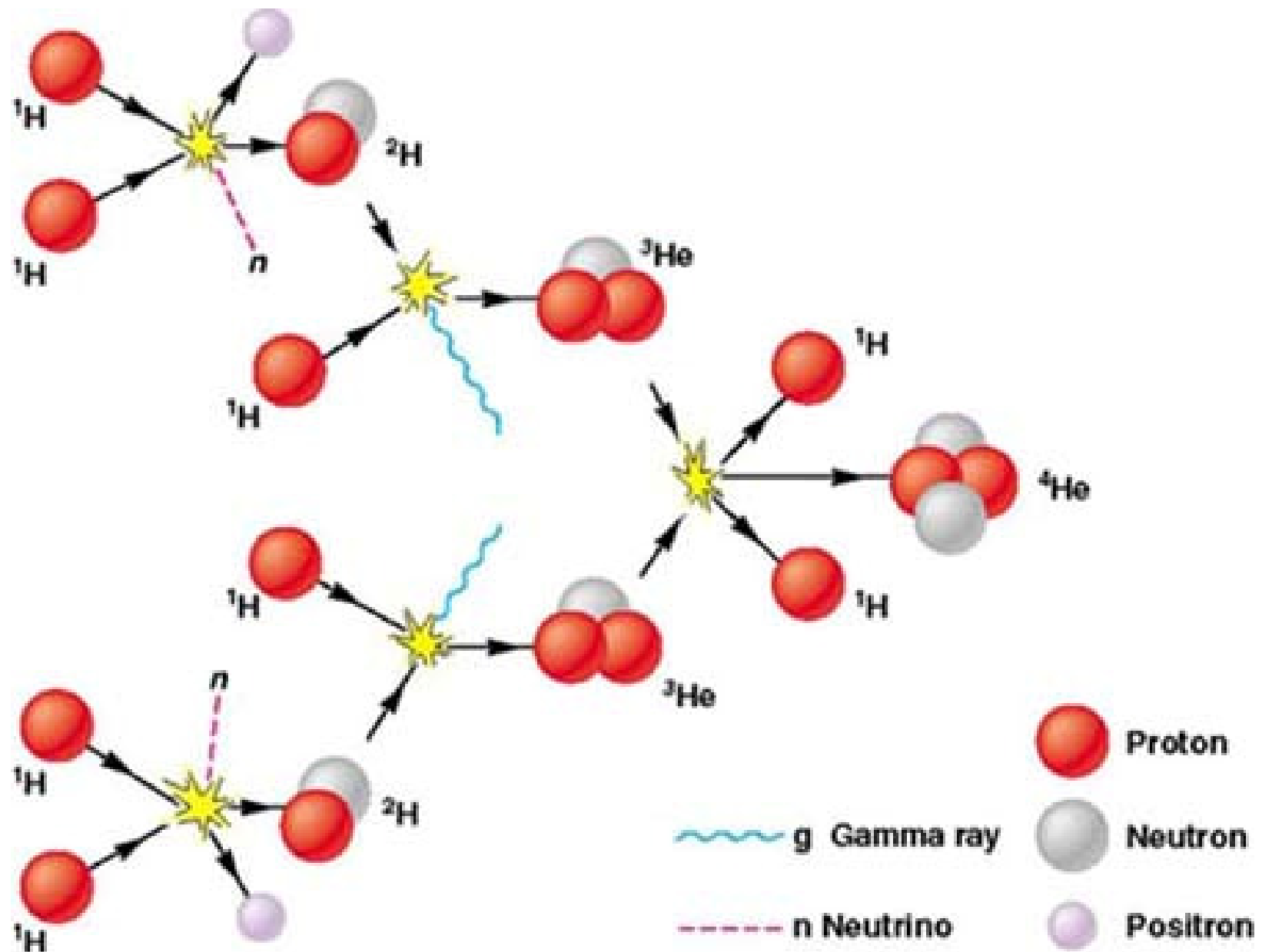


Fino a dieci anni fa non sapevamo neppure se esistessero pianeti orbitanti altre stelle. Oggi se ne conoscono 180....

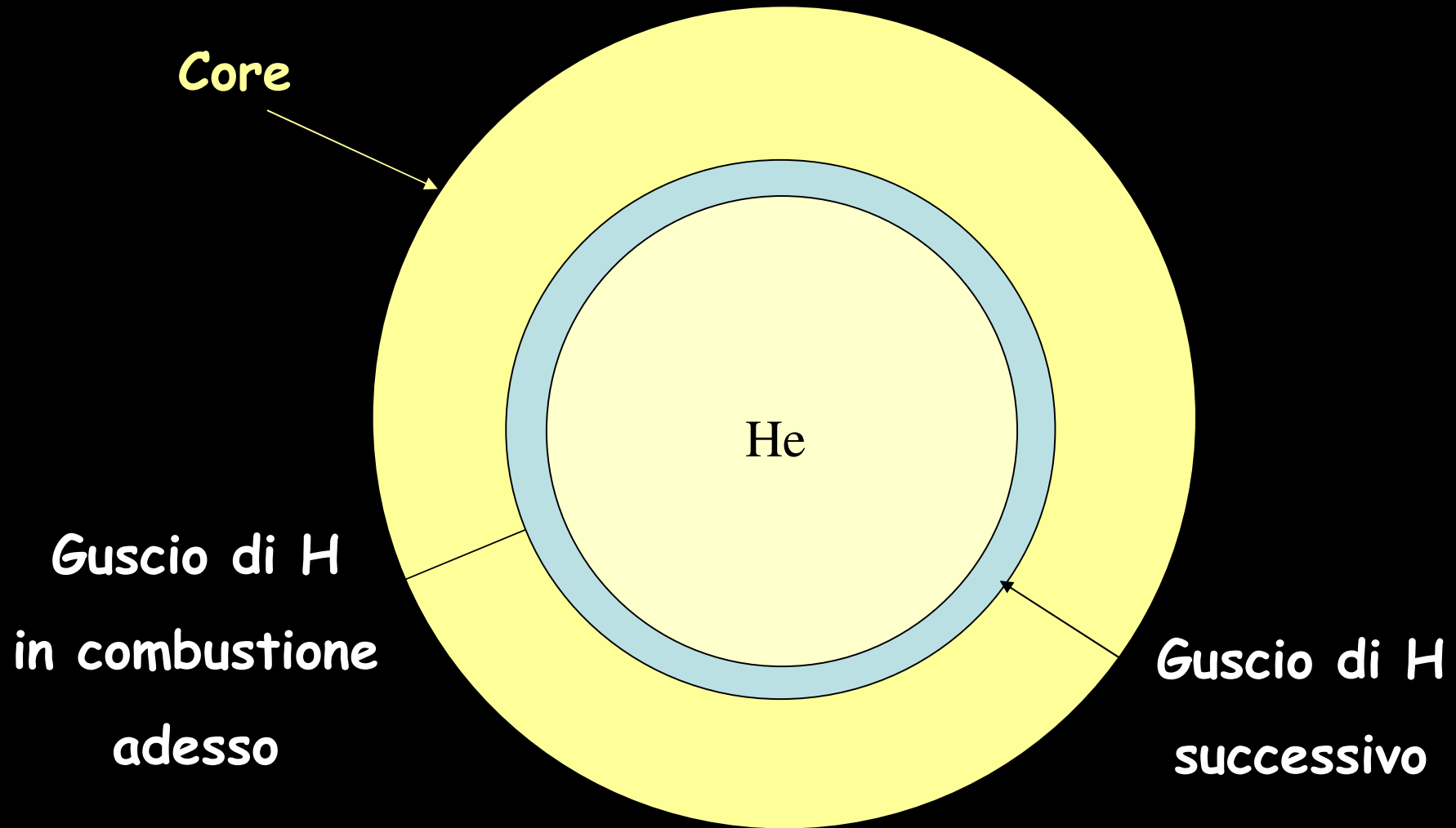
La sfida e' ottenere una immagine del pianeta: rapporto del contrasto richiesto $1:10^{-7}$ → $1:10^{-10}$



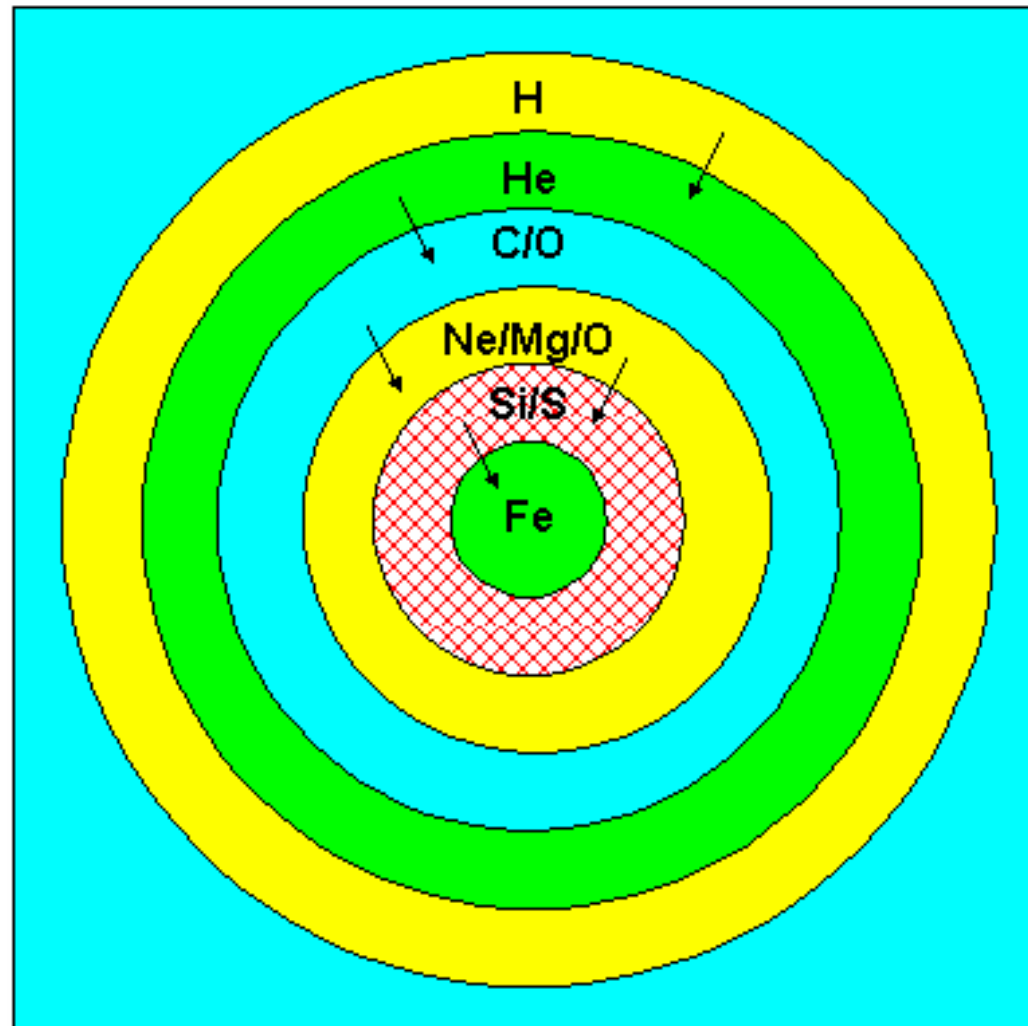
La sorgente di energia delle stelle



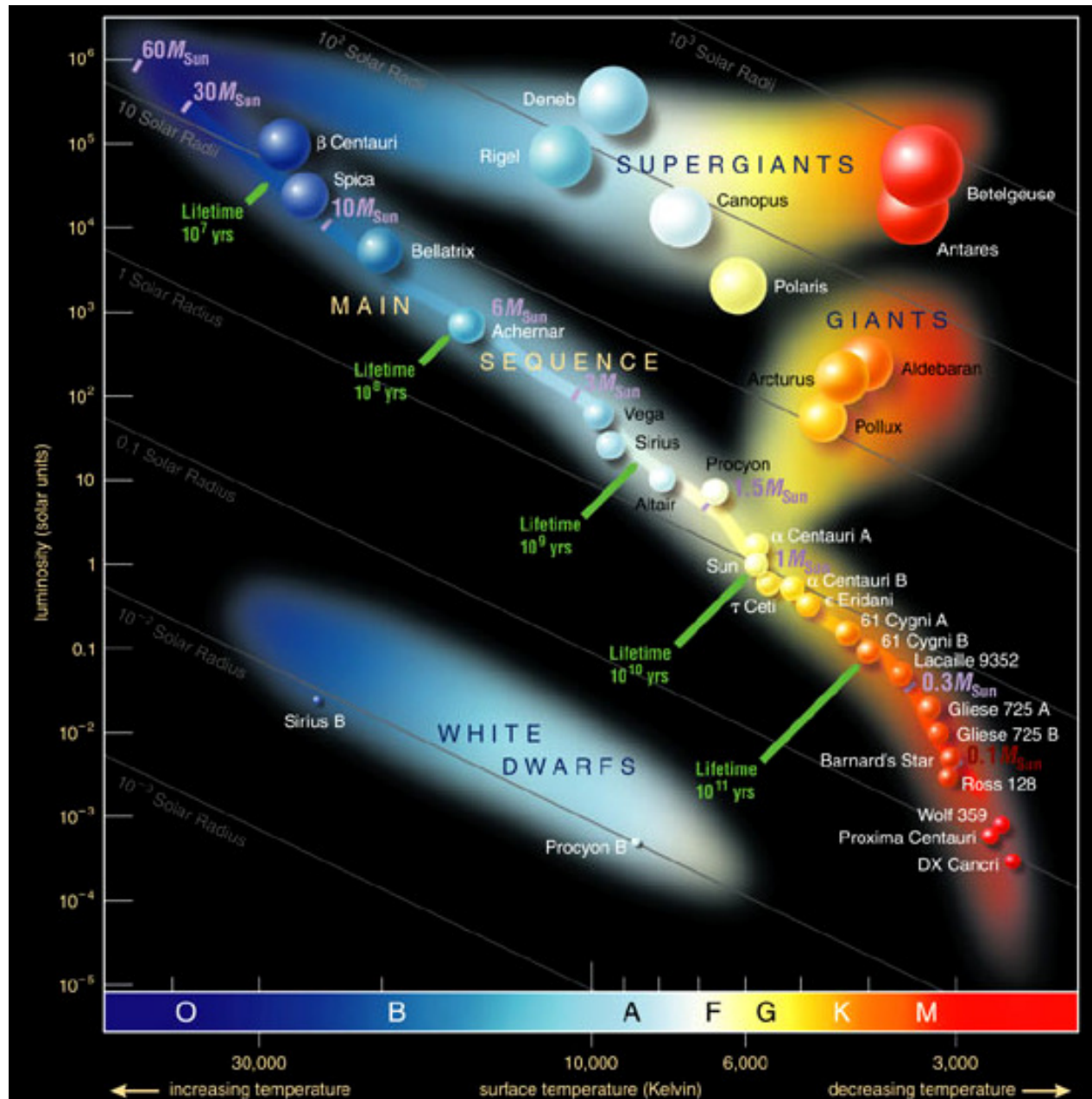
La "cucina" stellare



Struttura di una stella Pre-SN di 20 masse solari



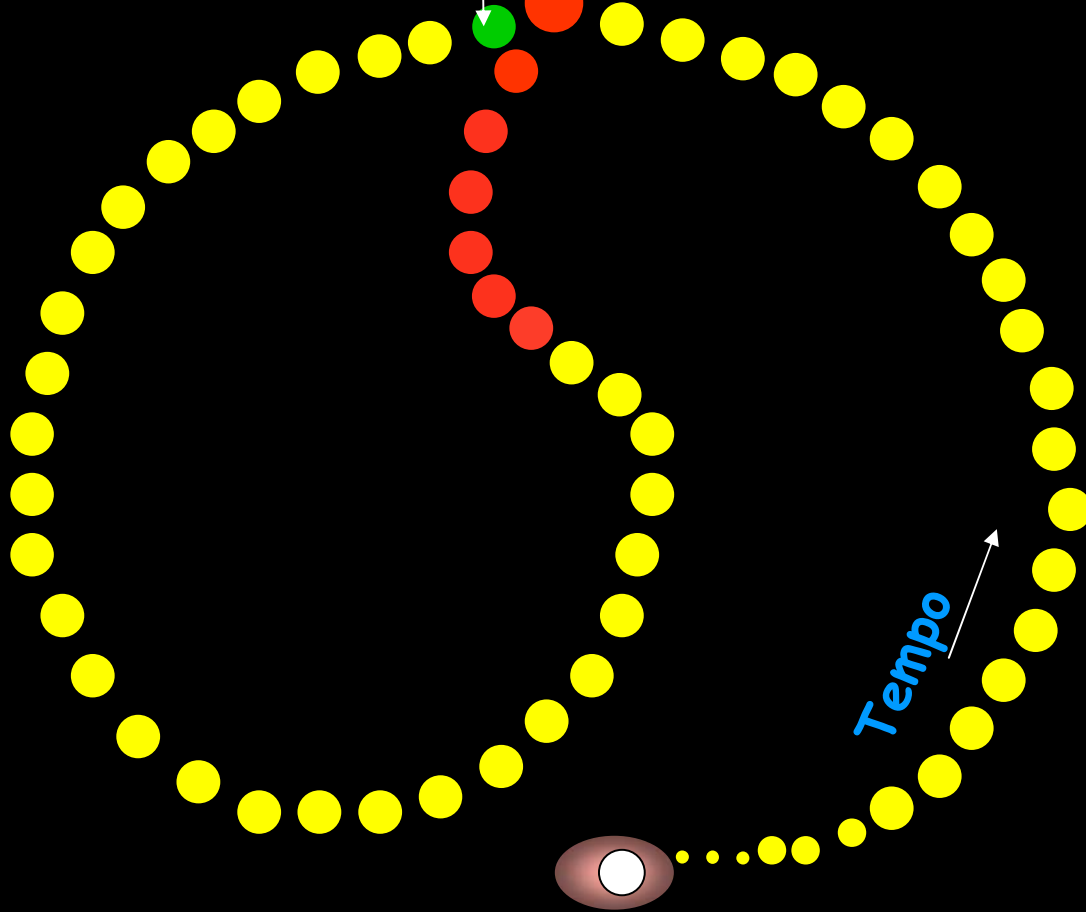
Evoluzione stellare



Il Sole oggi circa
a metà della
combustione
dell'H, 10 miliardi
di anni

Gigante Rossa

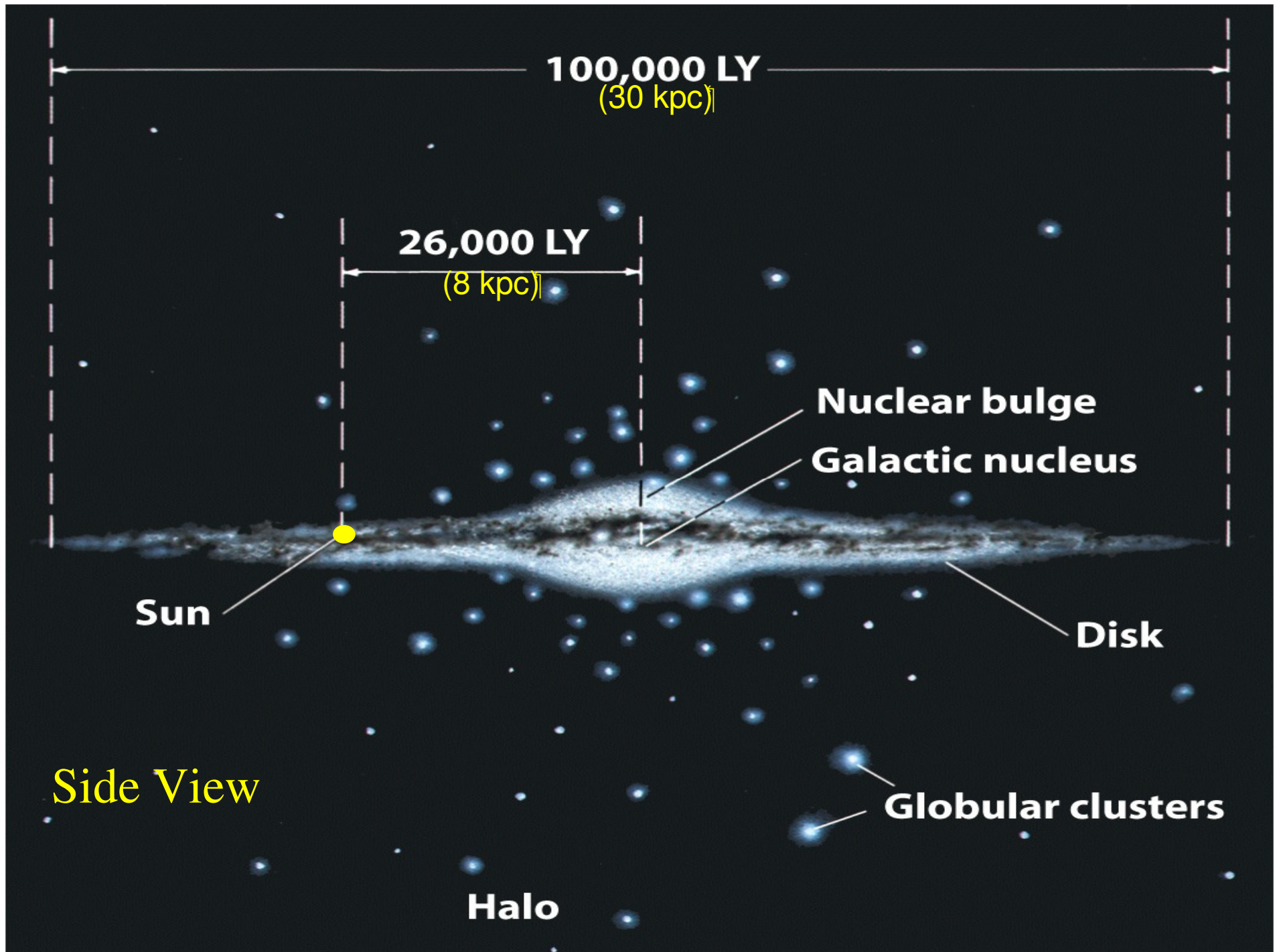
Nana Bianca



Nube

Tempo

L'intervallo tra due
posizioni consecutive
equivale a 150 milioni
di anni



100,000 LY
(30 kpc)

26,000 LY
(8 kpc)

Sun

Nuclear bulge
Galactic nucleus

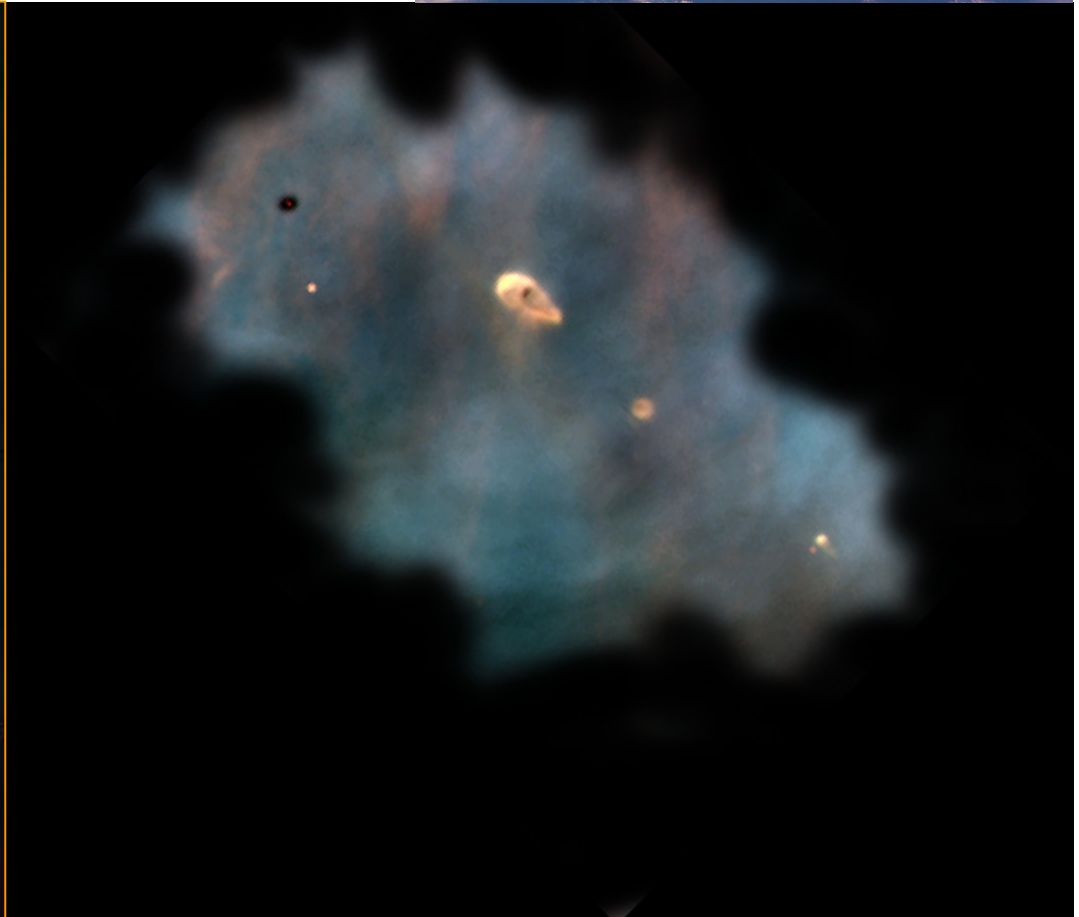
Disk

Side View

Globular clusters

Halo

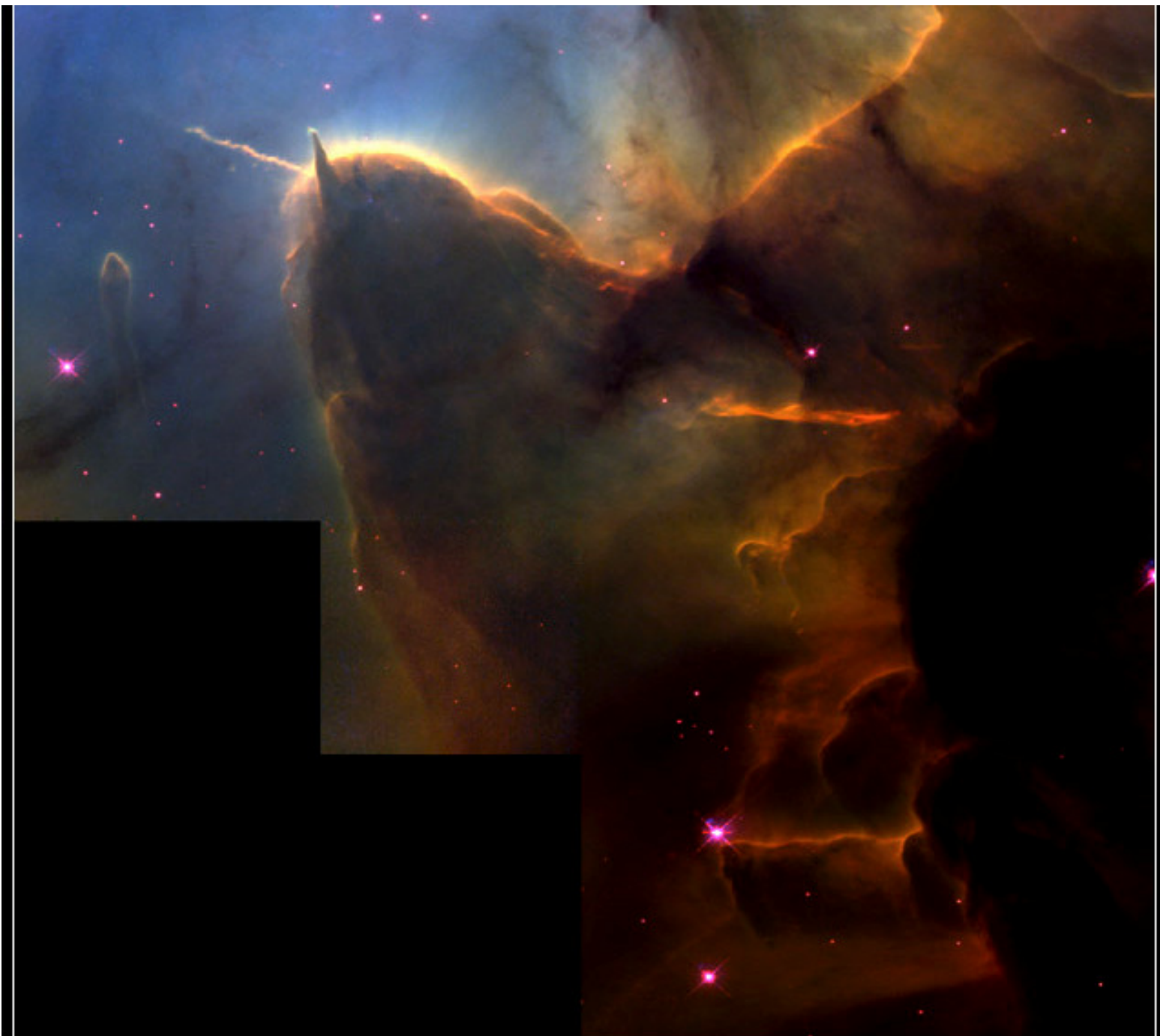
Formazione stellare



Carina Nebula







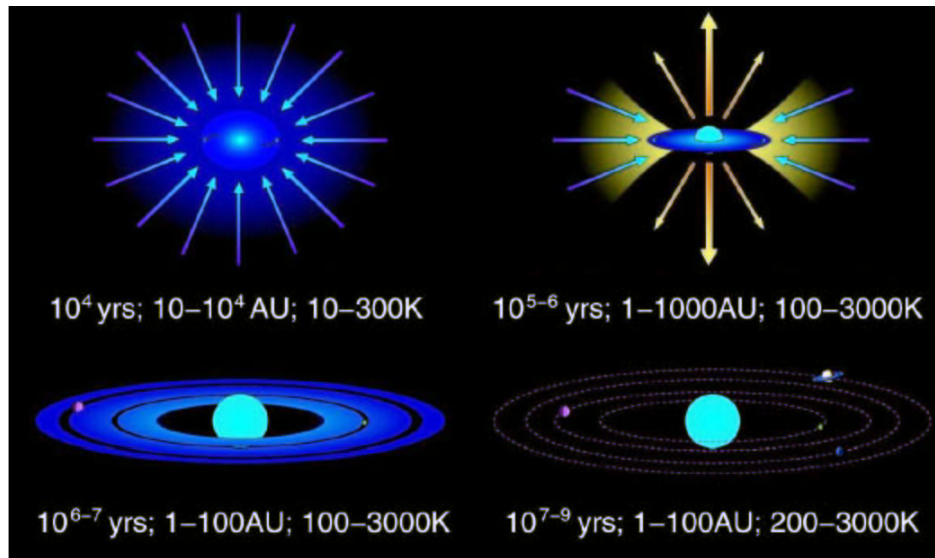
Trifid Nebula • M20

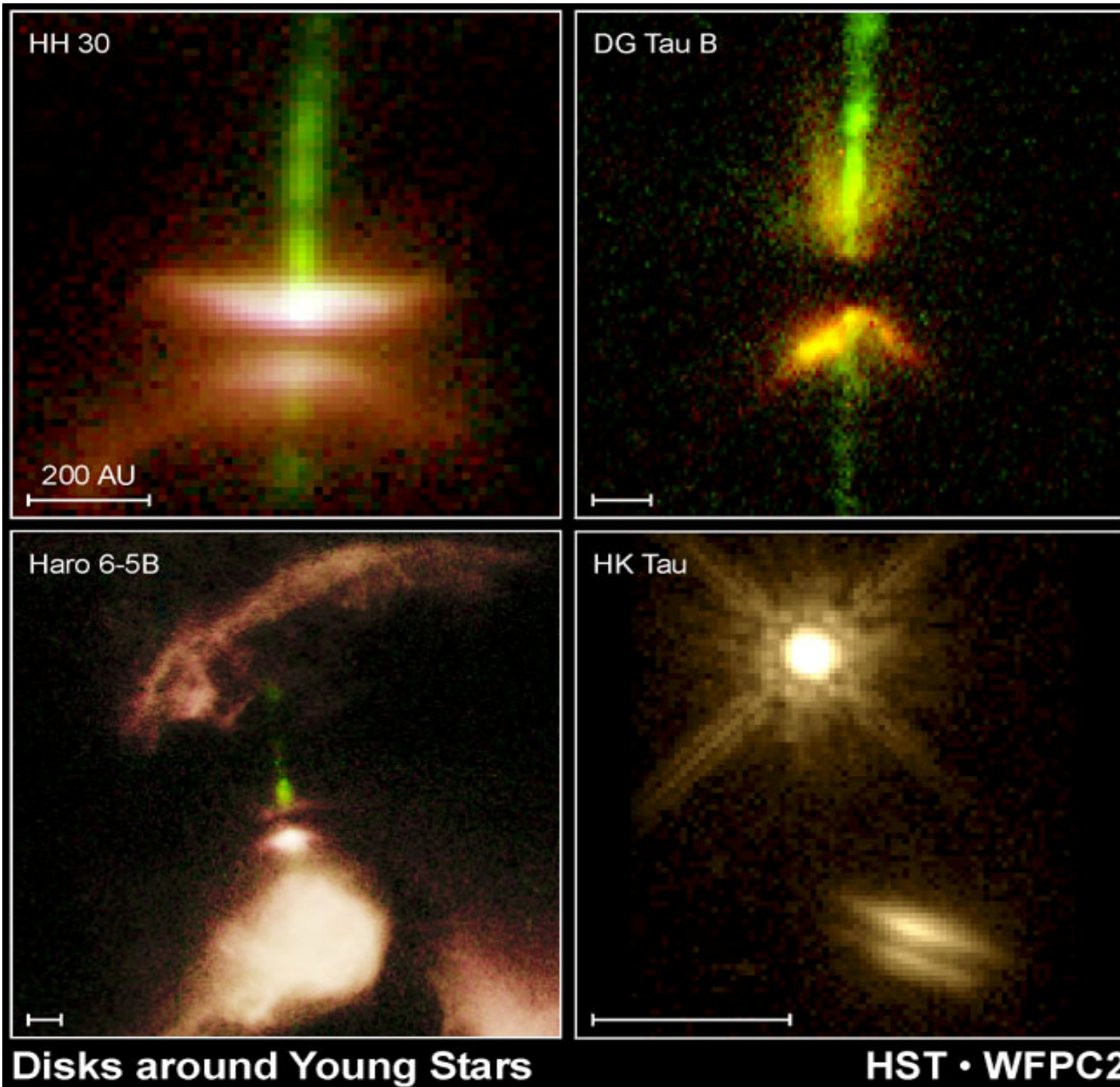
HST • WFPC2

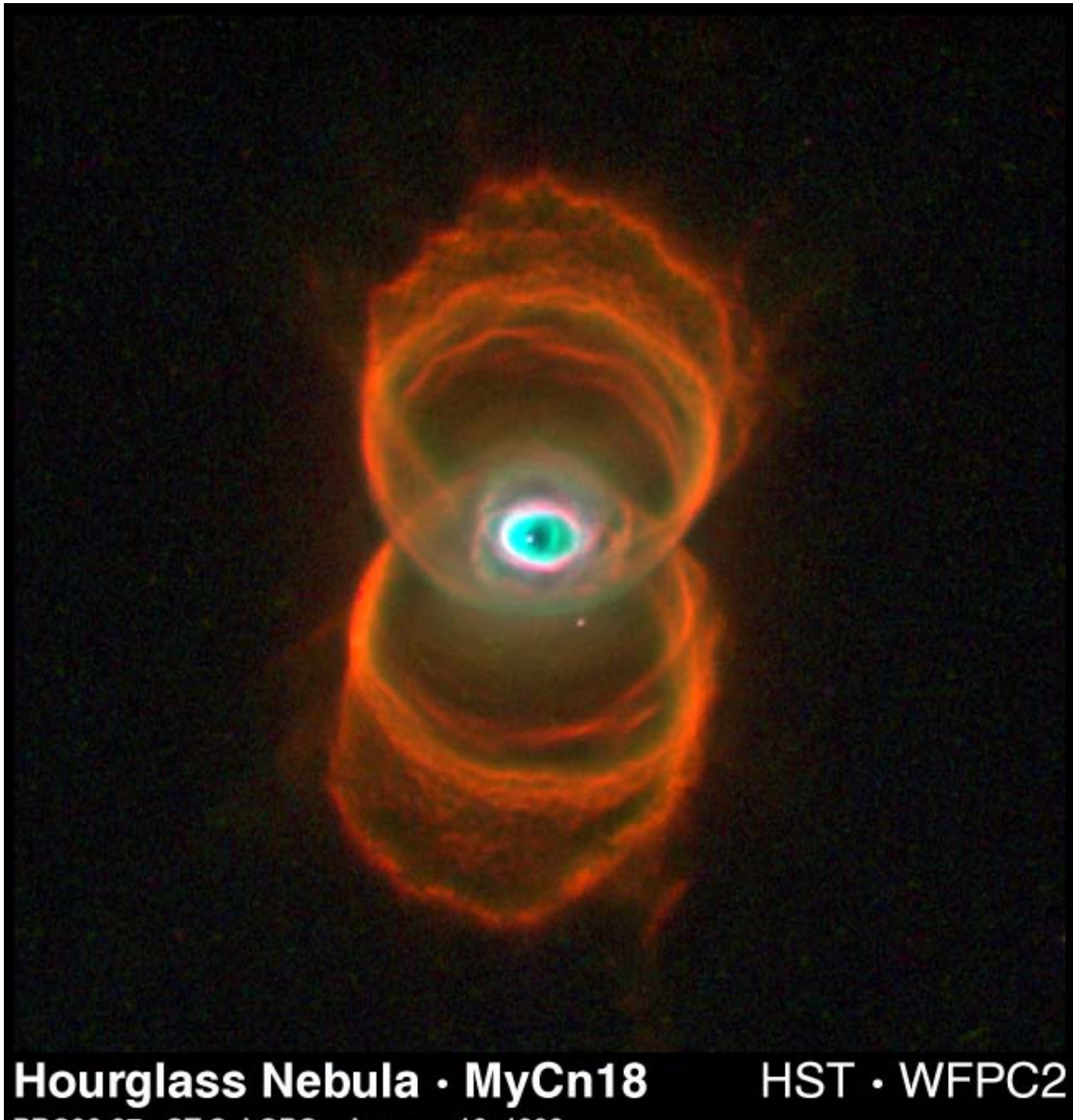
Nascita di stelle e sistemi protoplanetari



Eagle Nebula HST

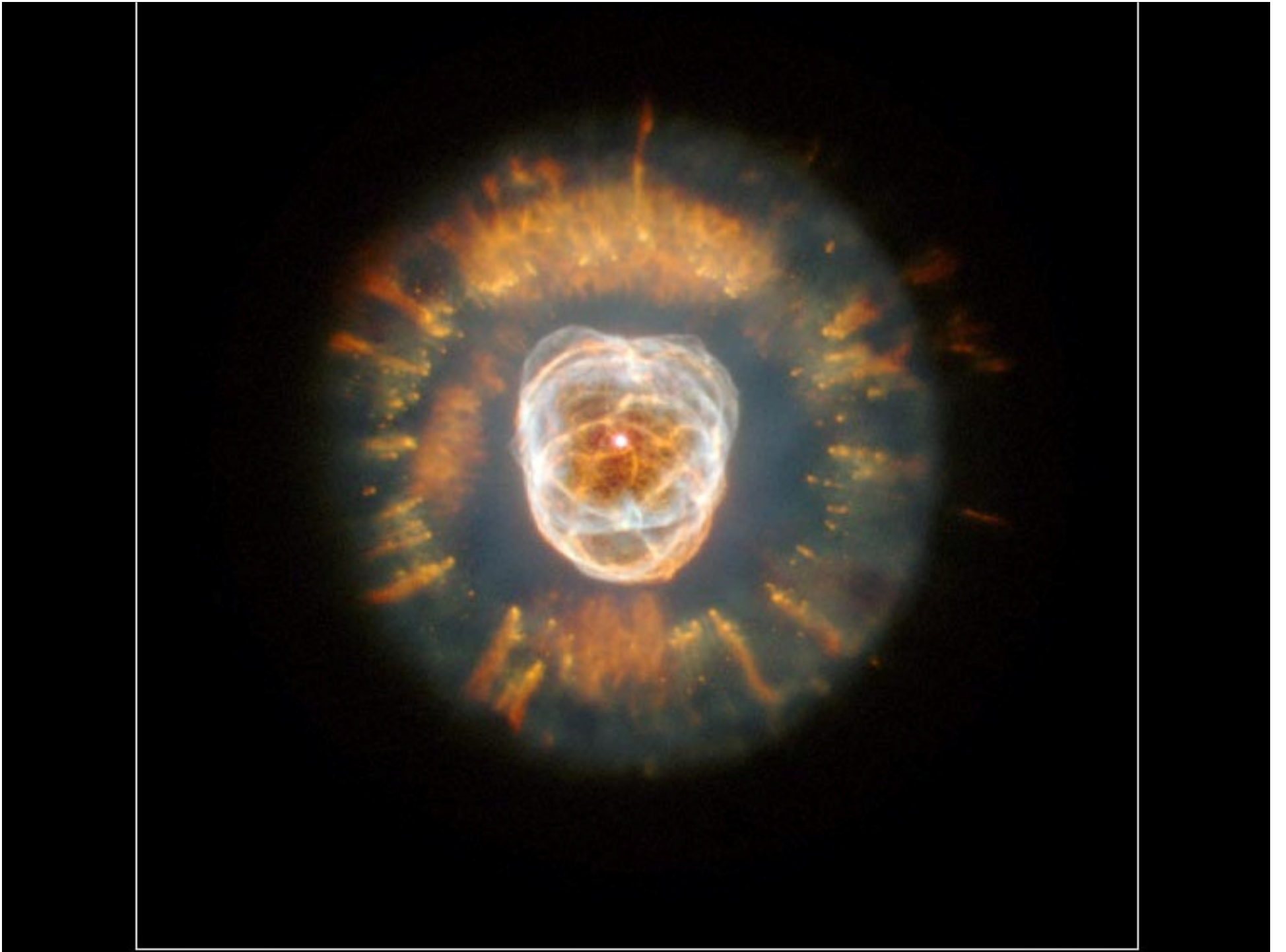






Hourglass Nebula · MyCn18

HST · WFPC2





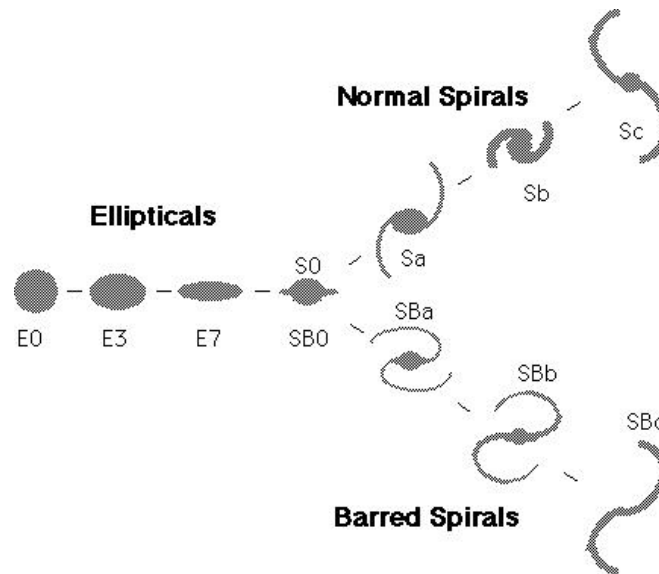


Edwin P. Hubble

(1889-1953)

Mt. Wilson
Observatory
Pasadena, CA

Galassie



La sequenza di Hubble







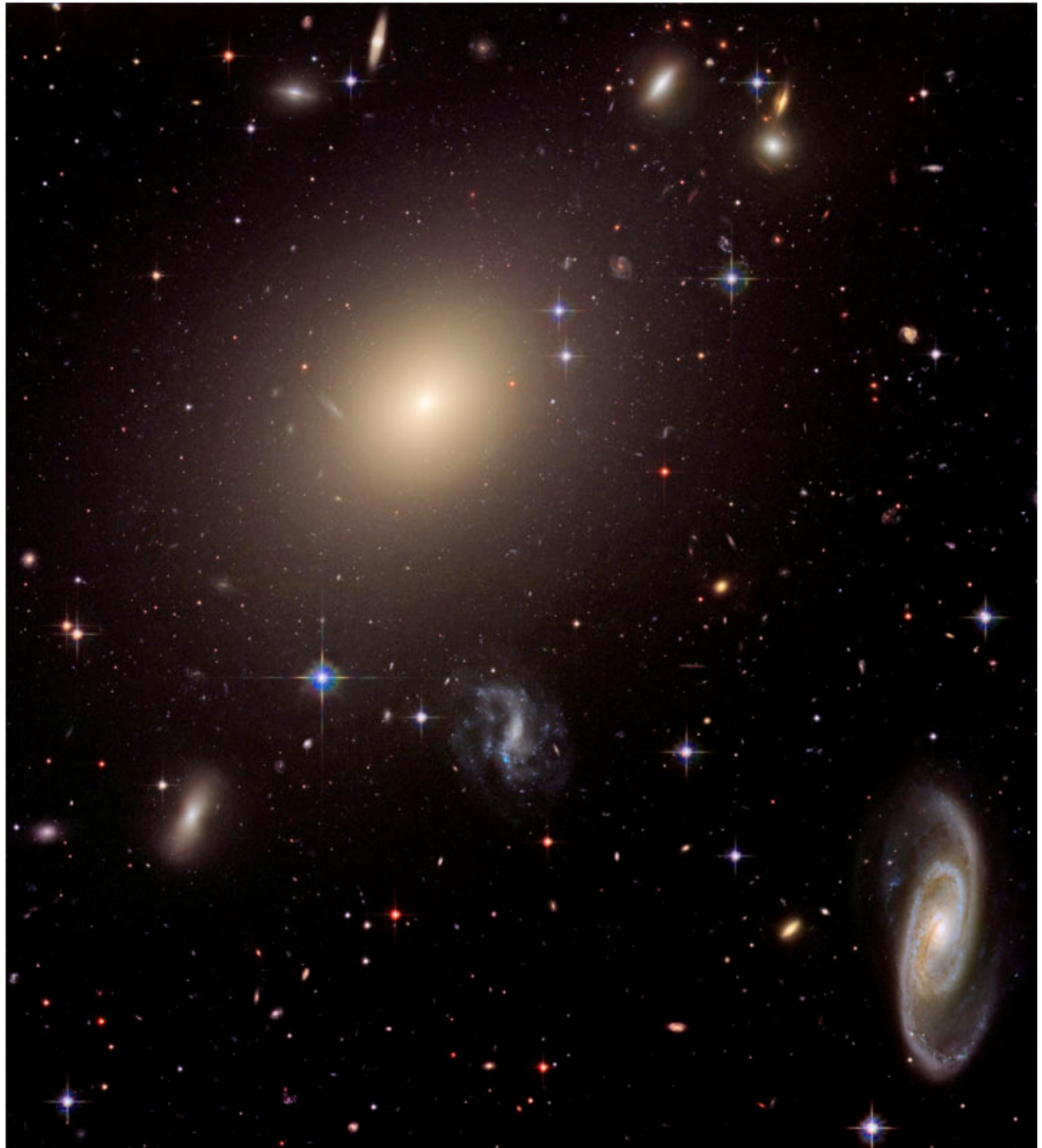


Galaxies NGC 2207 and IC 2163

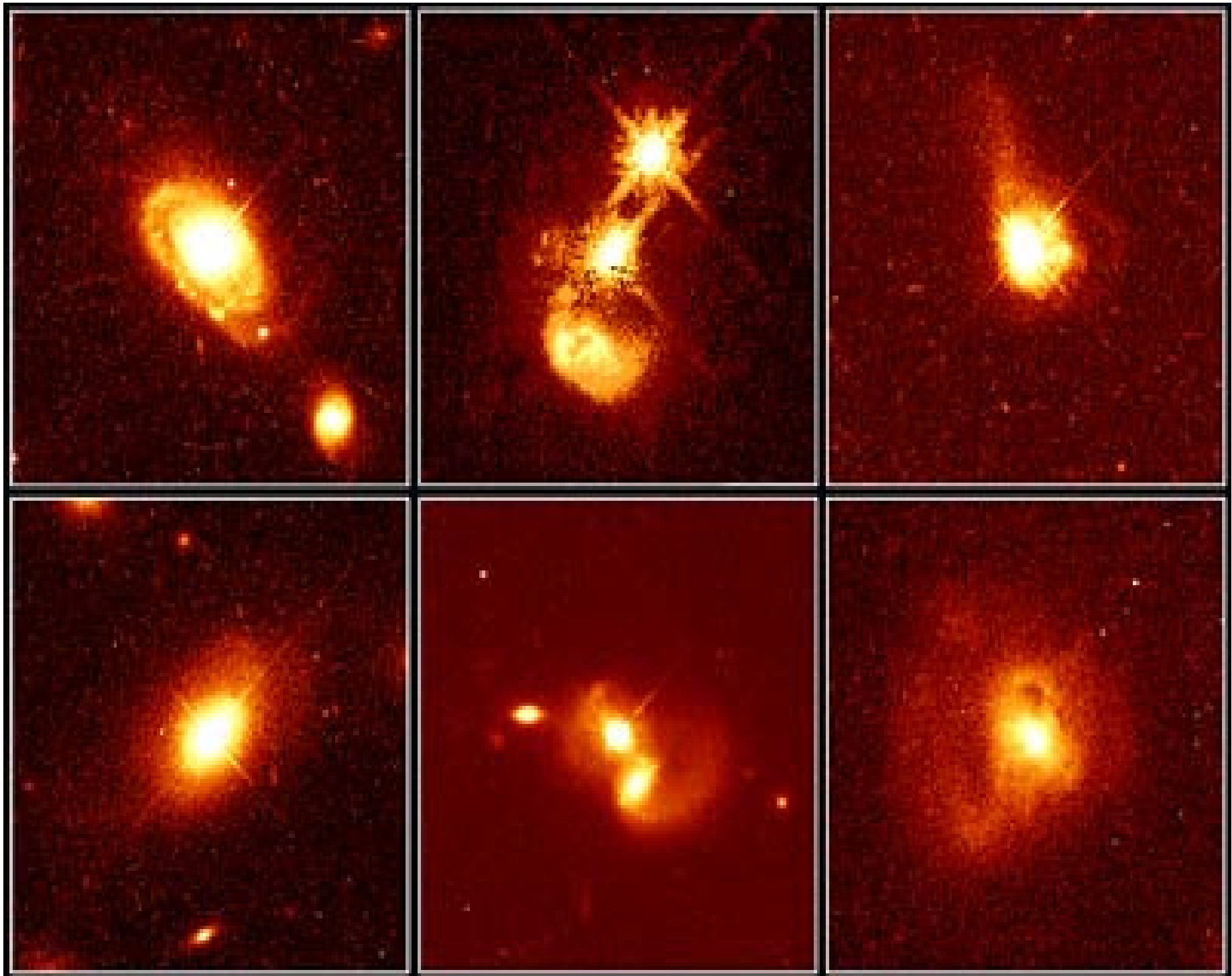


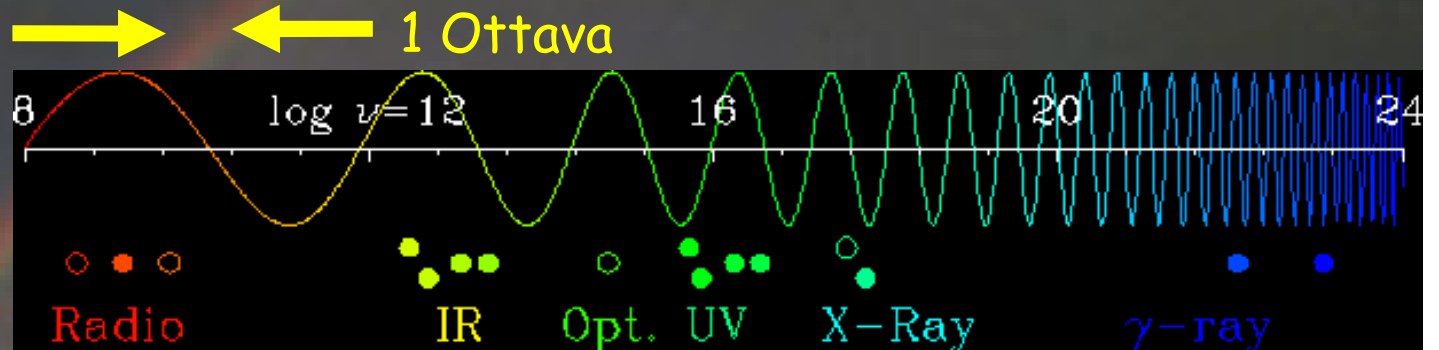
Hubble
Heritage

Ammassi di Galassie









50 Ottave (7 „Piani a coda“)

7 octaves



7 octaves



7 octaves



7 octaves



7 octaves



7 octaves



7 octaves

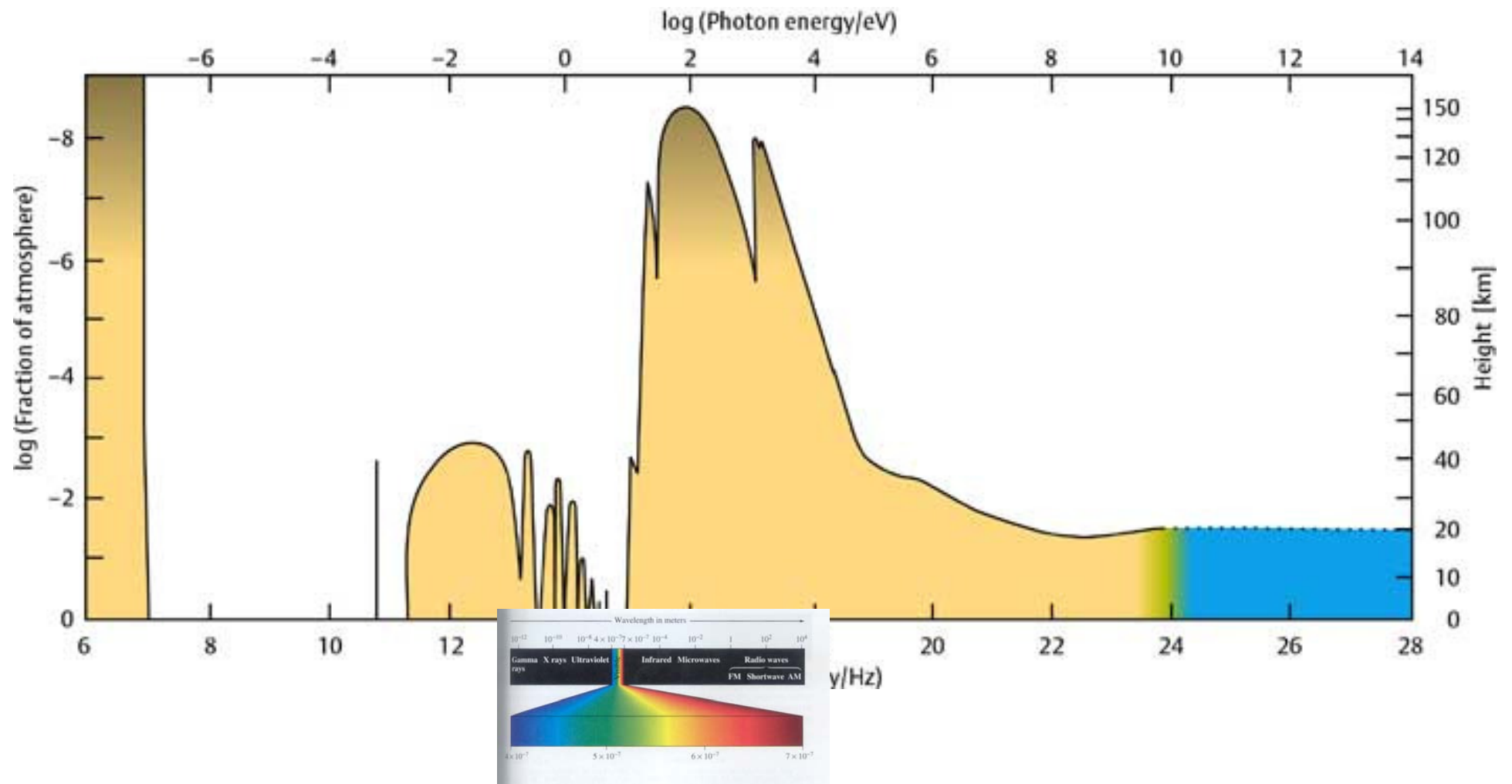


cosmic
concerto

Transpose up by ~50 octaves = 7 pianos (~7 octaves each)



human
concerto



Trasparenza atmosferica alla radiazione elettromagnetica



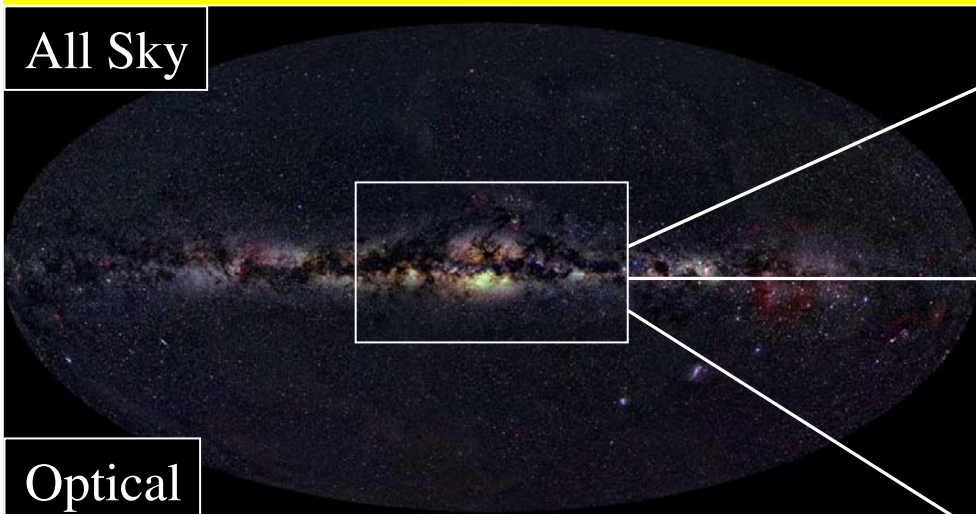
Edward Hopper
Rooms by the sea
(1951)

Il caso favorisce solo le menti preparate

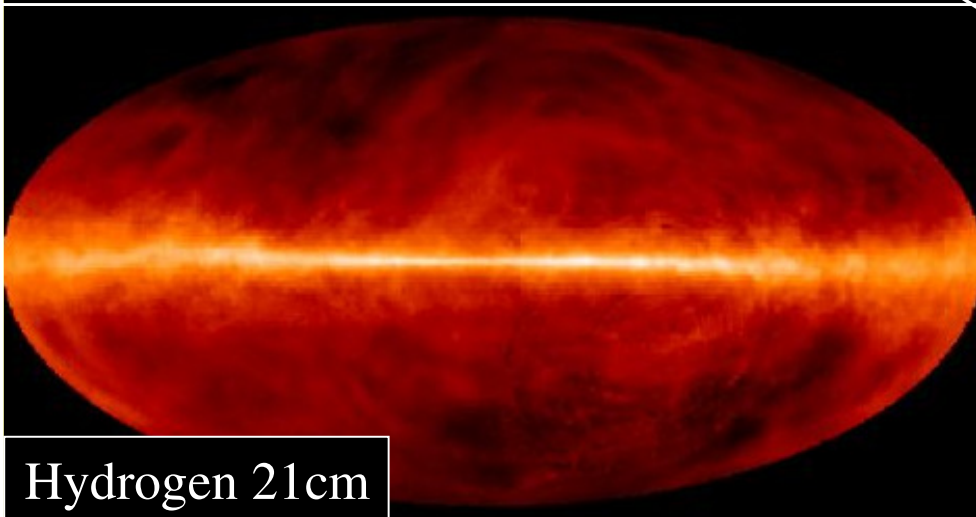
Louis Pasteur

Ottico/Infrarosso/Radio

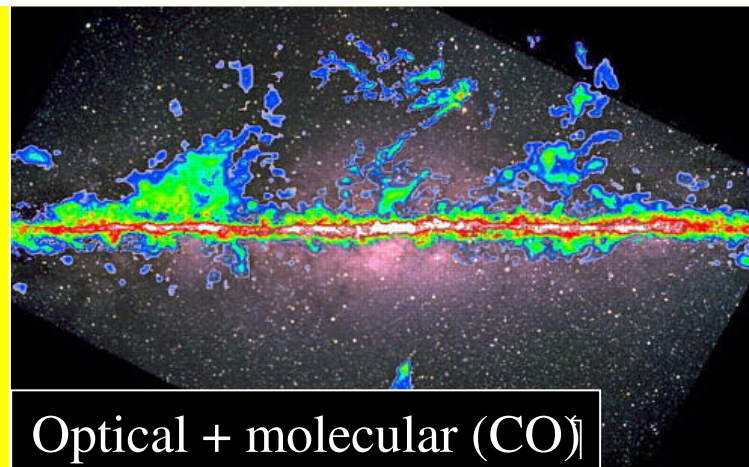
All Sky



Optical



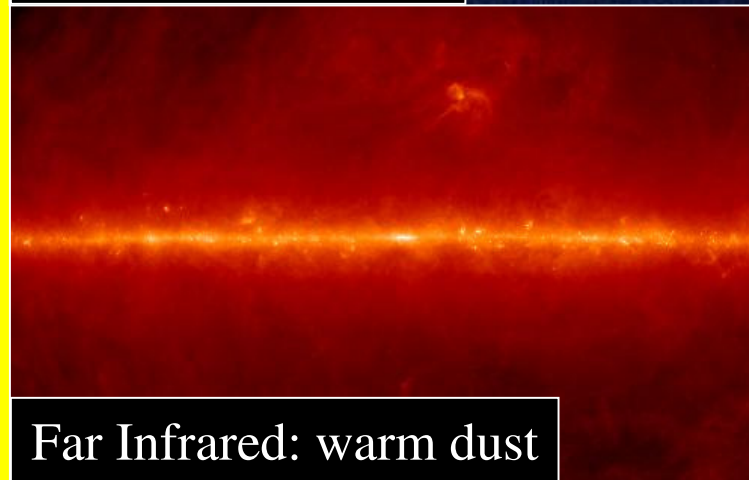
Hydrogen 21cm



Optical + molecular (CO)

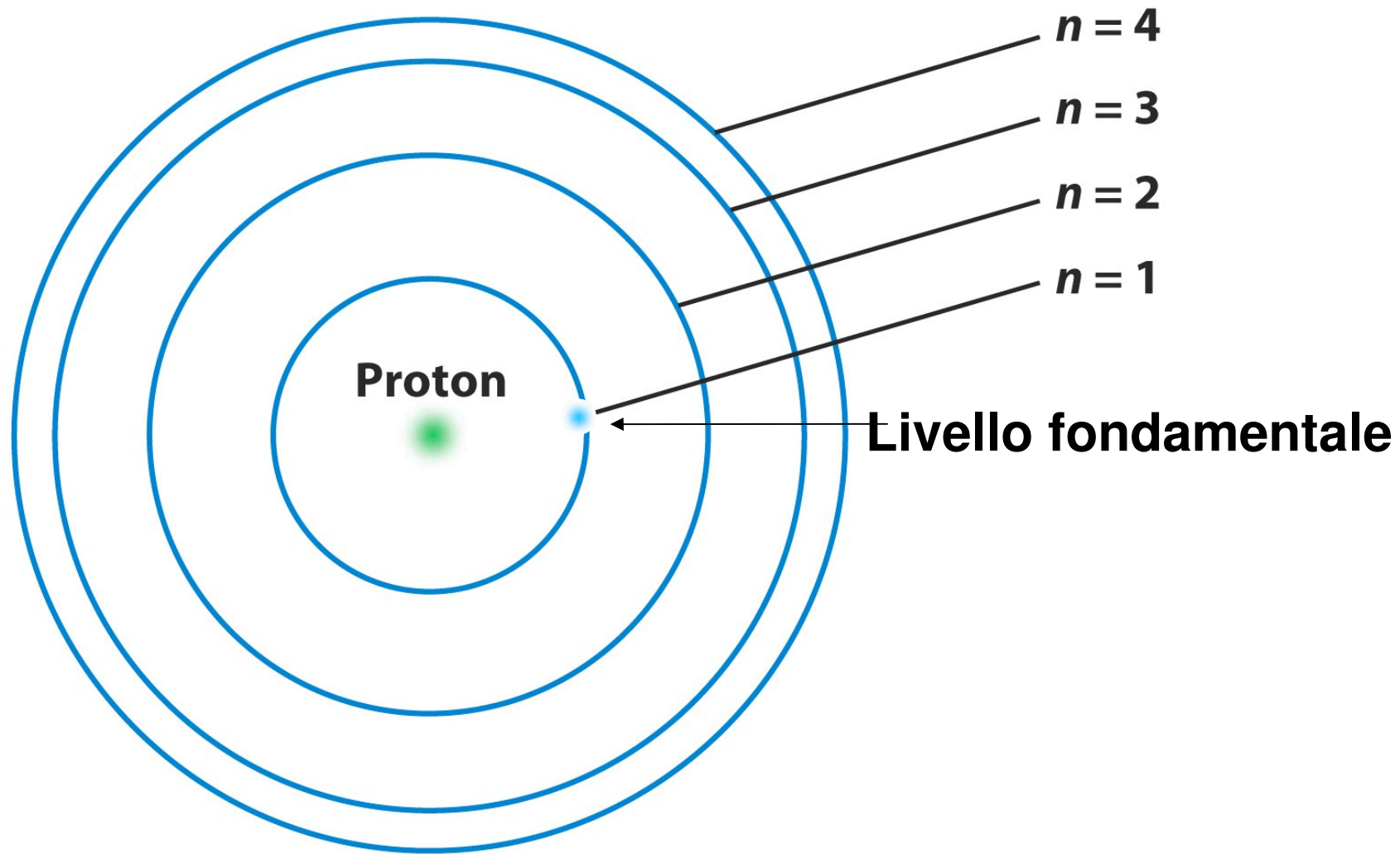


Near Infrared: stars



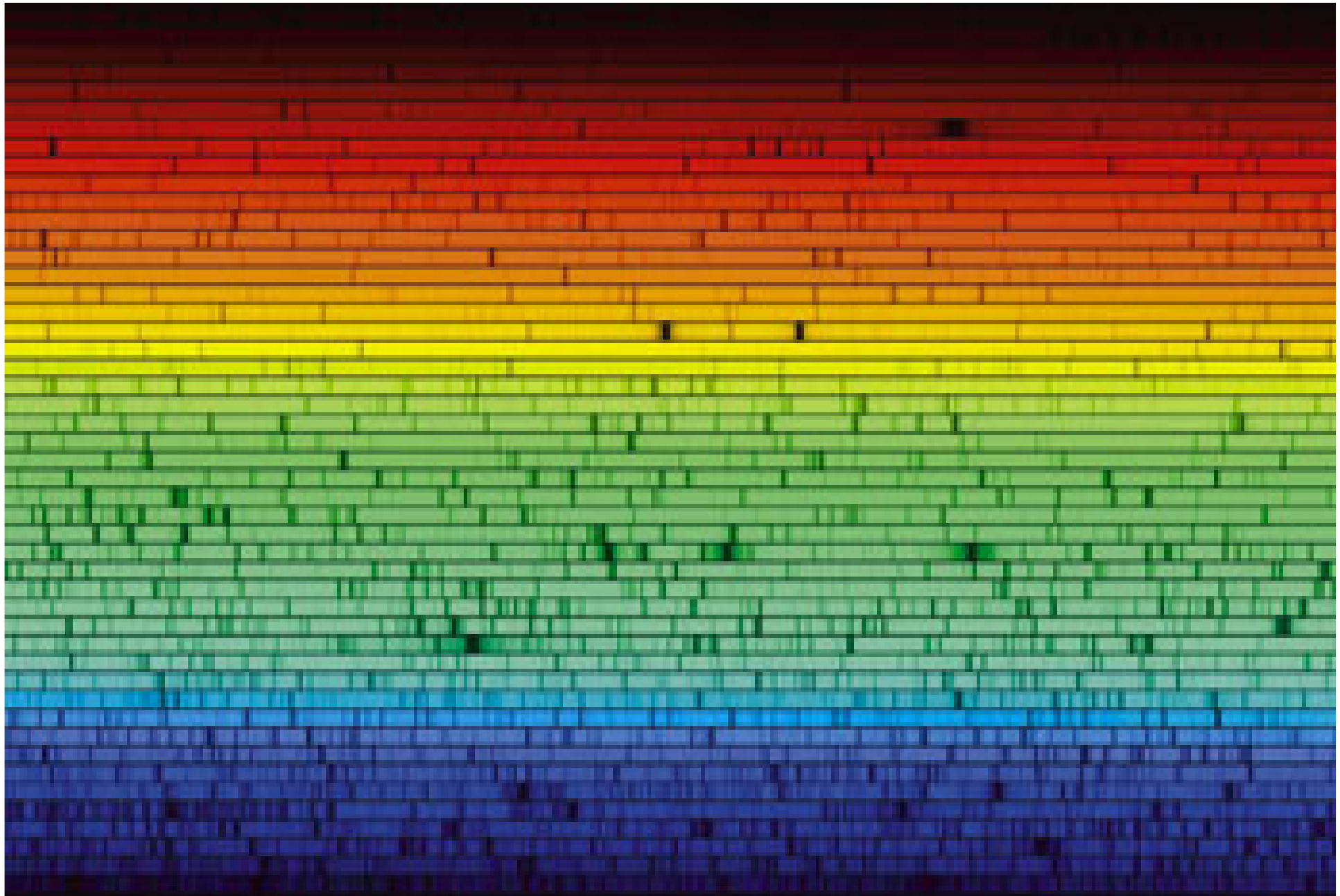
Far Infrared: warm dust

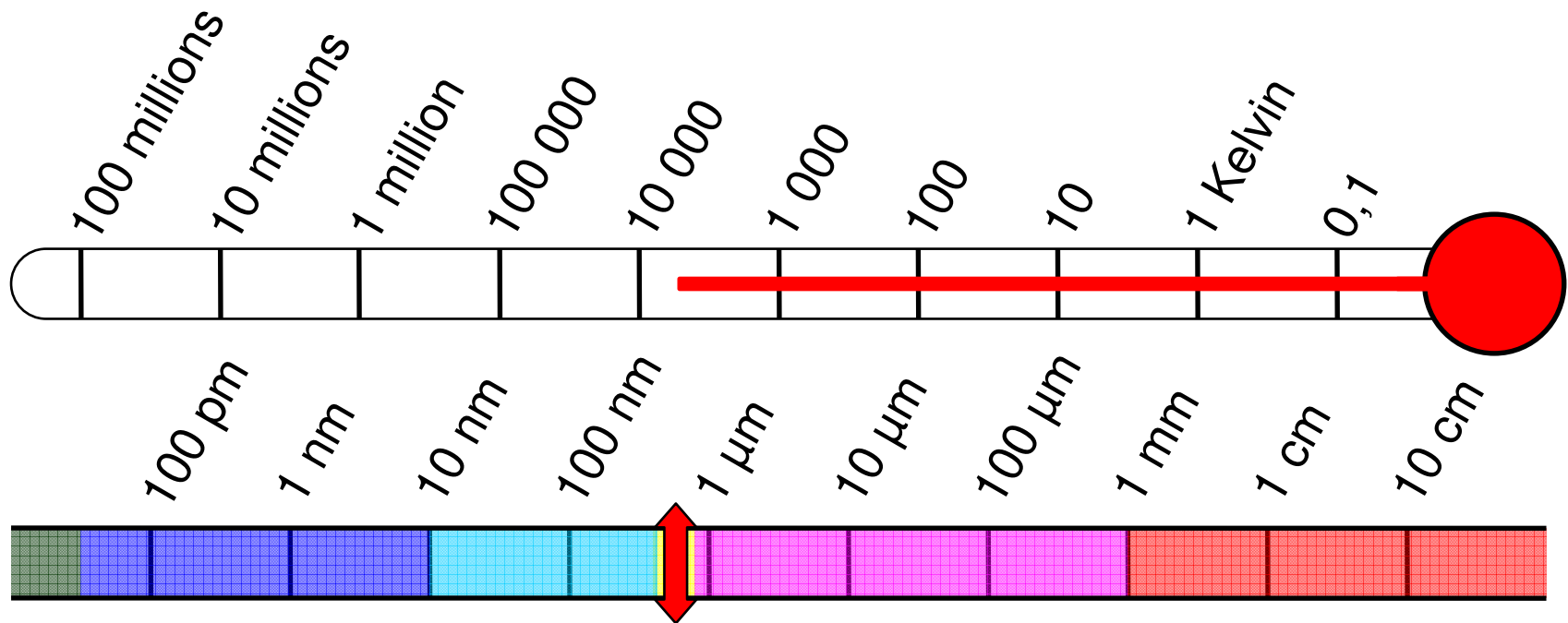
Struttura Atomica



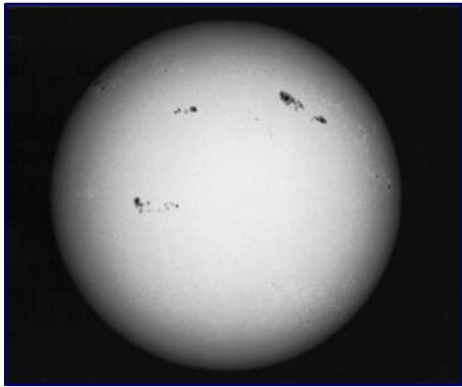
Atomo di H

Fine dell' 800 lo spettro del Sole

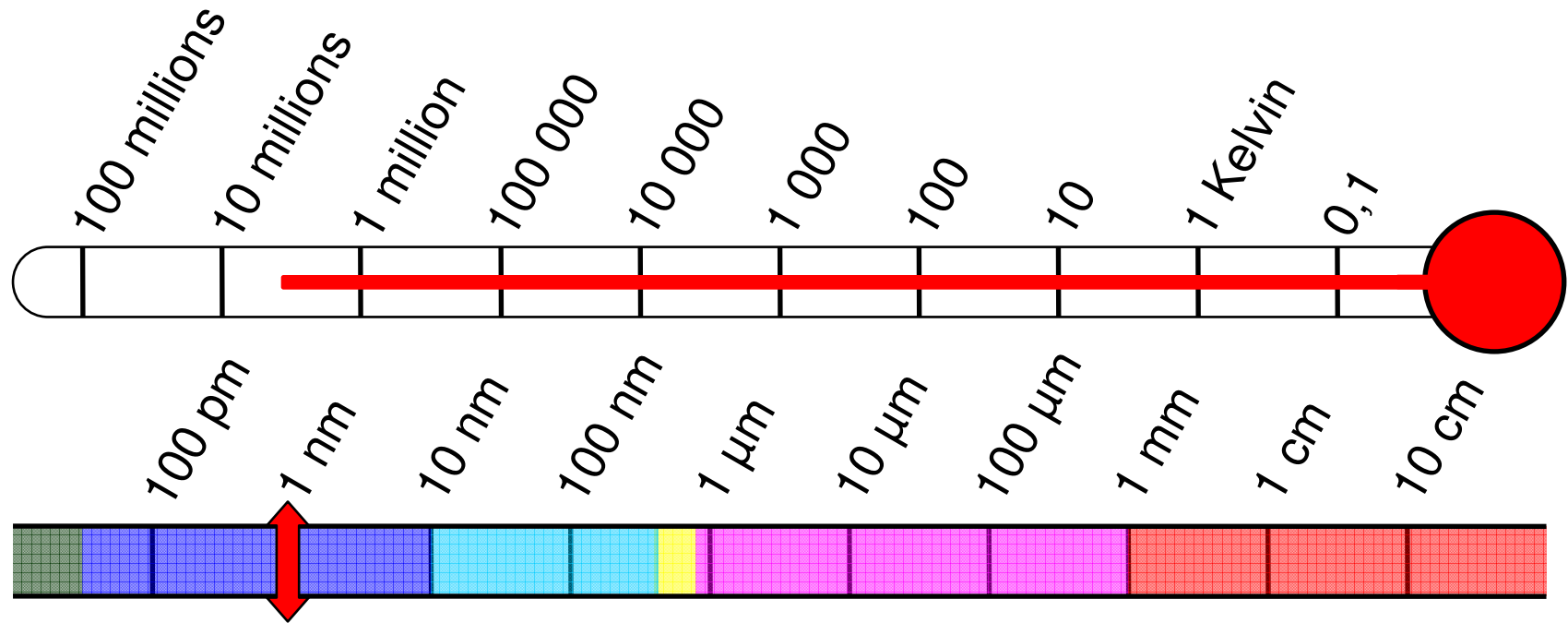




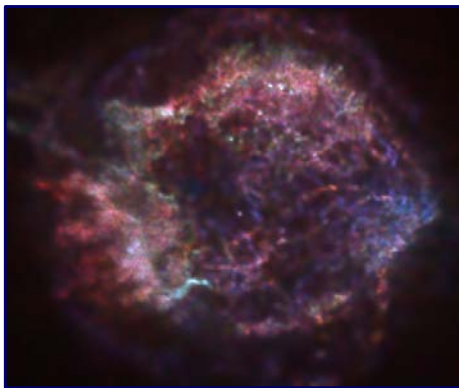
II VISIBILE



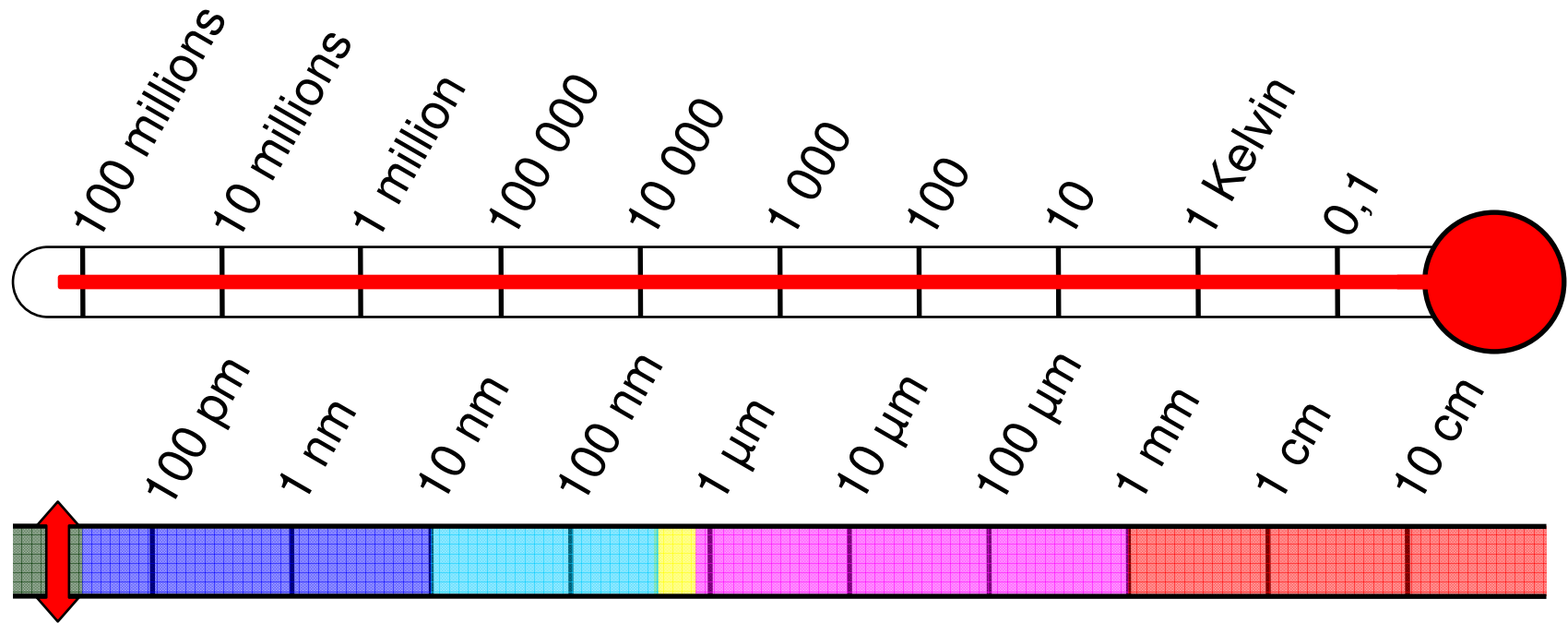
II SOLE



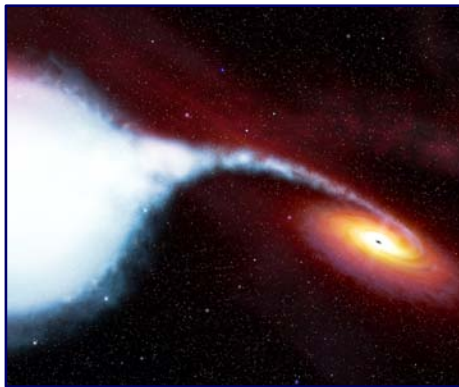
I RAGGI X



L'Universo
Caldo
Resti di
Supernova



I raggi γ



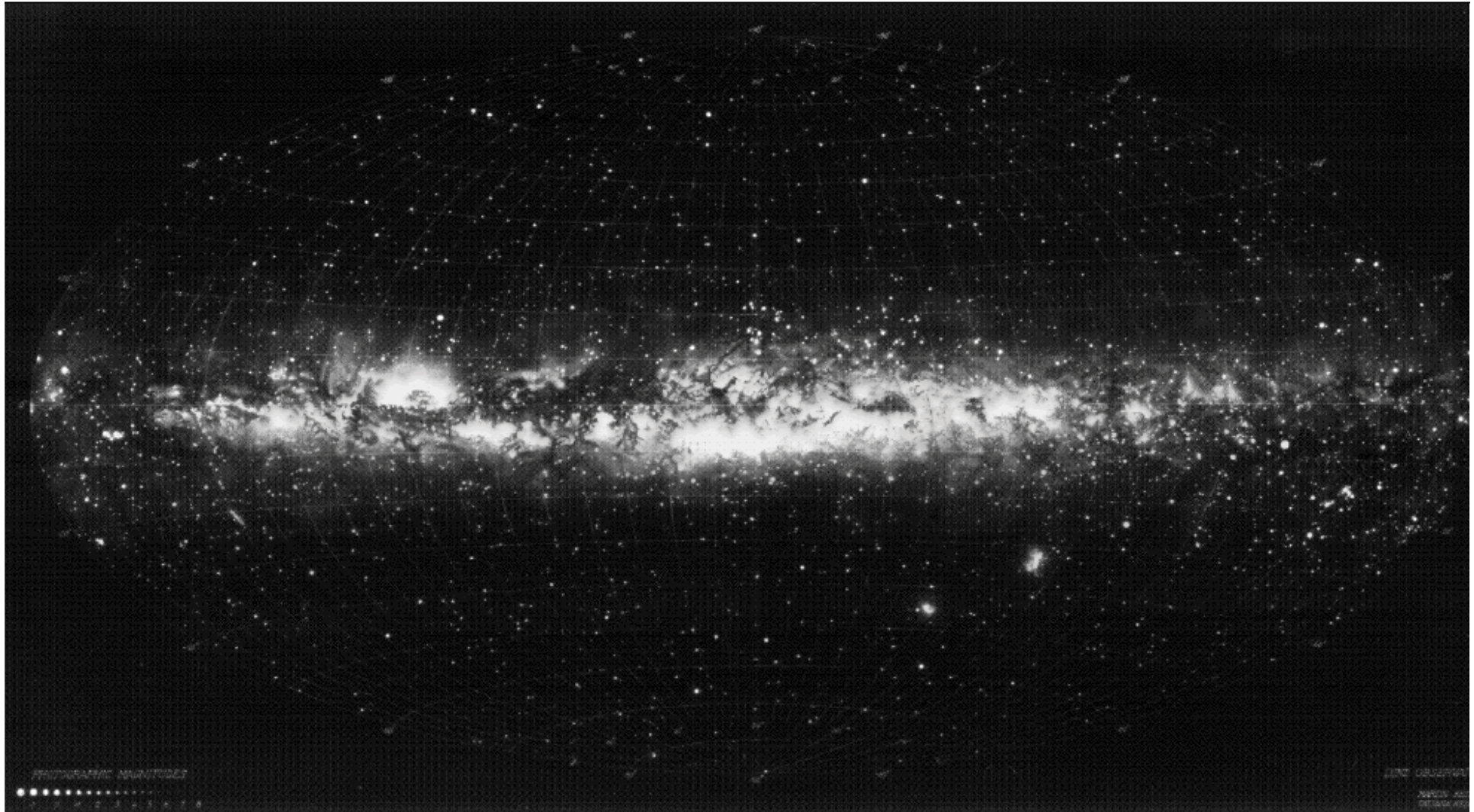
L'Universo
estremo
I buchi neri

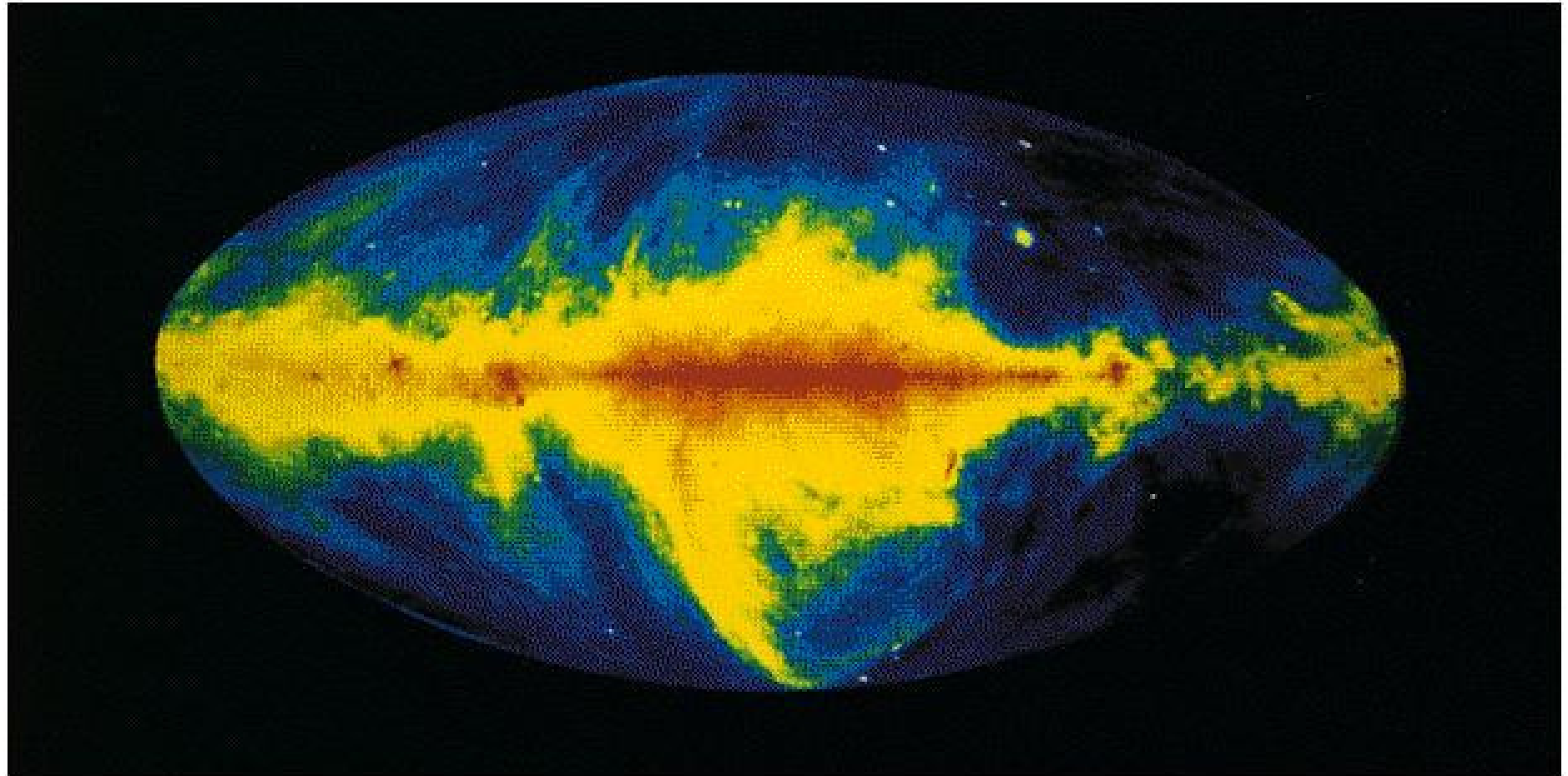
I radiotelescopi





La Via Lattea





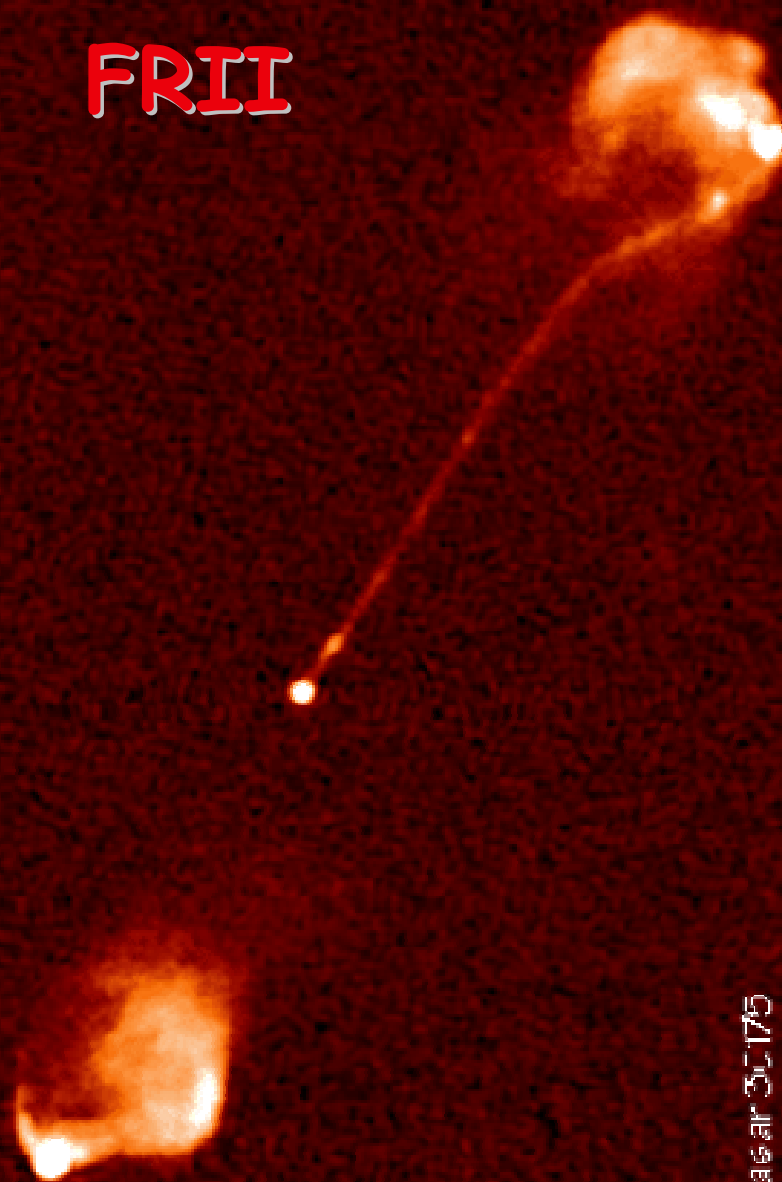
Radio Galaxy Zoo II
VLA 20cm image

FRI

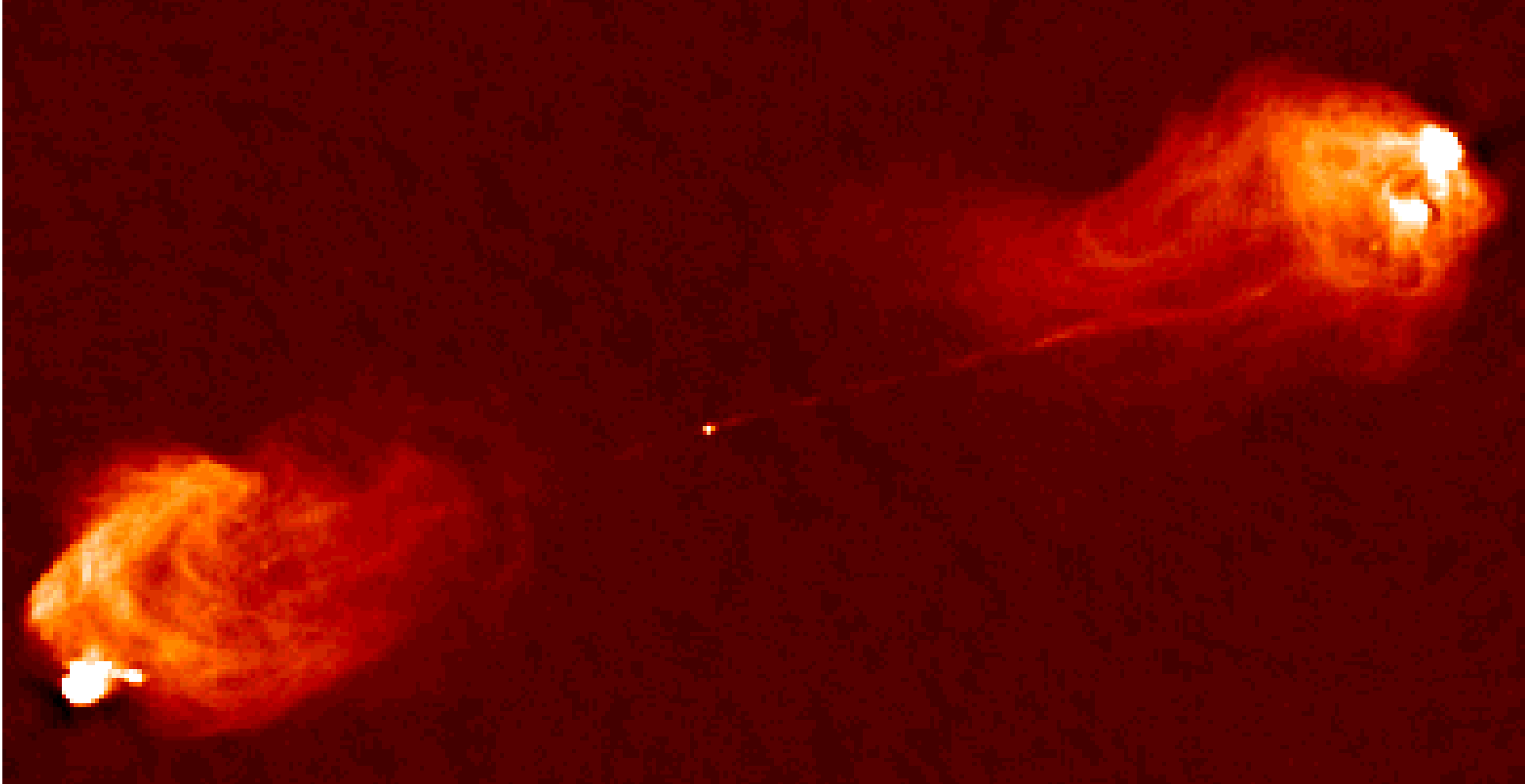


Copyright (c) NRAO/AUI 2009

FRII



Quasar 3C 175
VLA 6cm image (c) NRAO 1986



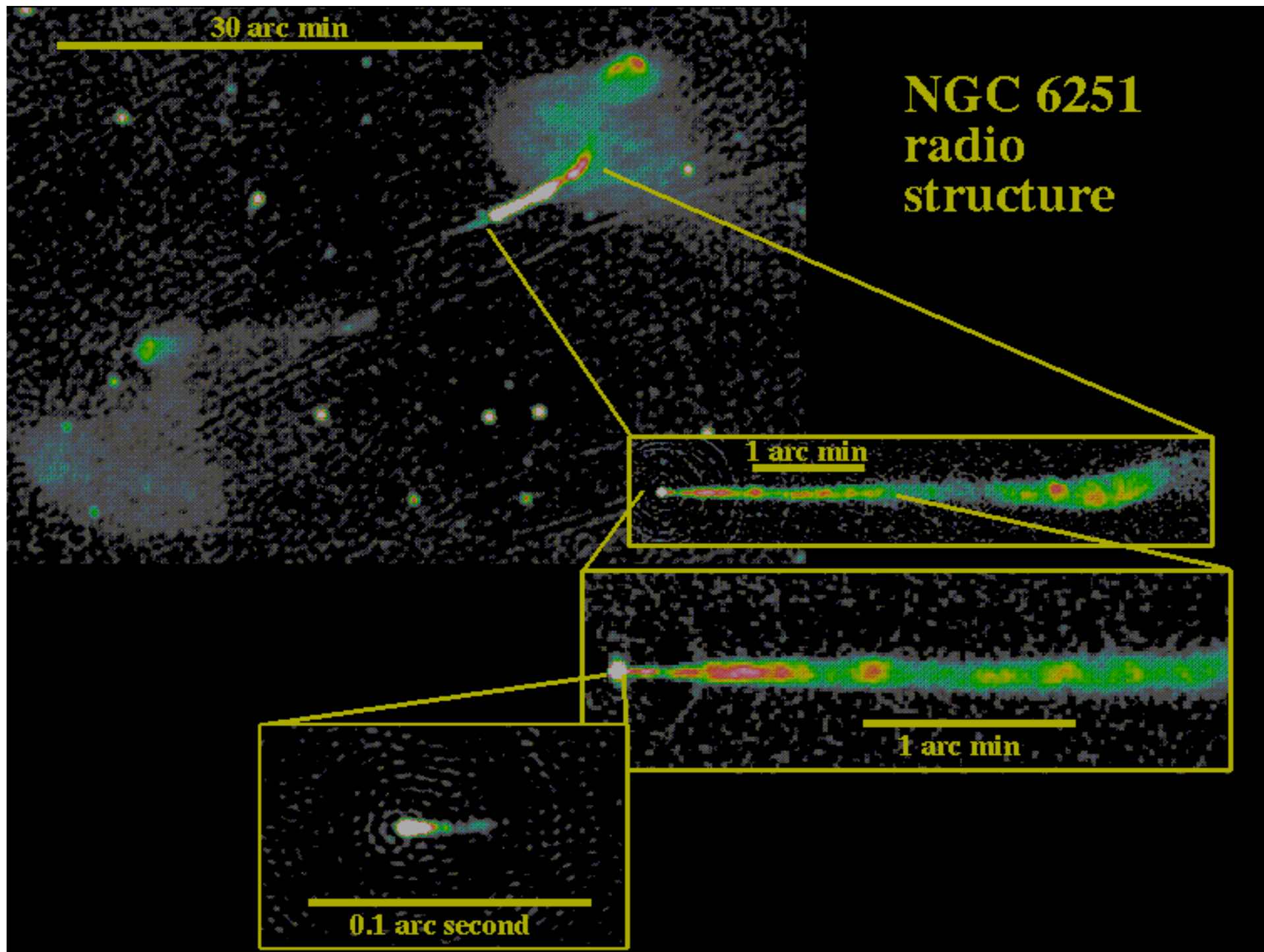
30 arc min

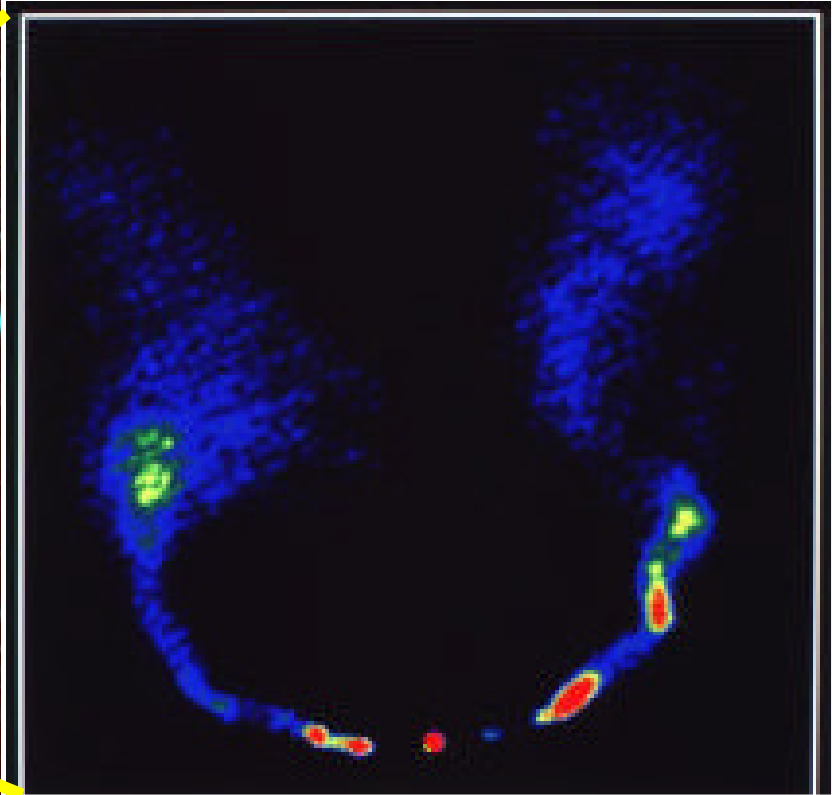
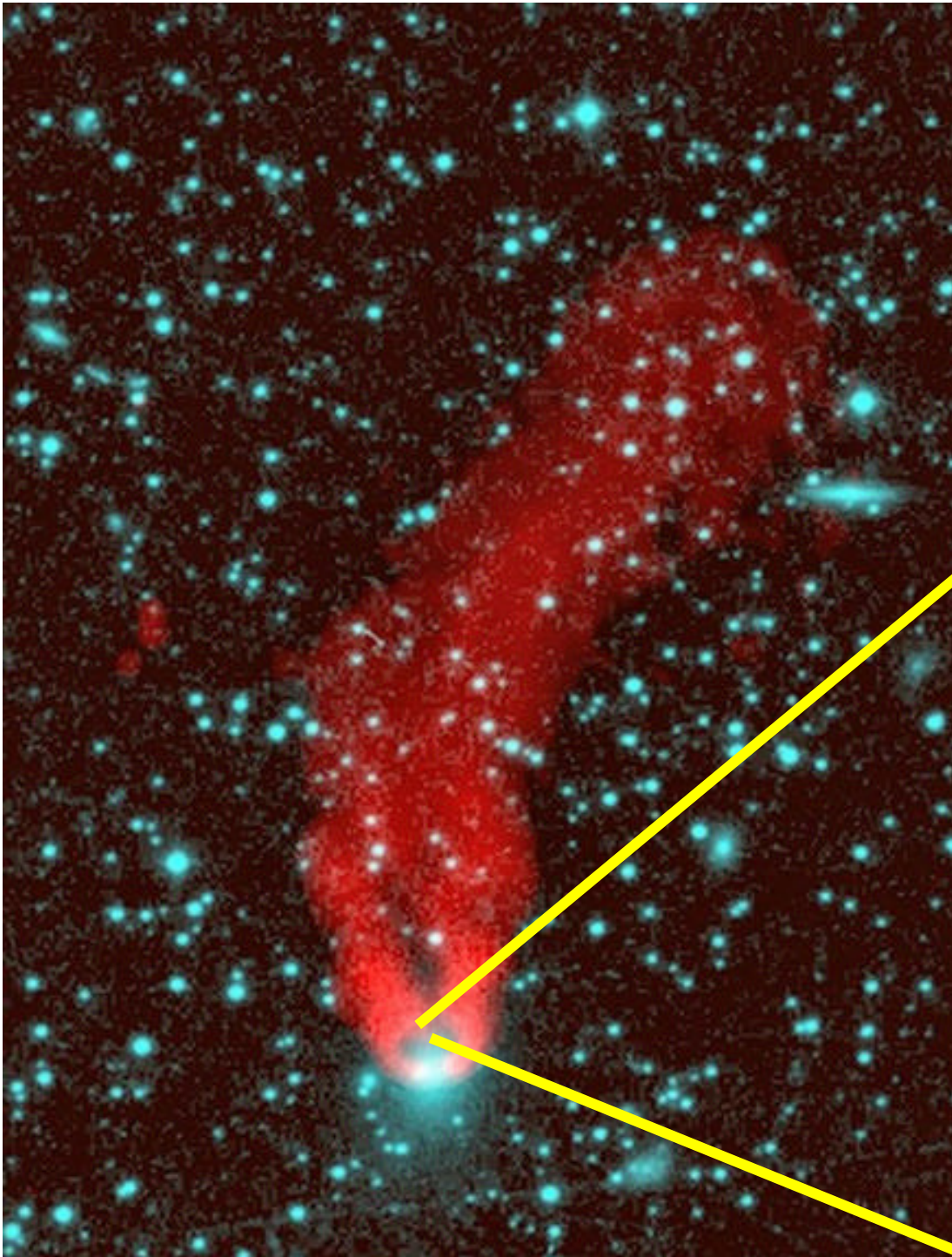
NGC 6251
radio
structure

1 arc min

1 arc min

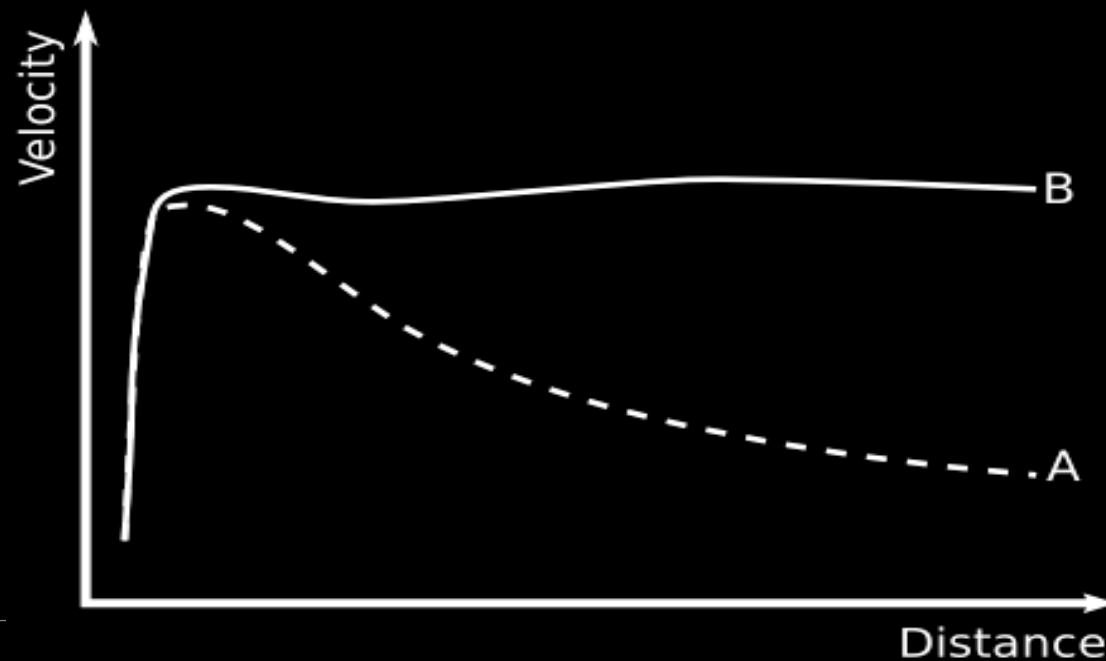
0.1 arc second



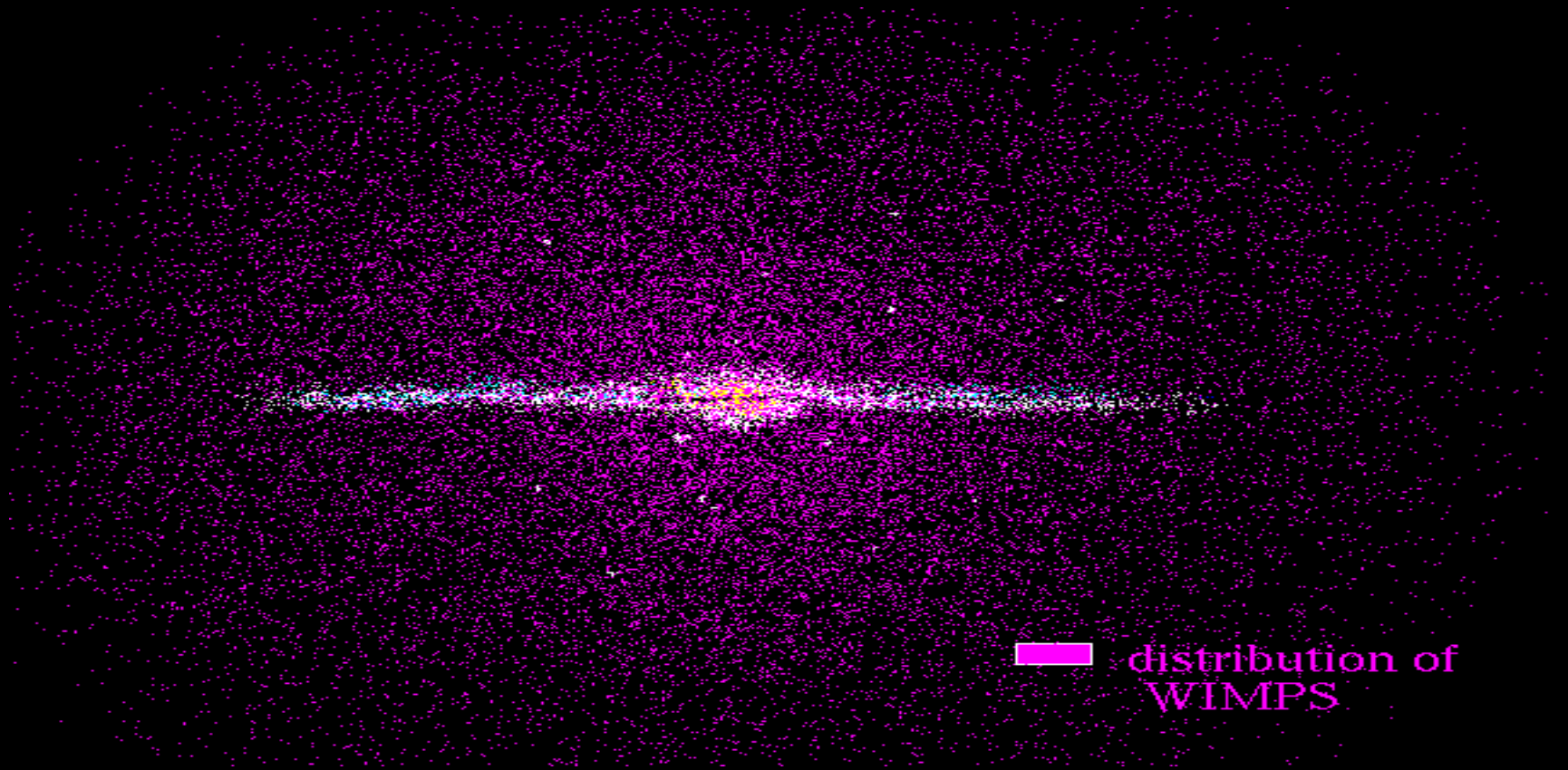


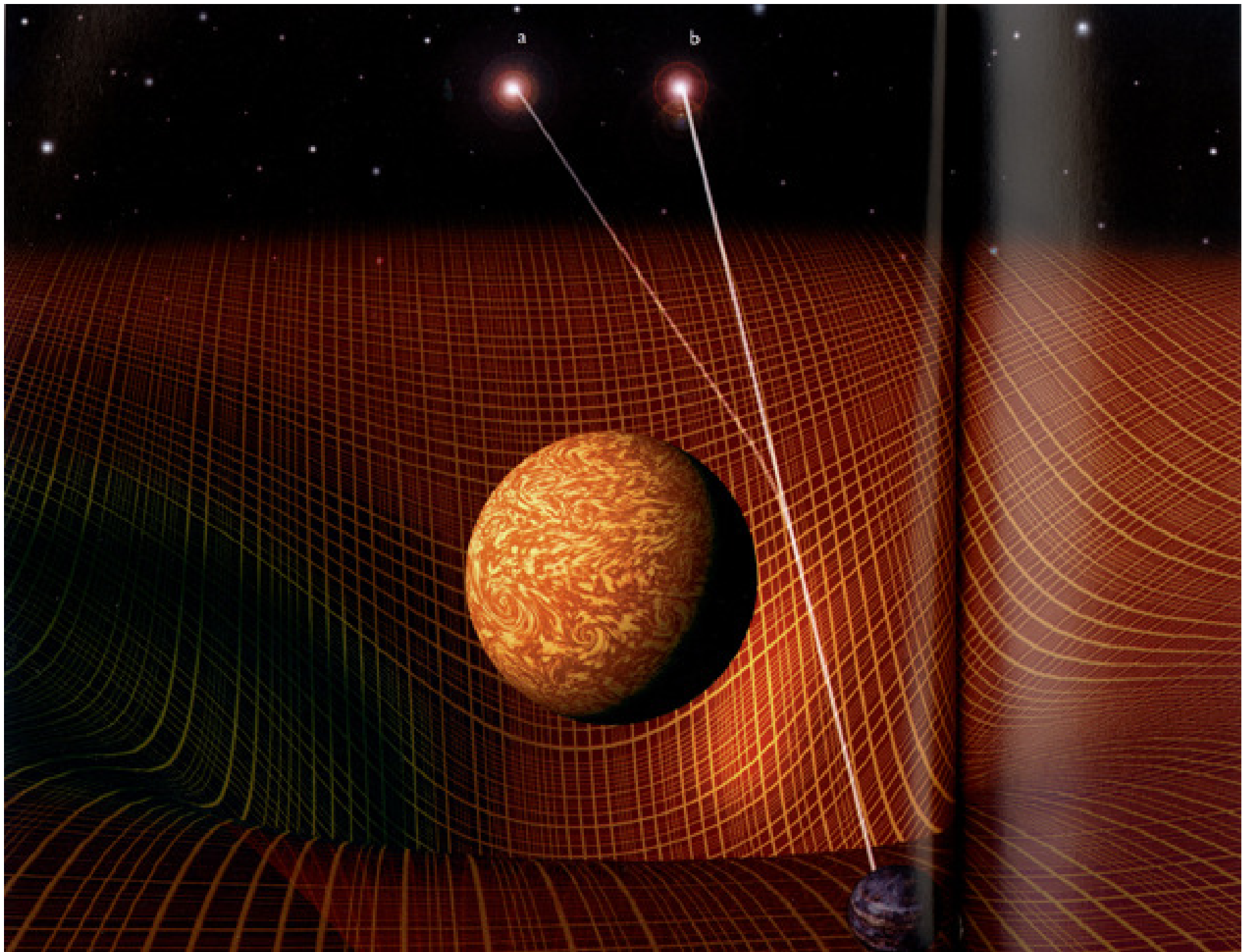


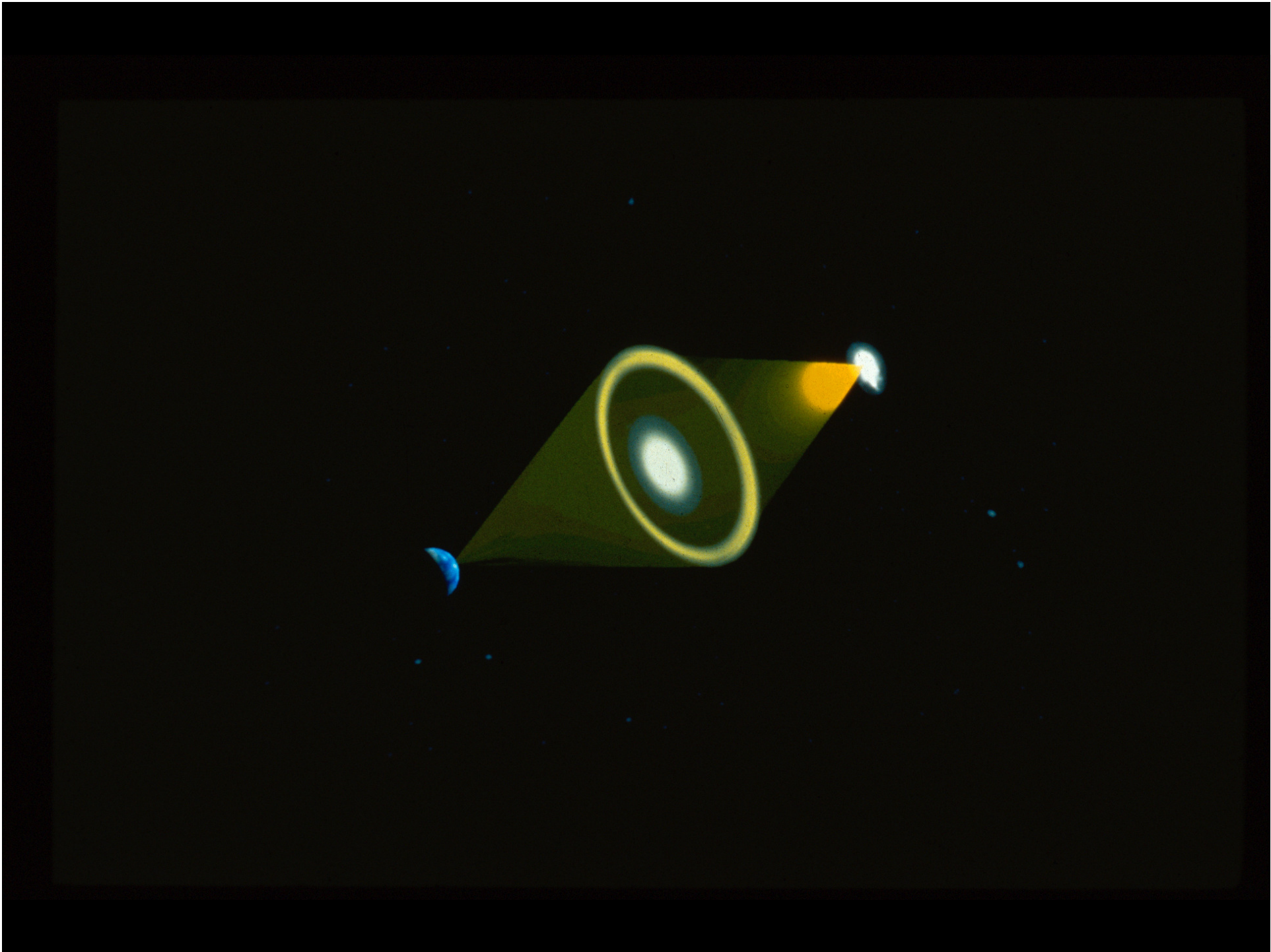
Curve di rotazione delle Galassie a Spirale



La materia scura





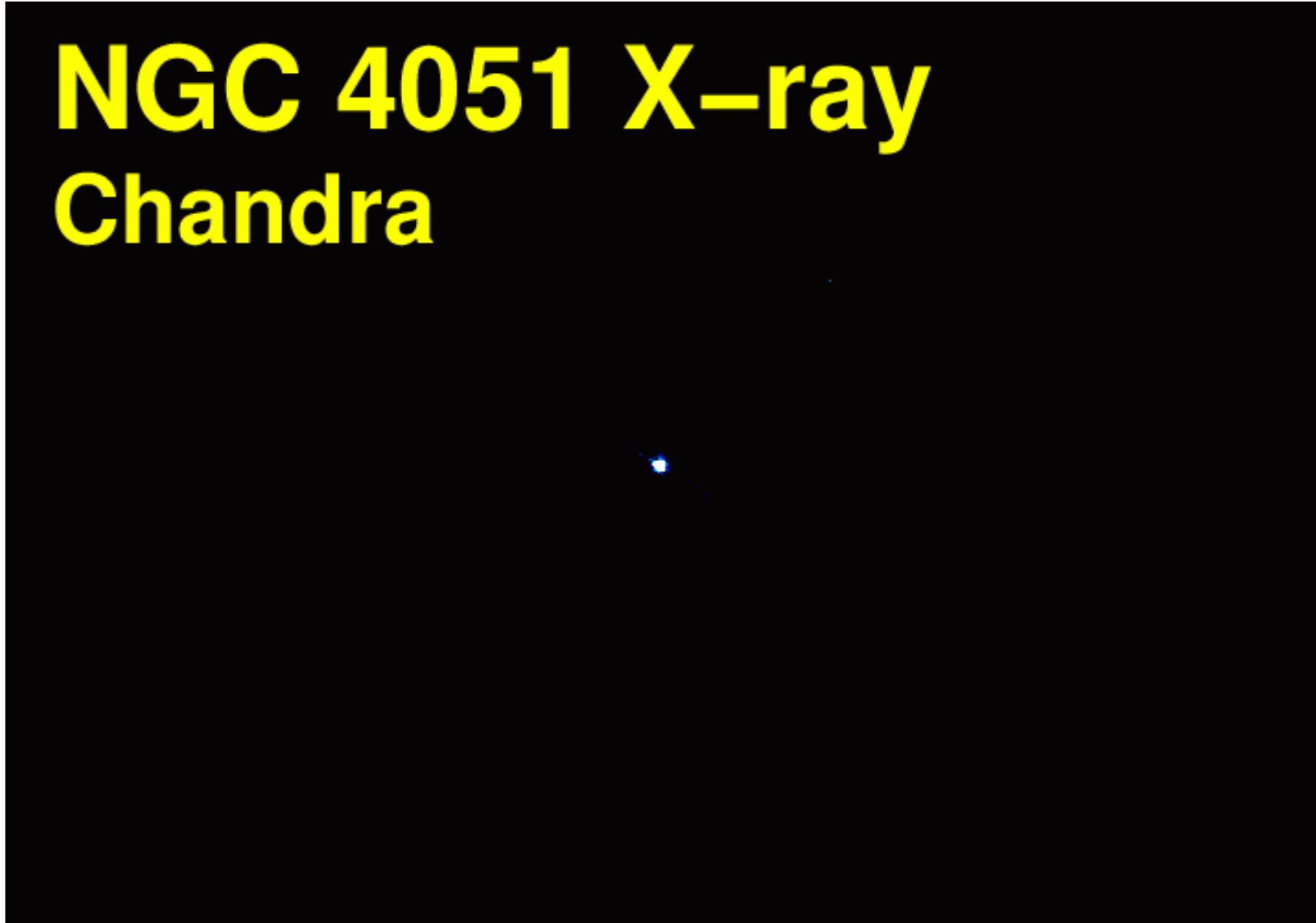


**~90% della massa
dell'Universo è di un qualche tipo
di materia scura (dark),
probabilmente di natura diversa
da quella di cui siamo fatti noi.**

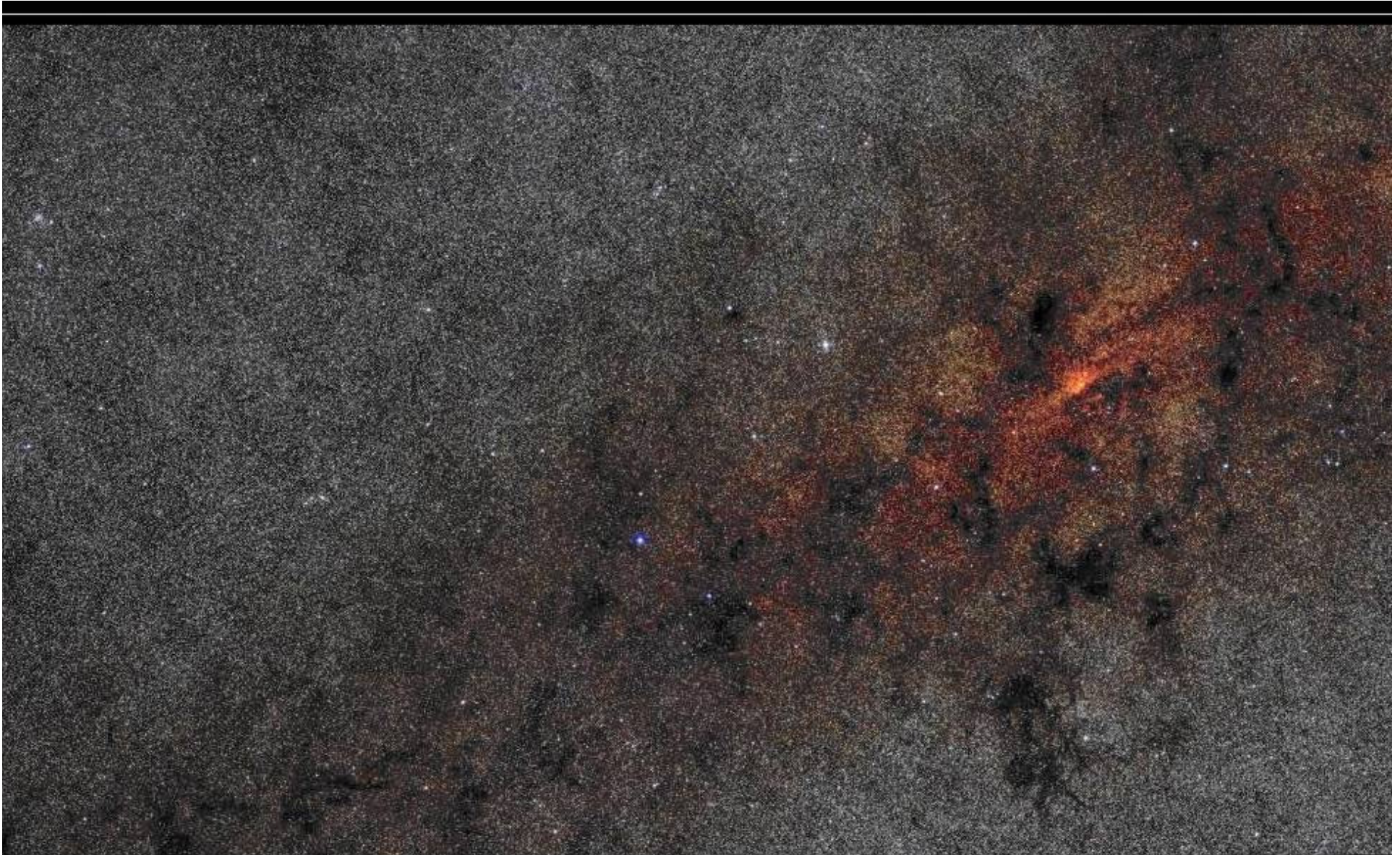
**90% materia scura ?
questi astronomi devono essere
pazzi !**

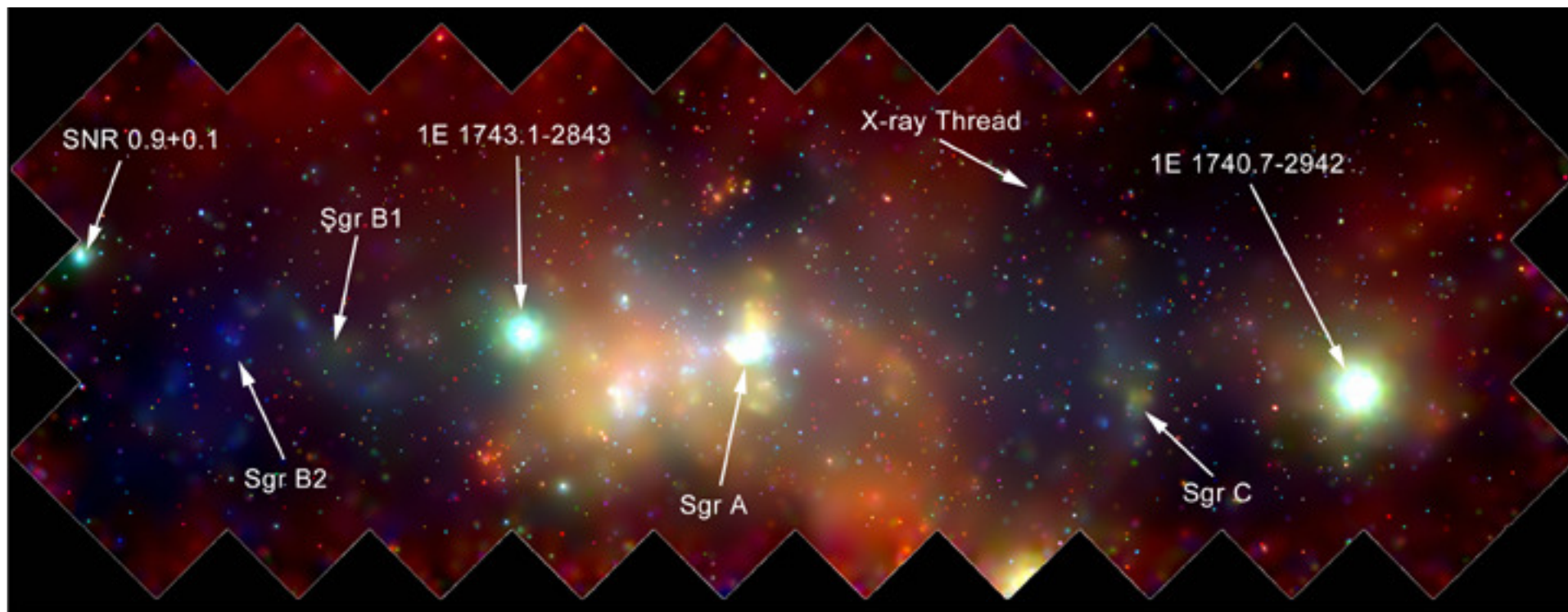
Emissione X da un AGN

NGC 4051 X-ray
Chandra



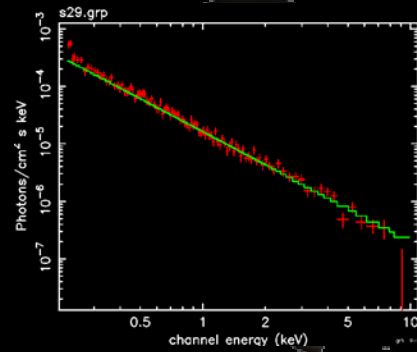
Il centro della nostra Galassia



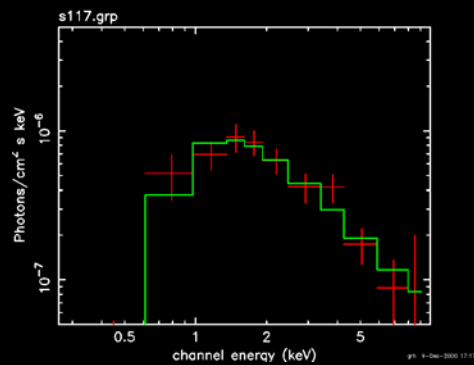
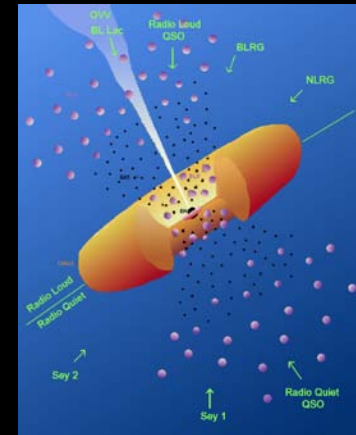


Il centro della Galassia in raggi-X (Chandra)

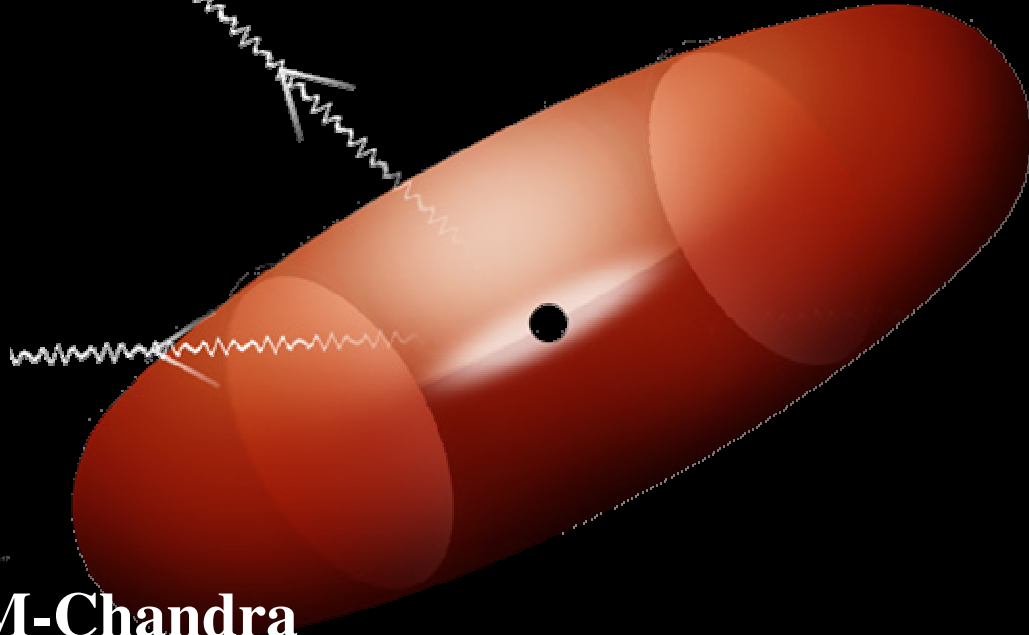
Spettri X e il Modello Unificato degli AGN



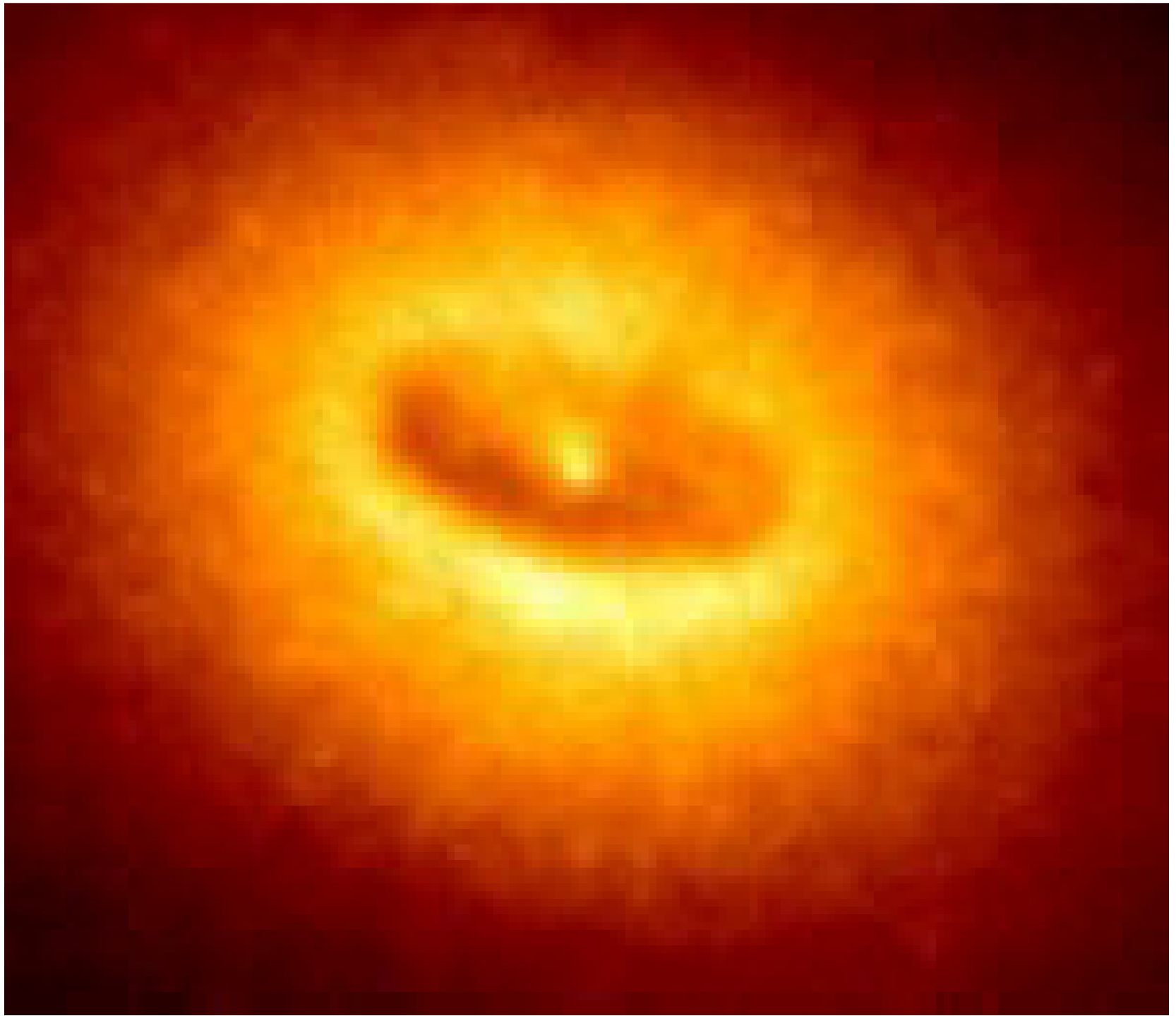
Einstein - ROSAT
E < 3 keV

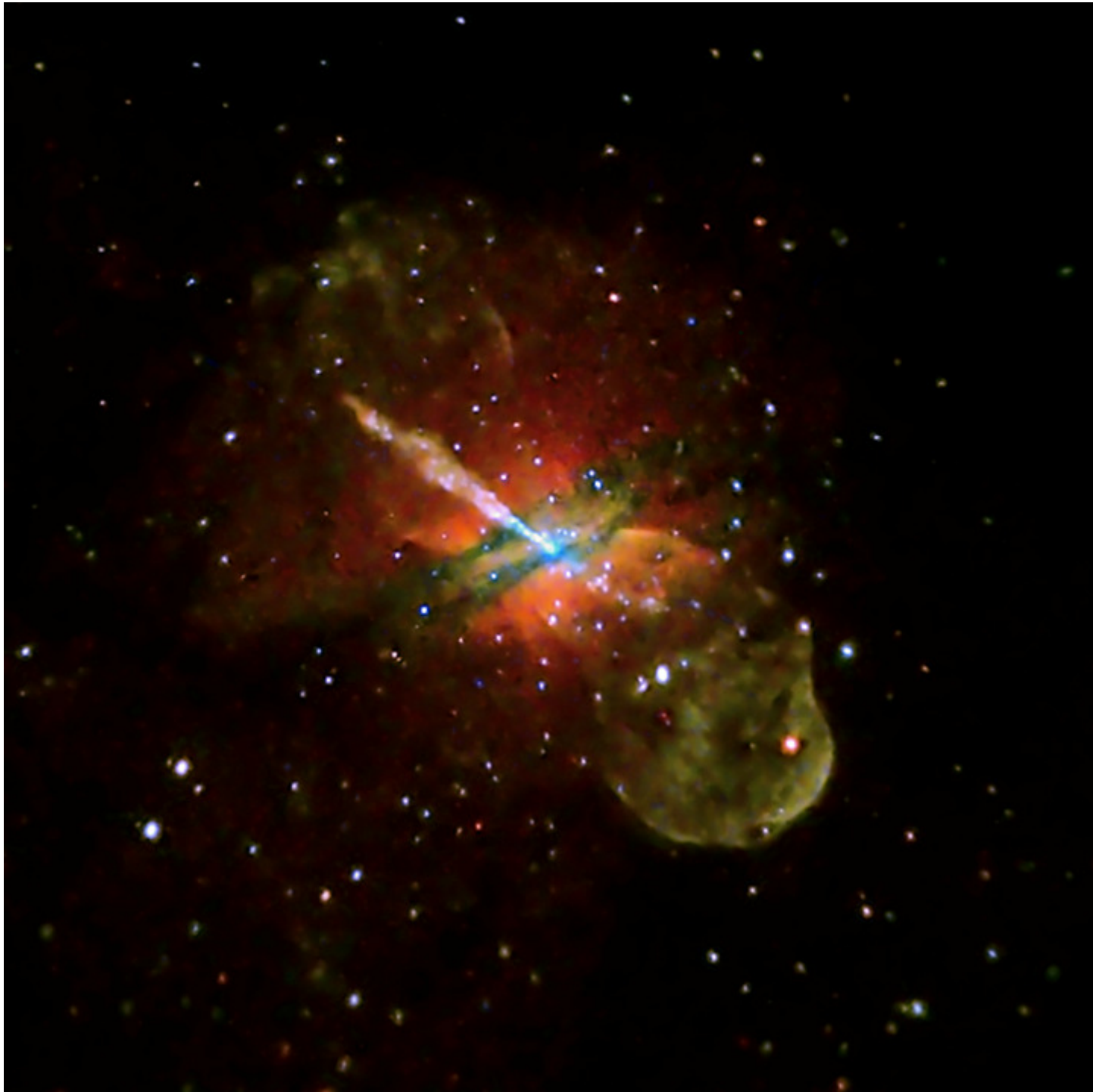


ASCA-SAX-XMM-Chandra
E > 3 keV



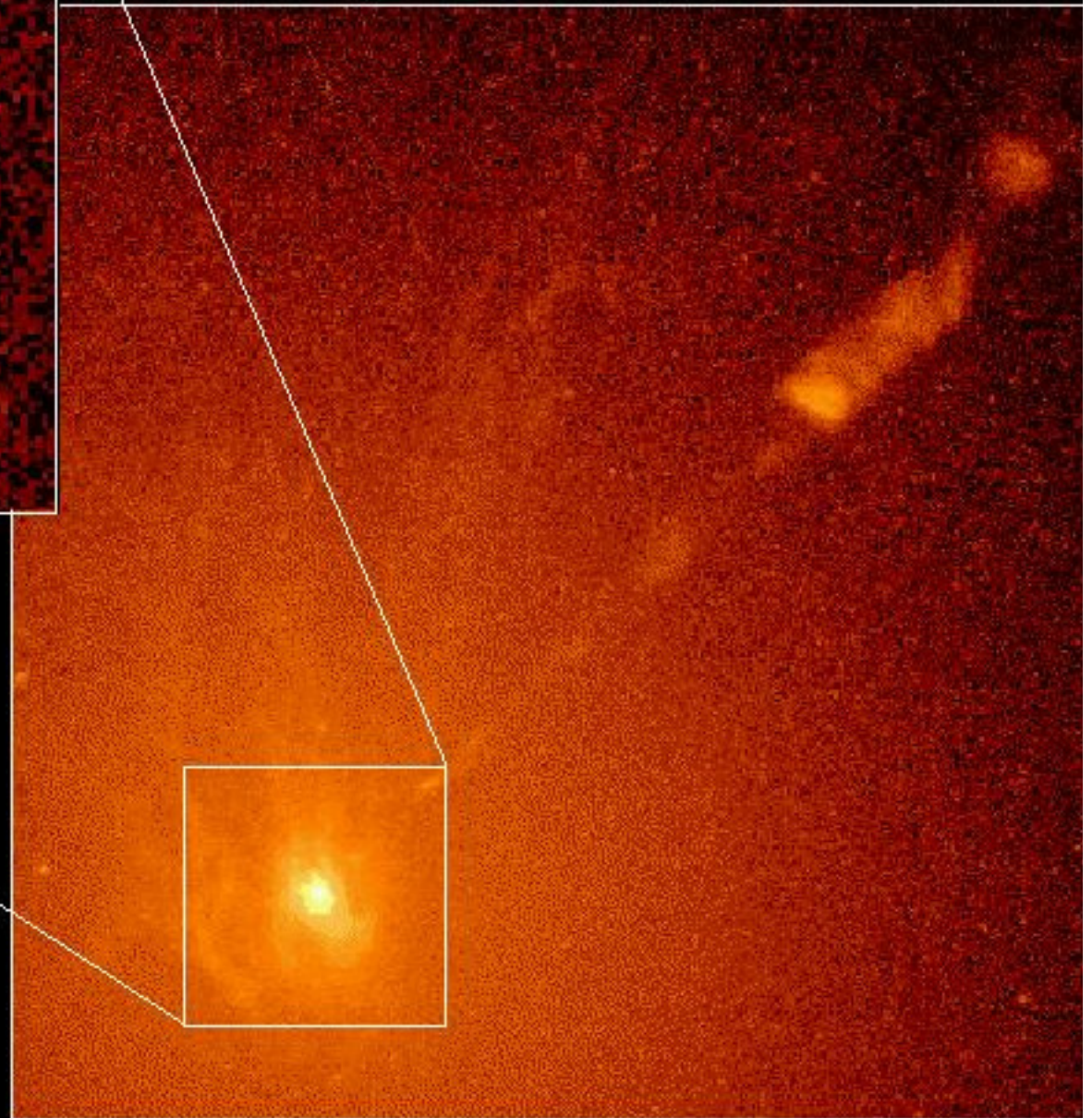
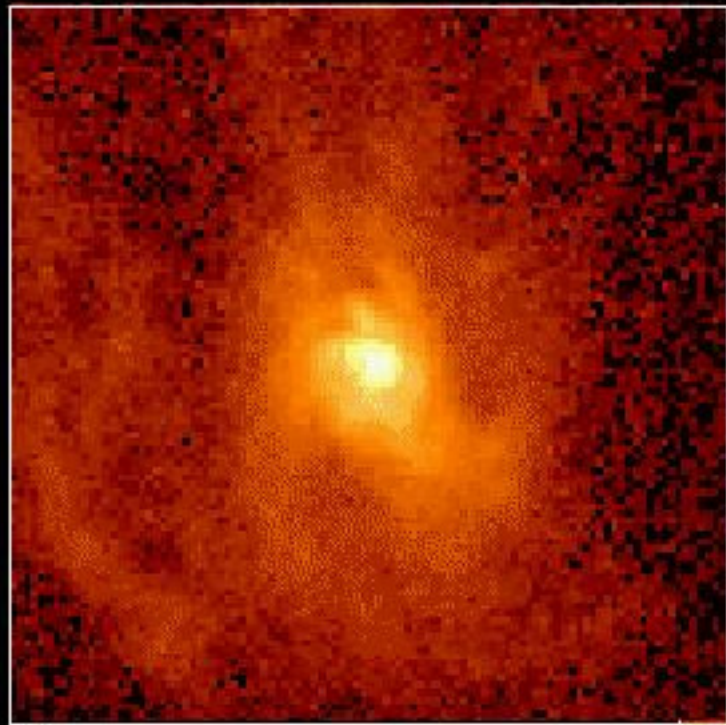
Adapted from G. Hasinger





CEN A

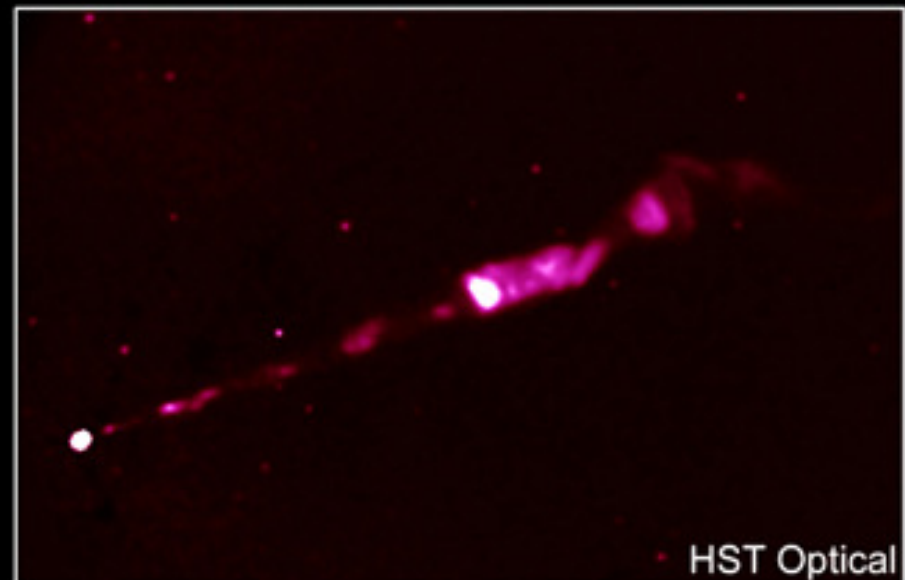
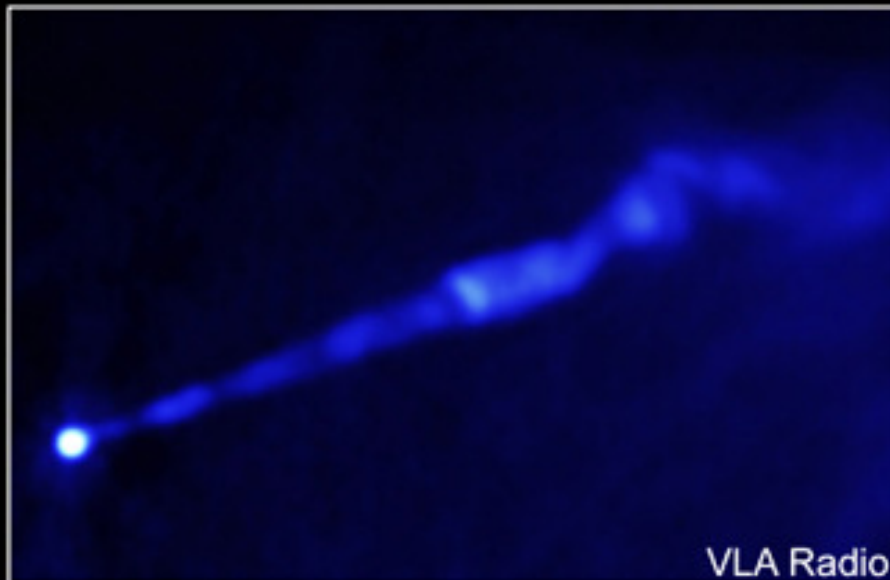
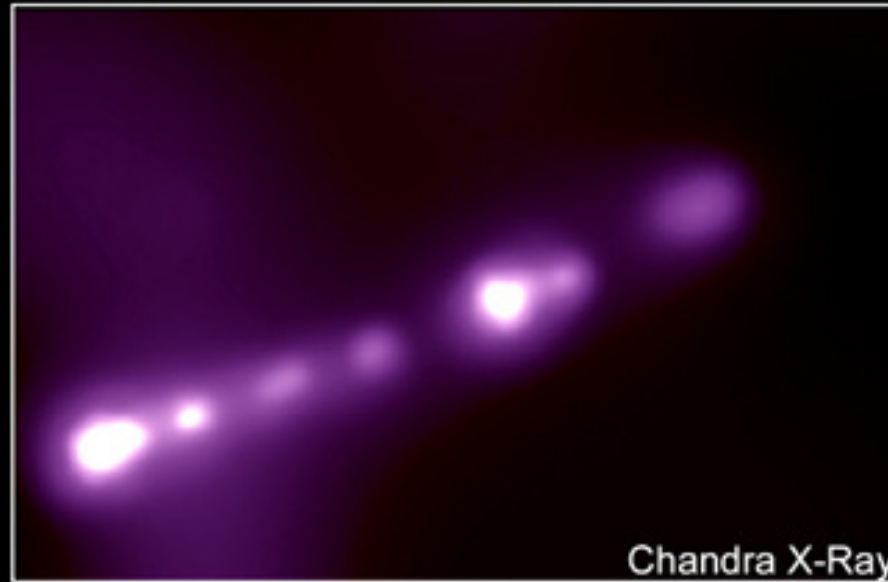
Gas Disk in Nucleus of Active Galaxy M87

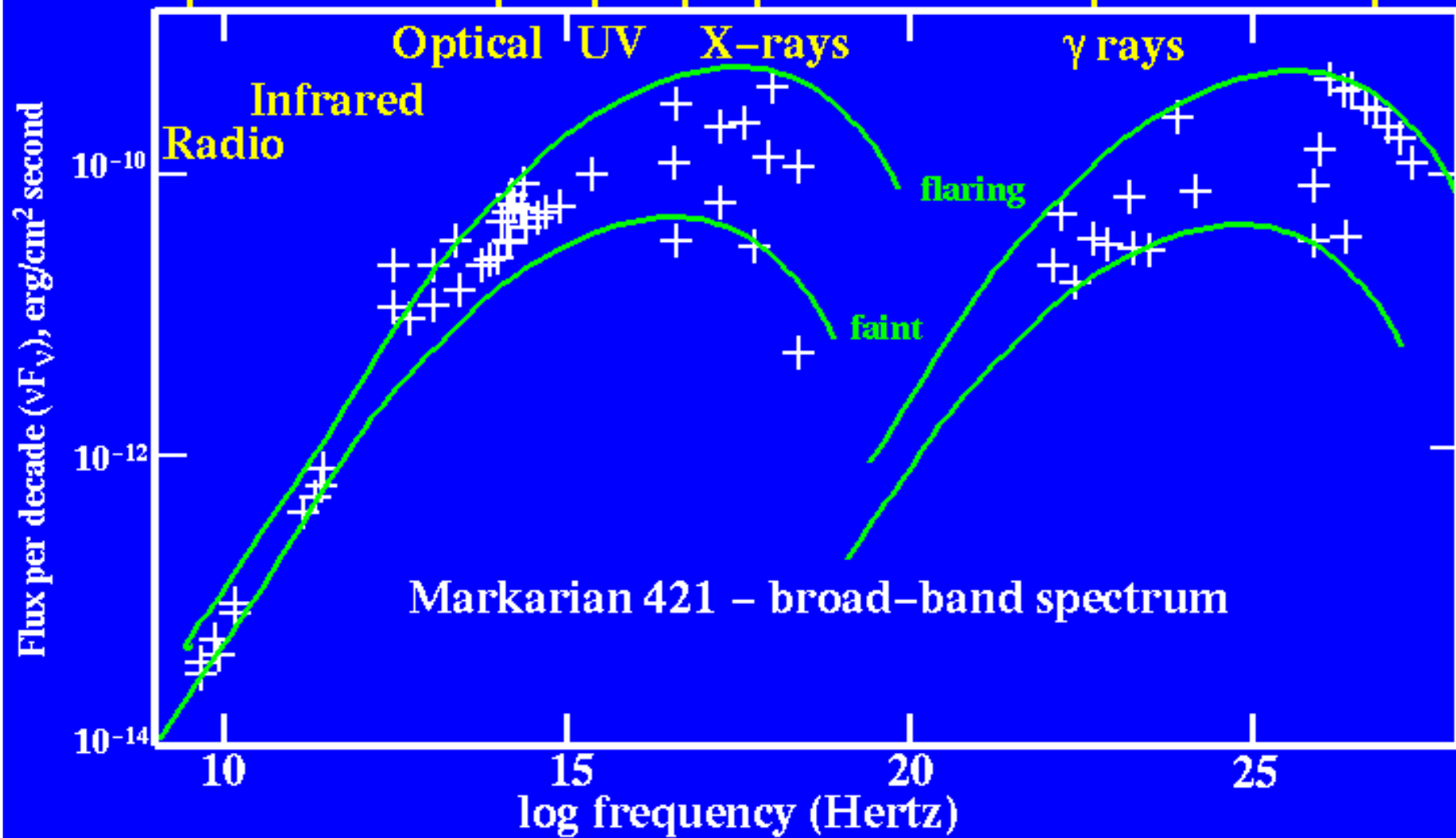
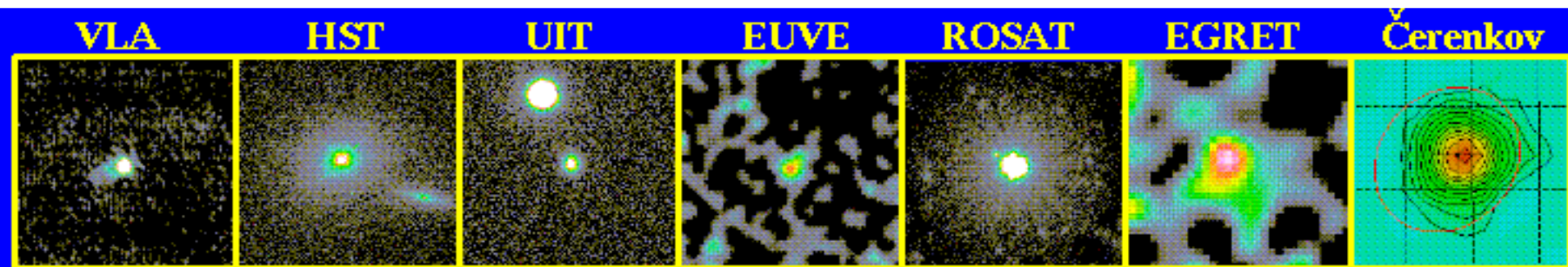


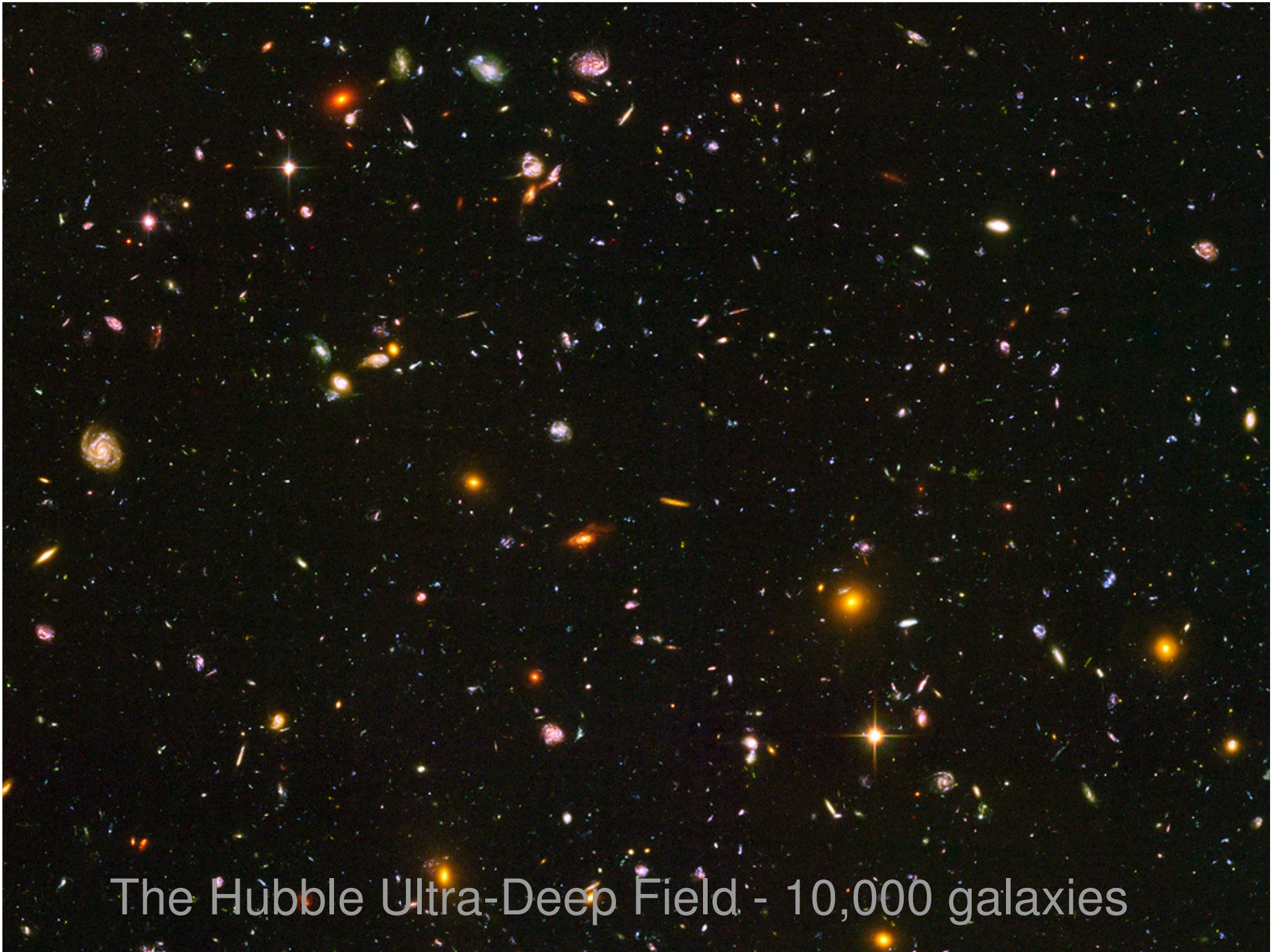
Hubble Space Telescope
Wide Field Planetary Camera 2

Getti in raggi-O/X/R

M87

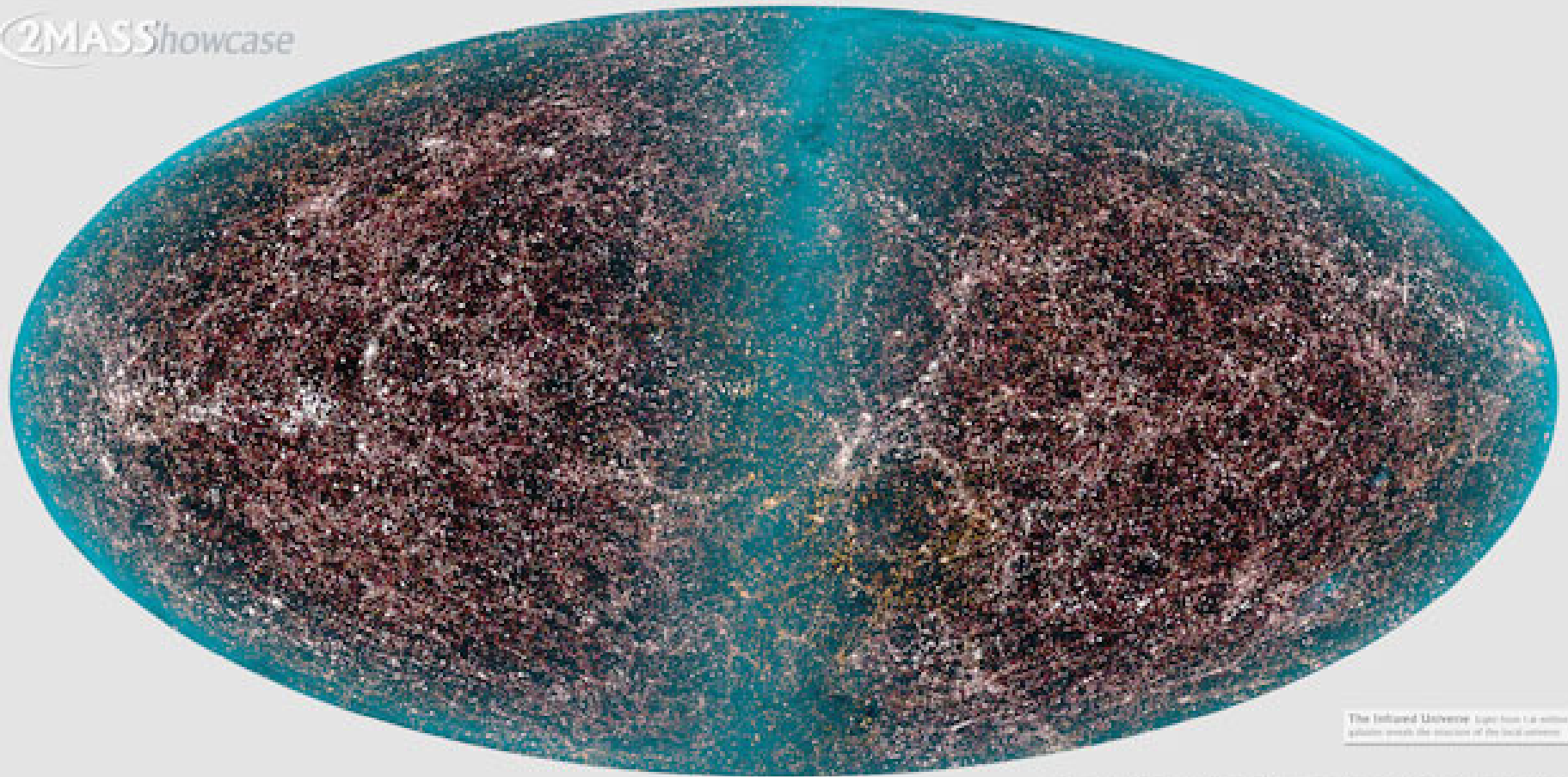






The Hubble Ultra-Deep Field - 10,000 galaxies

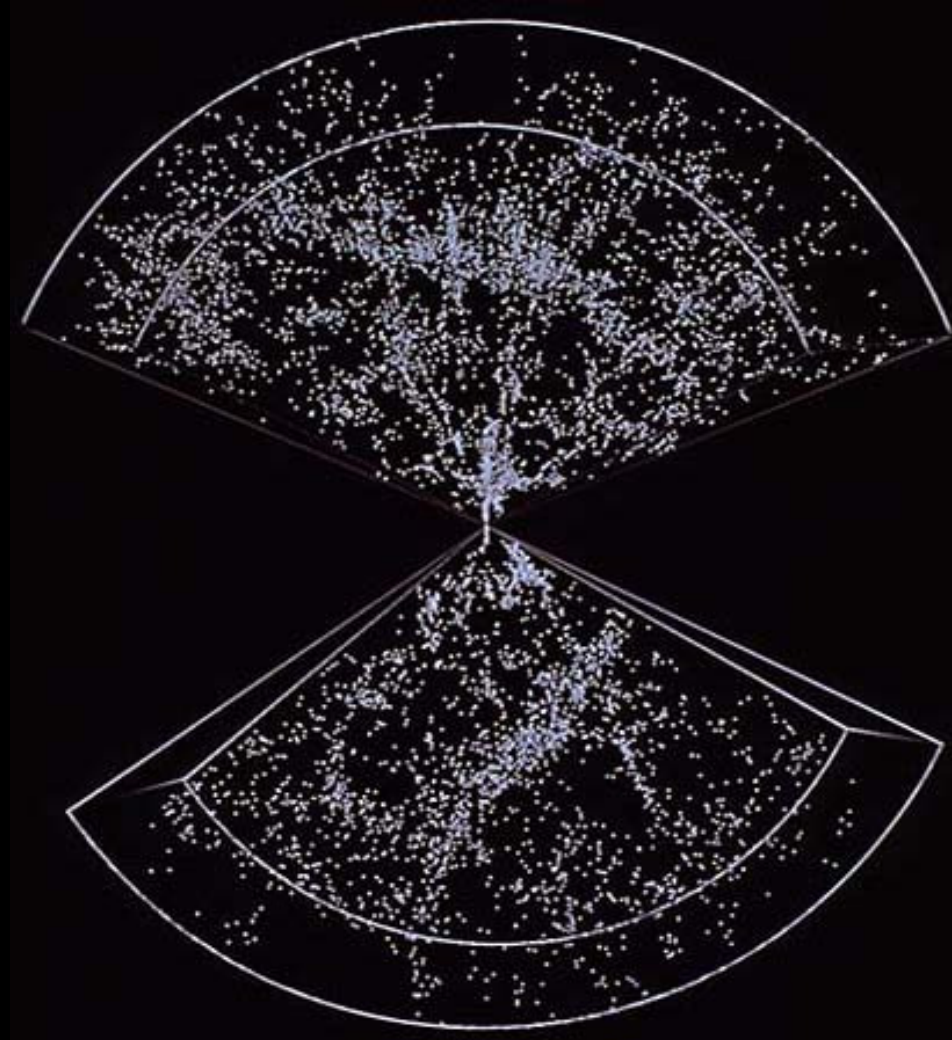
2MASS Showcase



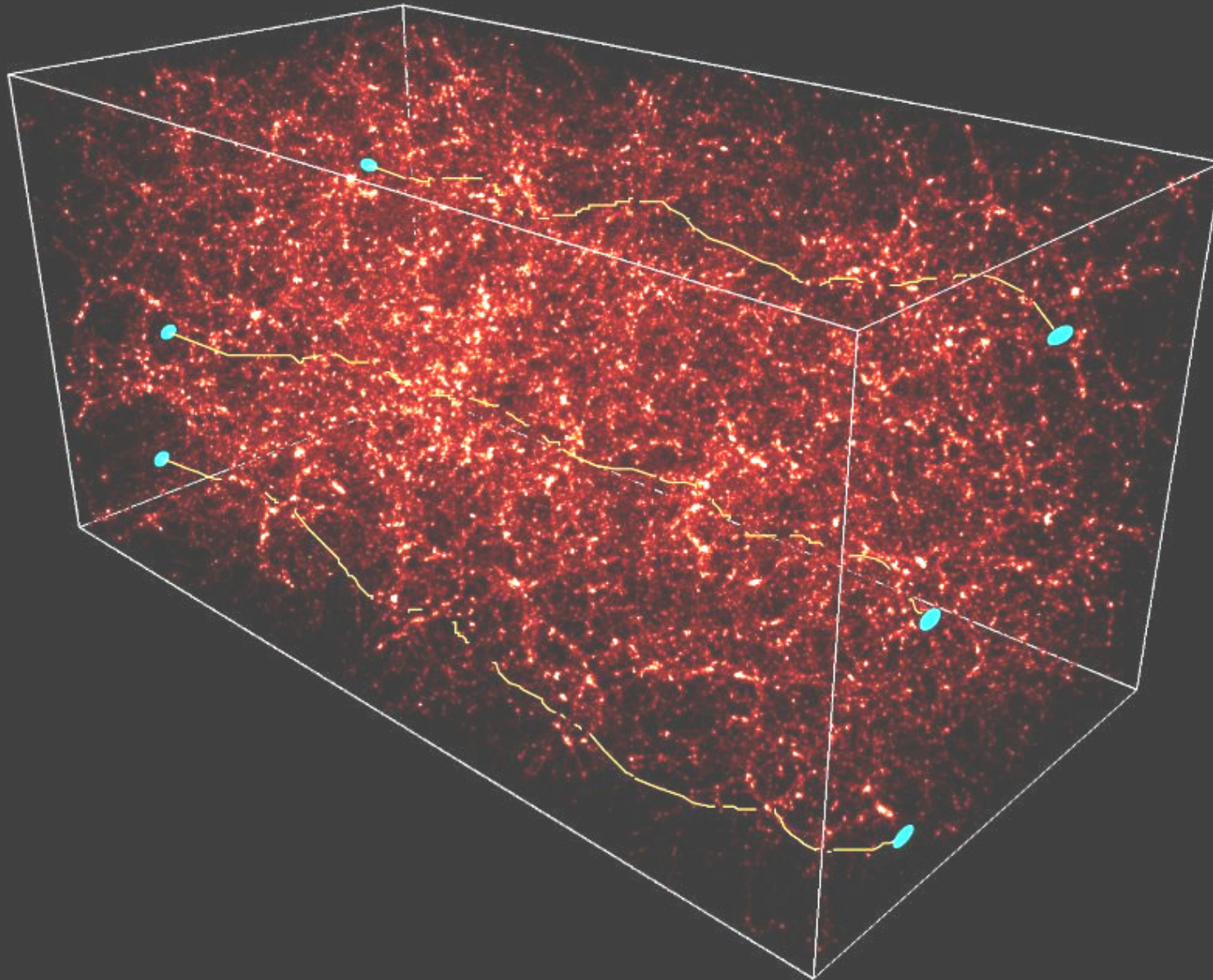
The Infrared Universe: Light from 1.4 billion galaxies reveals the structure of the local universe.

The World of Sky (Image: Bruce; infrared Processing and Analysis Center/CITRIS & University of Massachusetts)

Le grandi strutture



DEFLECTION OF LIGHT RAYS CROSSING THE UNIVERSE, EMITTED BY DISTANT GALAXIES



Supernovae lontane non sono
facilmente distinte dalle galassie
che le ospitano

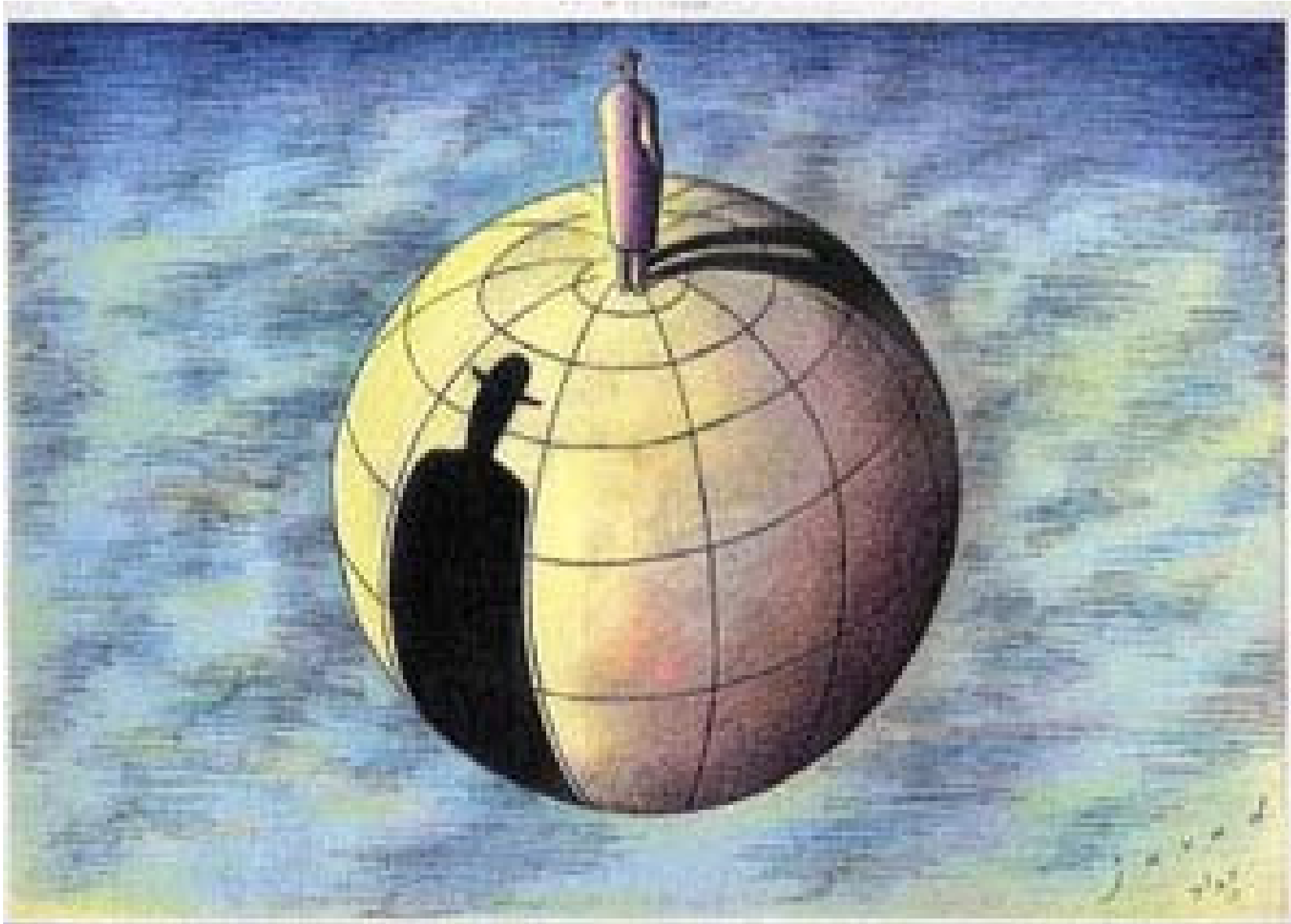


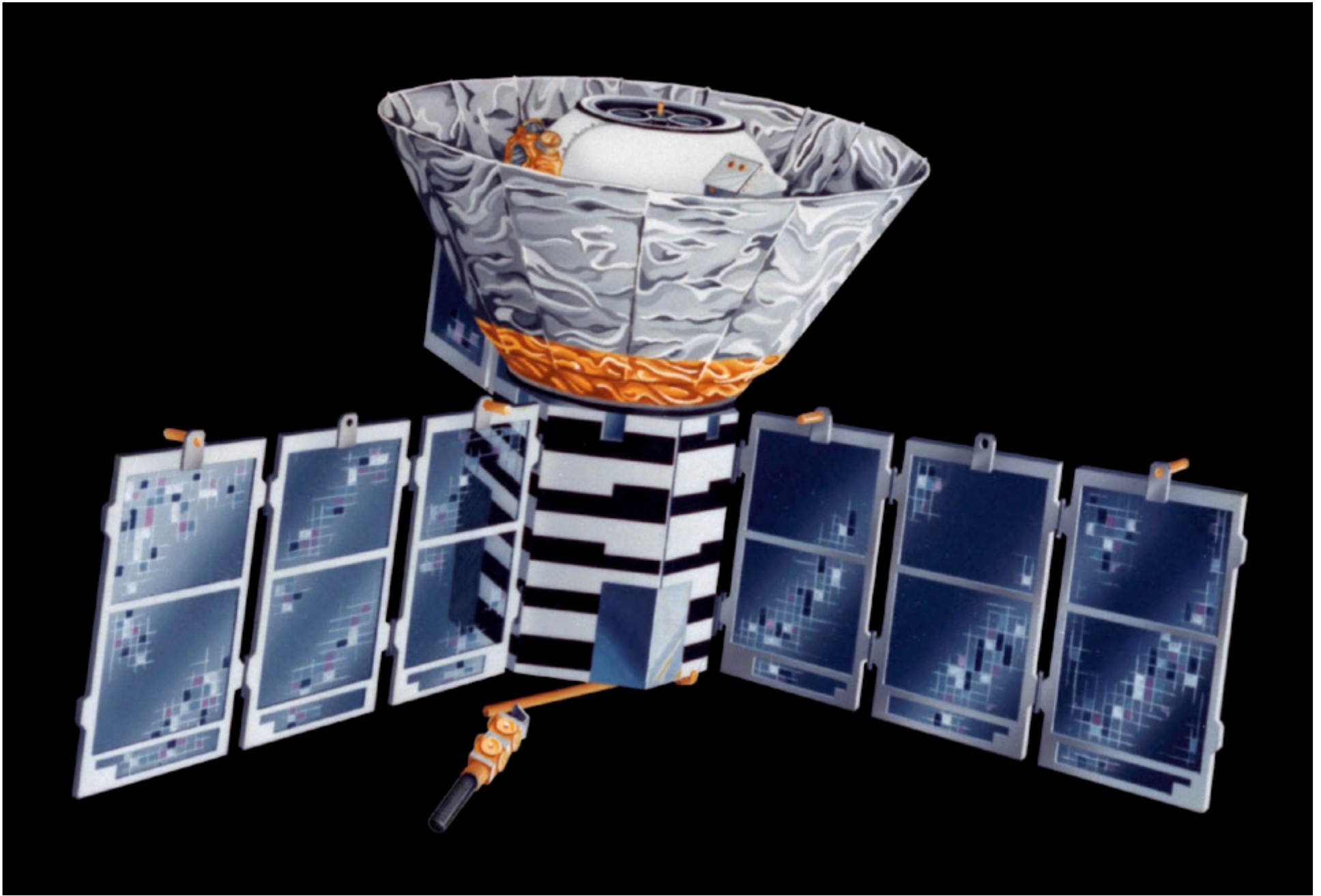
Supernovae le "candele standard"



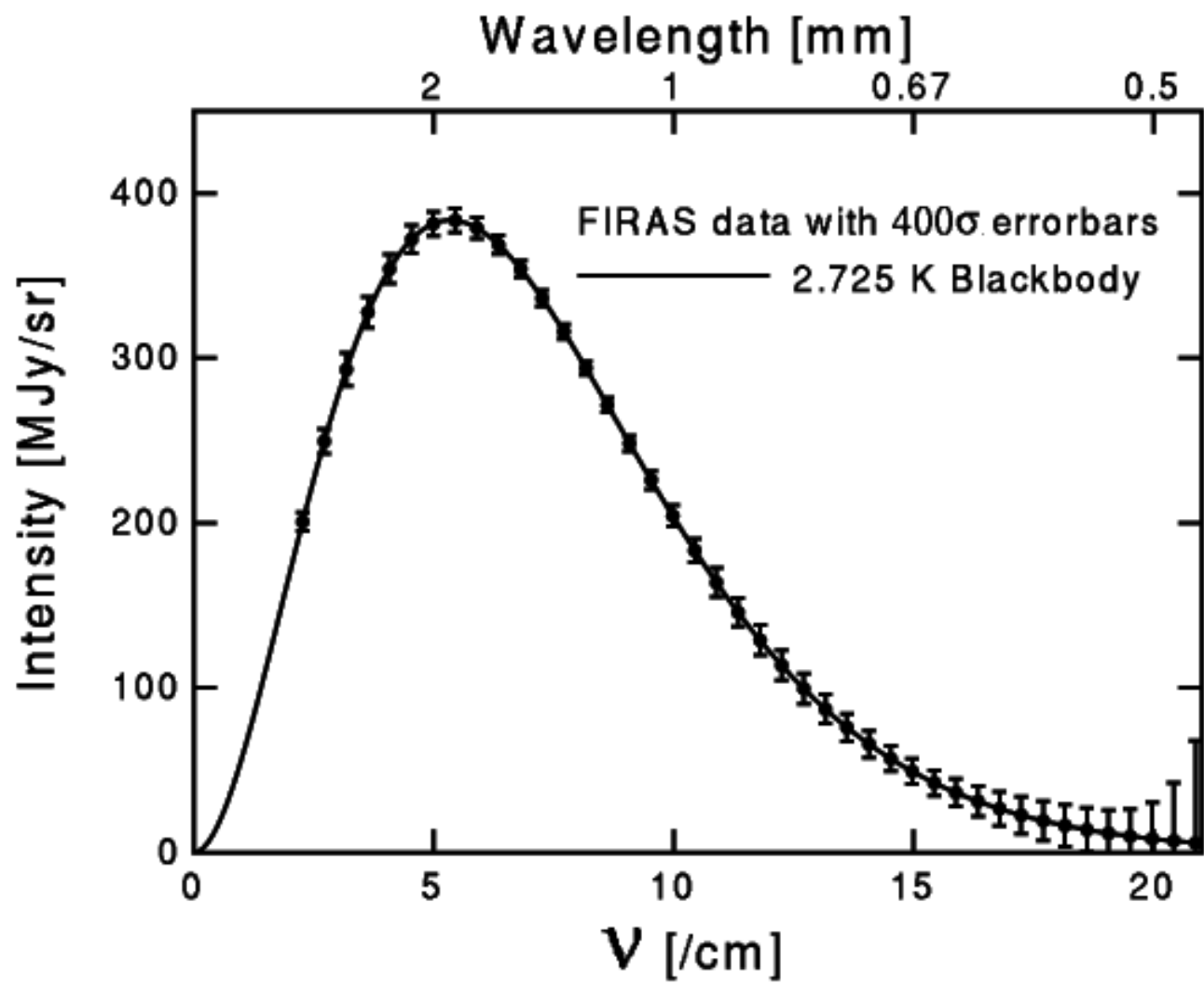
-Da misure di
velocità e distanza
si può calcolare il
tempo....e graficare
la Storia
dell'Universo

In che UNIVERSO viviamo?





COBE



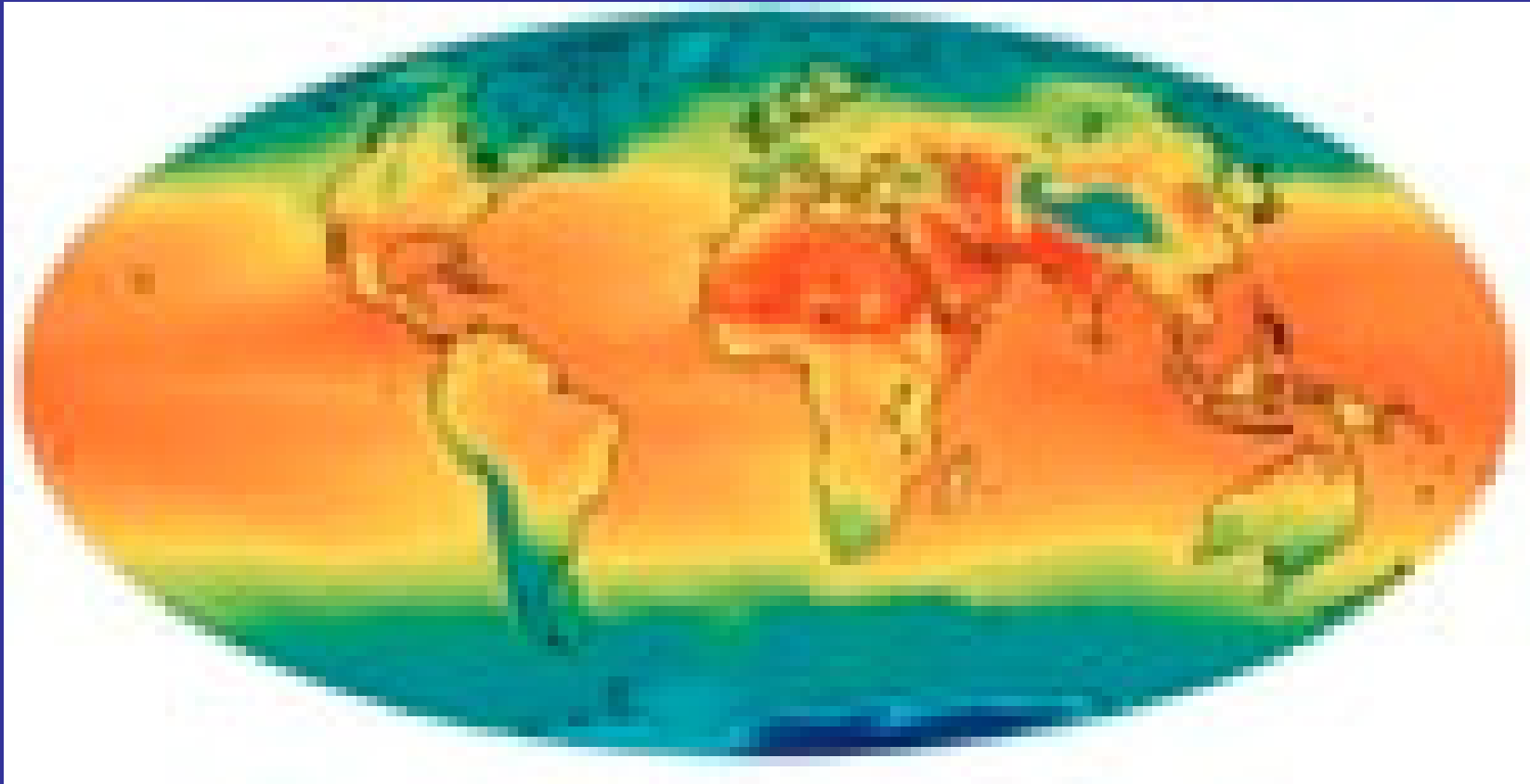
La temperatura dell'Universo

$T = 2,725 \text{ K}$ temperatura della "luce" che ha viaggiato verso di noi per il 90% del tempo dal Big Bang a oggi.

Sono pochi gradi ma una sorprendente quantità di calore.

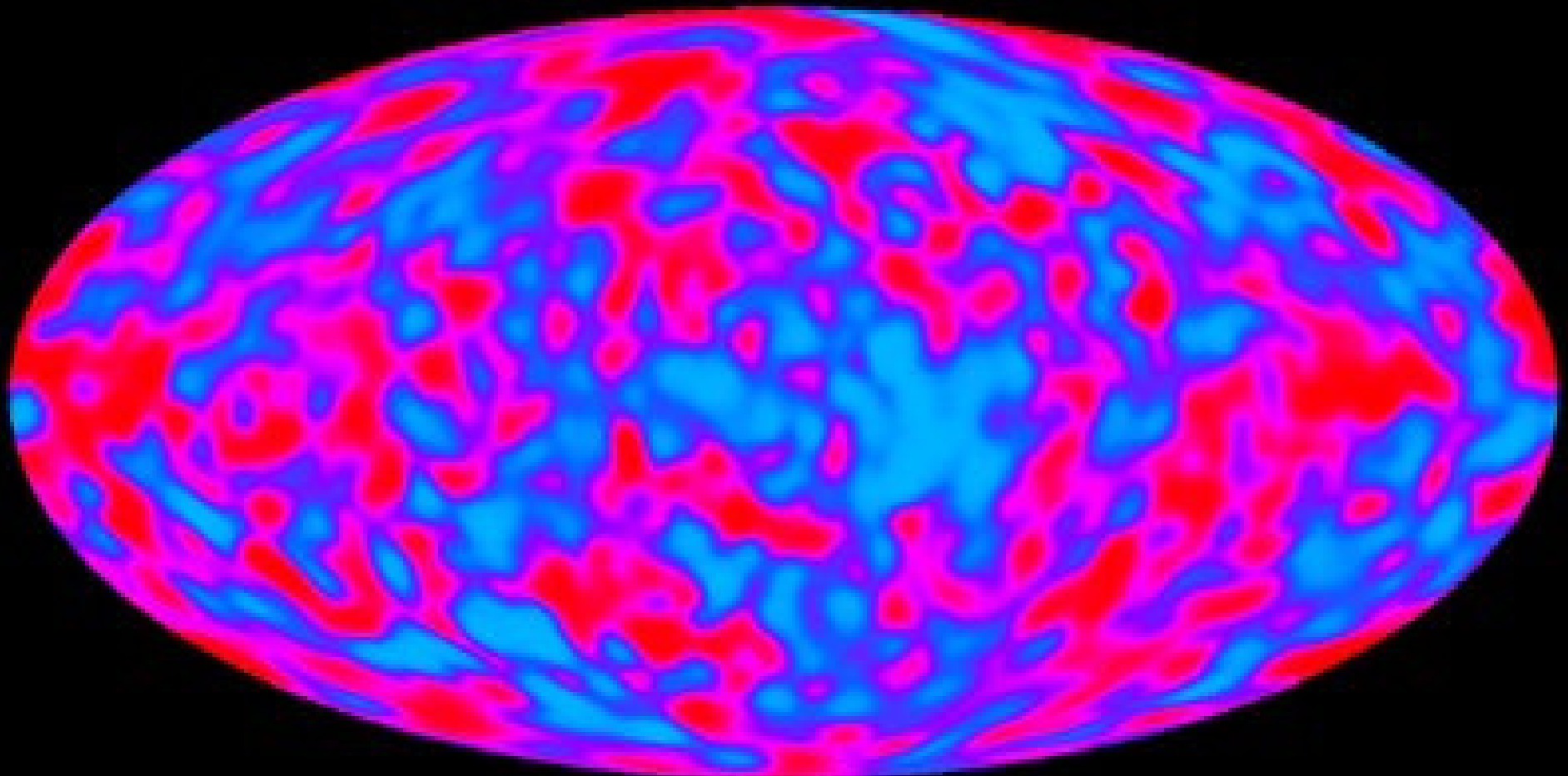
2 miliardi di fotoni per ogni atomo

La temperatura della Terra

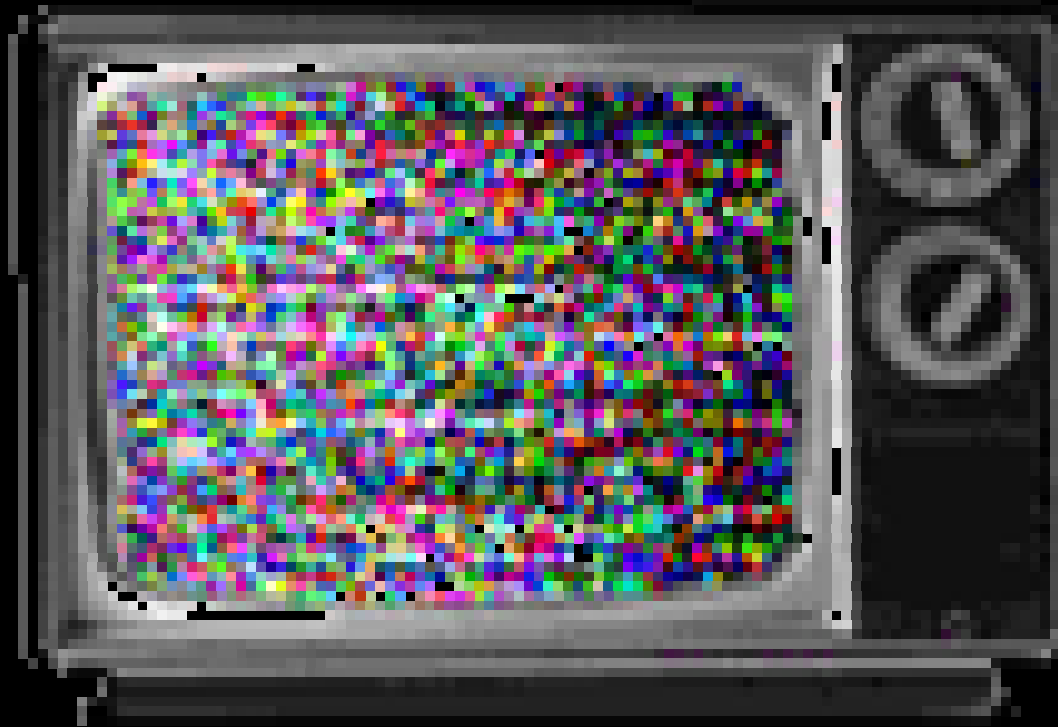


Fluttuazioni della radiazione cosmica di fondo

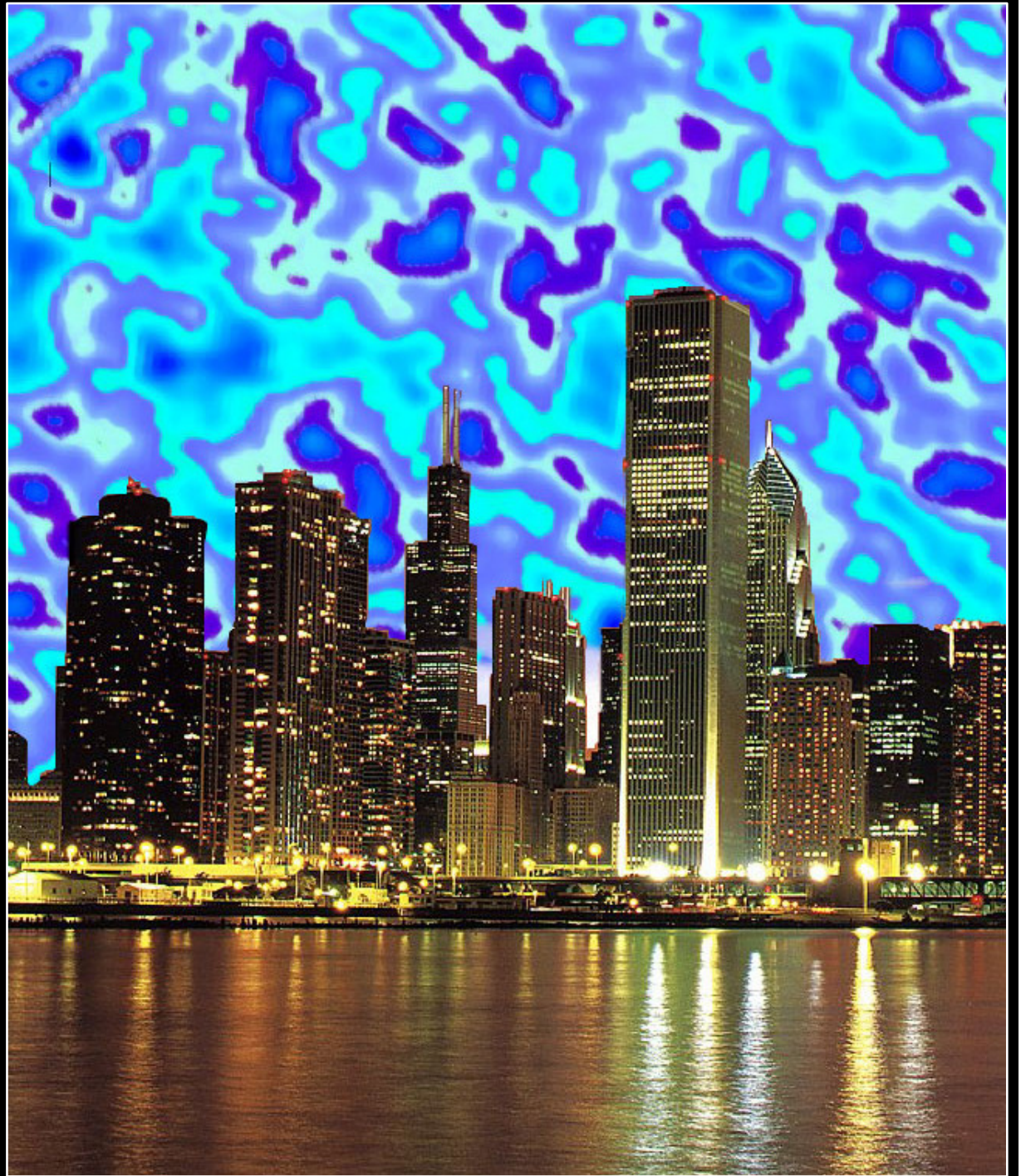
Nella radiazione cosmica di fondo vediamo le tracce delle fluttuazioni di densità che hanno dato origine alle strutture che osserviamo oggi.



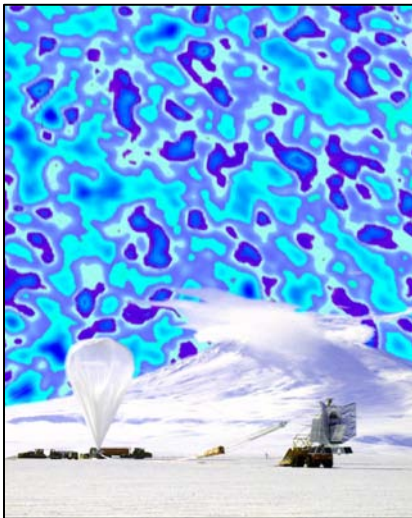
- 400 fotoni per cm cubo !



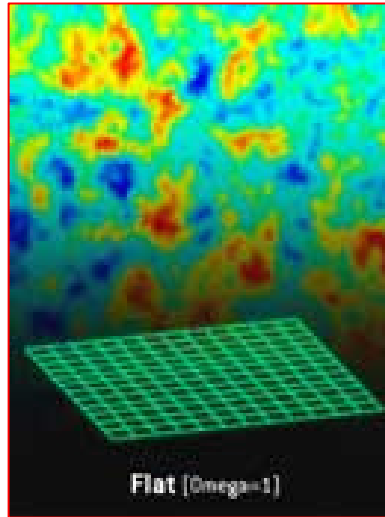
Se i nostri
occhi
vedessero le
microonde,
il cielo ci
apparirebbe
così



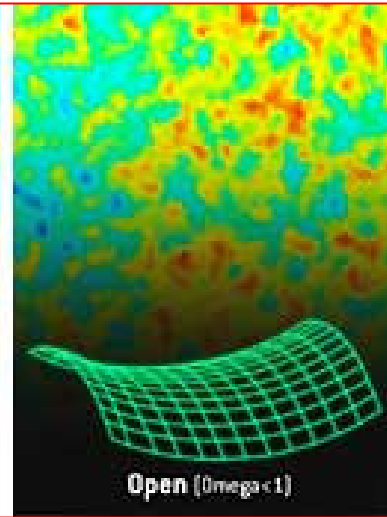
La geometria dell'universo



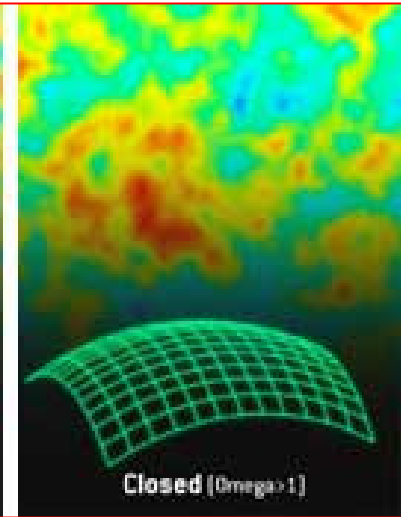
Boomerang



Piatto

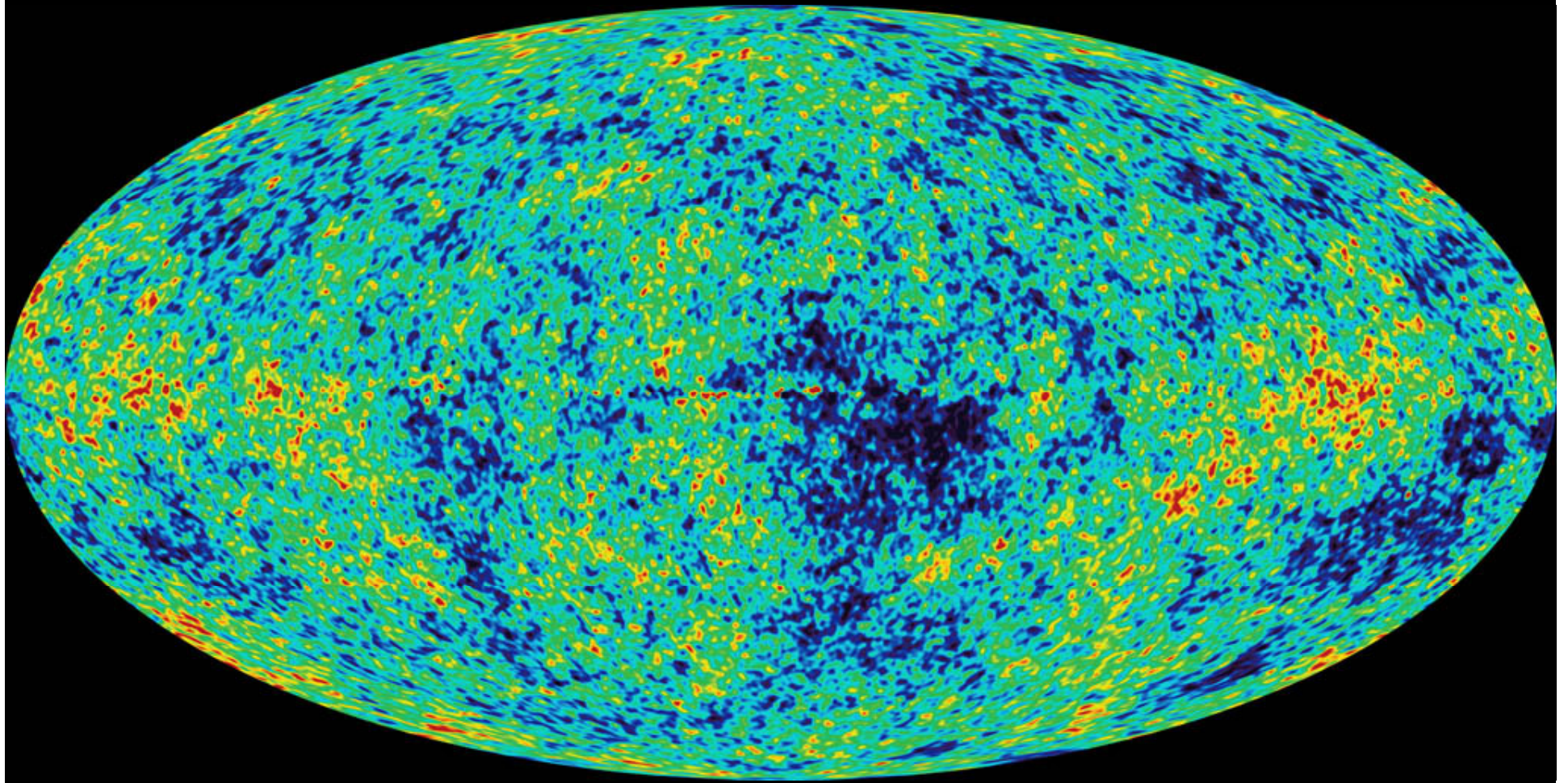


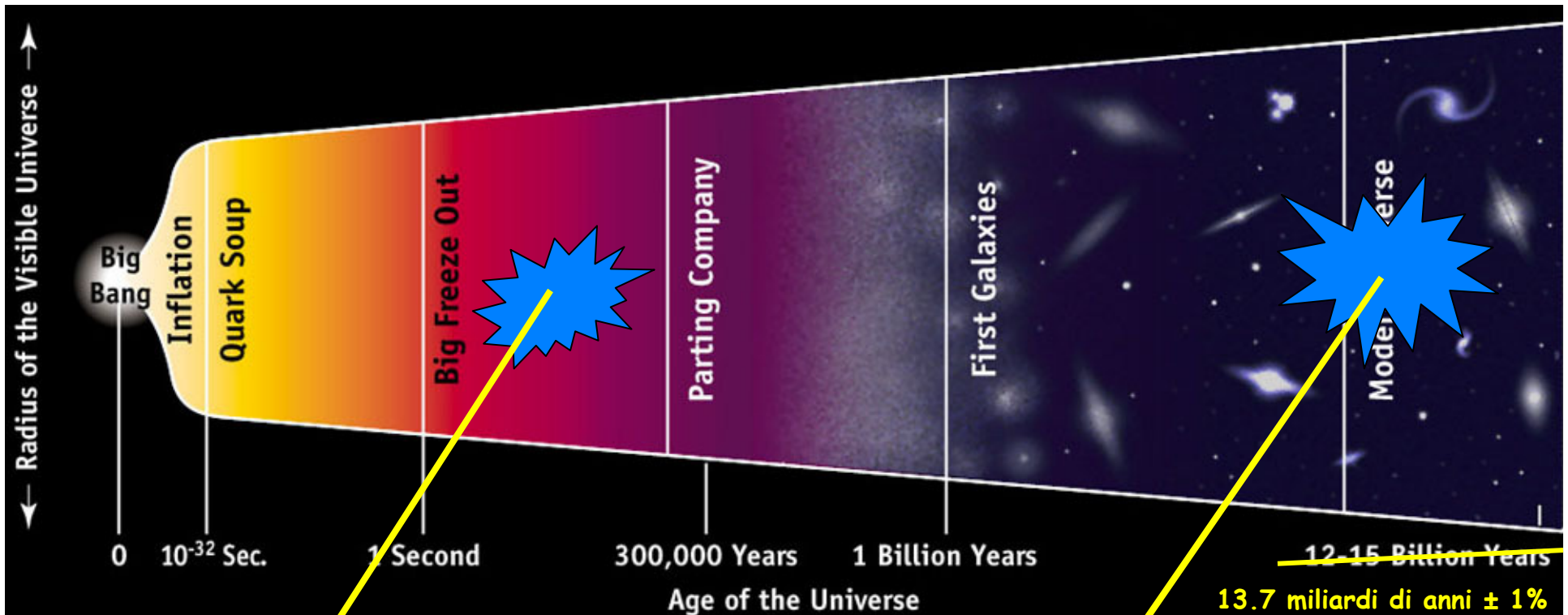
Aperto



Chiuso

L' Universo 380,000 anni
dopo il Big Bang





Nucleosintesi primordiale
 H ~ 75%
 He ~ 25%
 D, Li, B

Fusione nucleare nelle stelle
 He, C, O, N, ... Fe ...

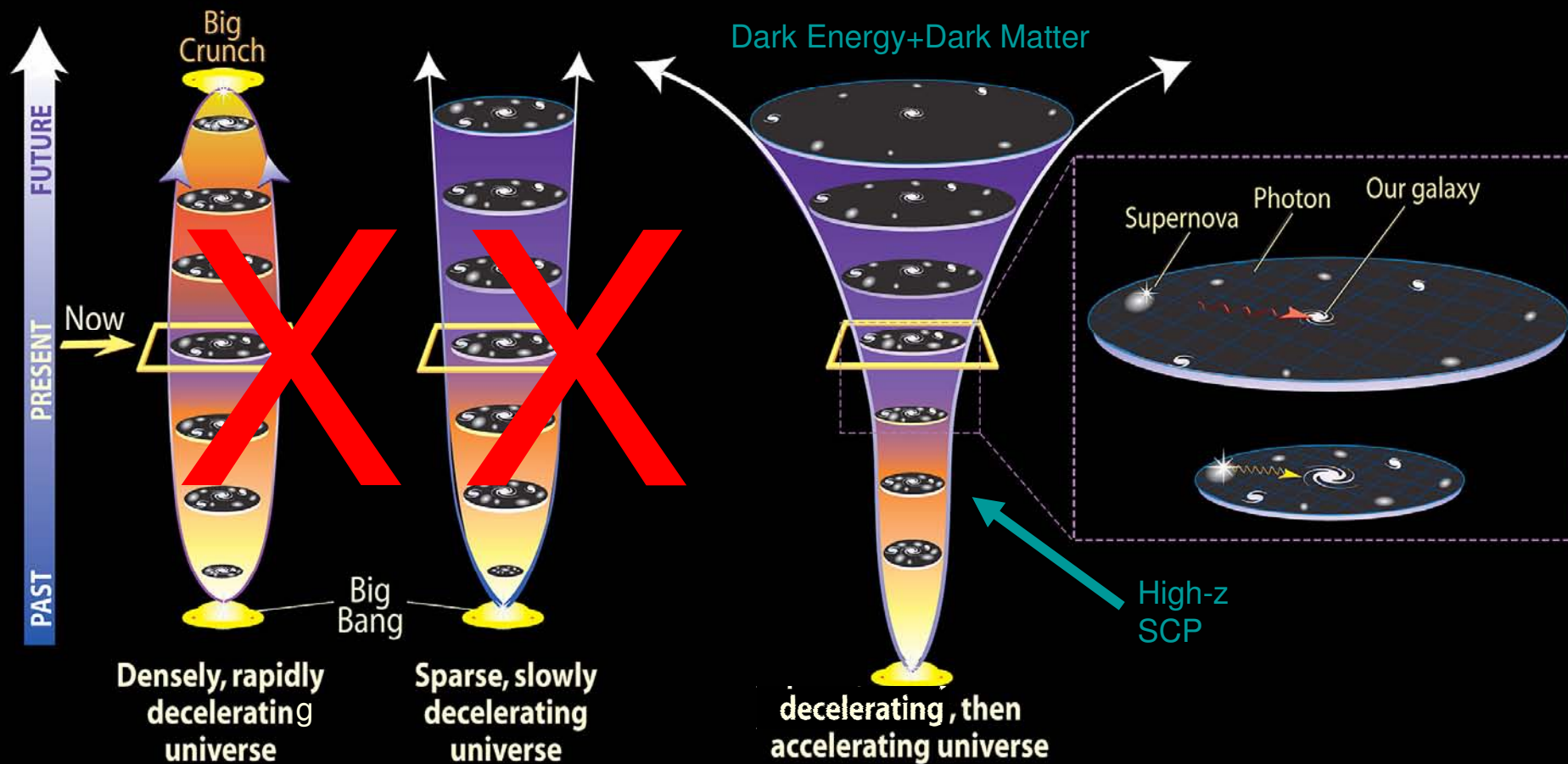
Diffusione nel mezzo interstellare
 Supernove, Venti stellari

Aggregazione di molecole
 Nubi molecolari, dischi protoplanetari

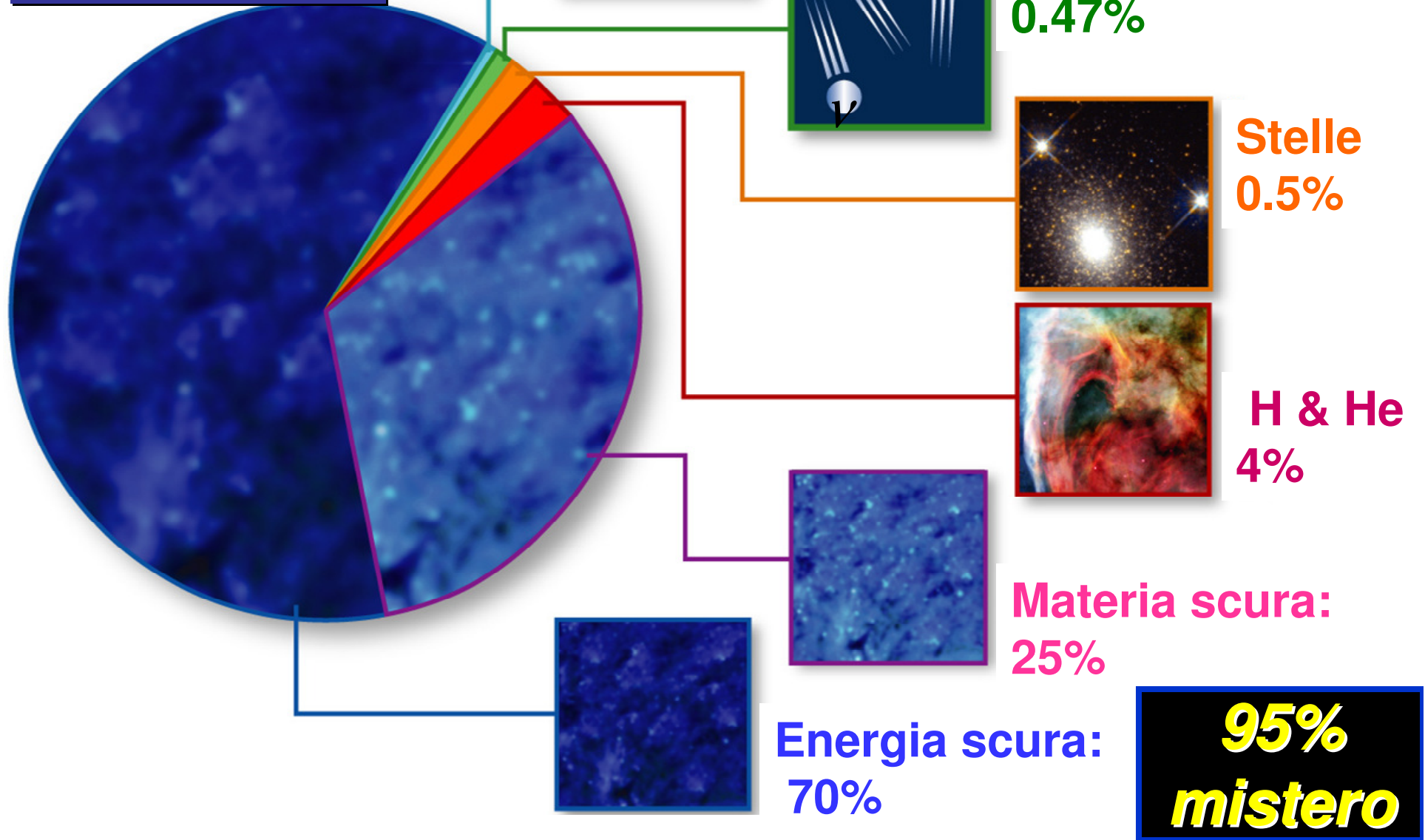
12-15 Billion Years
 13.7 miliardi di anni ± 1%

La storia dell'Espansione rivela: un **UNIVERSO** in espansione

Models of the Expanding Universe



Pizza cosmica



- “Grandissima mi par l’innezzia di coloro che vorrebbero che Iddio avesse fatto l’Universo più proporzionato alla piccola capacità del loro discorso...”

Galileo Galilei
Opere VII, 397



L'Astronomia



Una necessità primaria

*Ulterius progredi temporis angustia inhihet;
plura de his brevi candidus Lector expectet*

*La ristrettezza del tempo mi impedisce di
procedere oltre; il benigno Lettore aspetti tra
breve una piu' ampia trattazione di questo
argomento*

Galileo Galilei
Sidereus Nuncius